

Determination of the True Ionospheric Currents and Conductances from Combined Ground- and Space-Based Observations

Pulkkinen, A.^{1,2}, H. Korth³, E.
Cousins⁴ and S. Shepherd⁴

¹The Catholic University of America

²NASA GSFC, Space Weather Laboratory and CCMC

³JHU APL

⁴Thayer School of Engineering, Dartmouth College

Background

- Ionospheric electrodynamics one major controlling factor of the geospace dynamics.
- Ground-based or space-based magnetic field observations alone cannot provide the true full (field-aligned and horizontal) ionospheric electric currents.
- Combined usage of ground-based and space-based magnetic field observations required for direct derivation of the the true ionospheric currents.
- When combined further with ionospheric electric field observations, the true currents facilitate direct characterization of the full ionospheric electrodynamics.

Background

- Our goal is to use combined unprecedented global data sets to derive, for the first time, directly the full electrodynamic properties of the global ionosphere.
- Work supported by the NSF grant # AGS-1003513.

Methodology

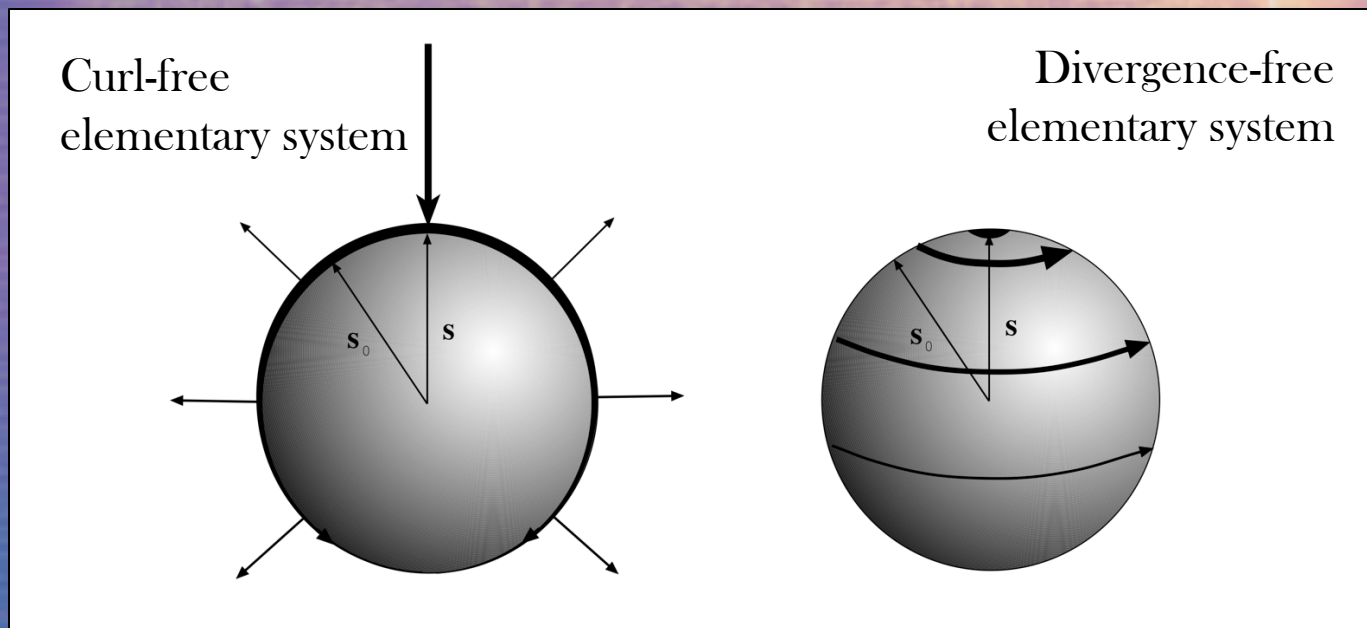
- We apply Spherical Elementary Currents Systems (SECS) to represent the divergence-free and curl-free parts of the system:

$$\bar{j} = \bar{j}_{df} + \bar{j}_{cf} = \int_V \bar{G}_{df} \times \bar{u} dV + \int_V \bar{G}_{cf} v dV \quad (1)$$

$$\begin{aligned} \nabla \cdot \bar{j}_{df} &= 0 & \nabla \times \bar{j}_{df} &= \bar{u} \\ \nabla \times \bar{j}_{cf} &= 0 & \nabla \cdot \bar{j}_{cf} &= v \end{aligned} \quad (2)$$

Methodology

- In SECS $\bar{G}_{df}, \bar{G}_{cf}$ in Eq. (1) are constructed using discrete elementary systems:



Methodology

- The amplitudes J_0^{df} , J_0^{cf} of the discrete elementary systems are solved using information about ground magnetic field fluctuations and field-aligned current densities, respectively. For details, see *Amm* (JGR, 2001) and *Pulkkinen et al.* (Earth, Planets, Space, 2003).

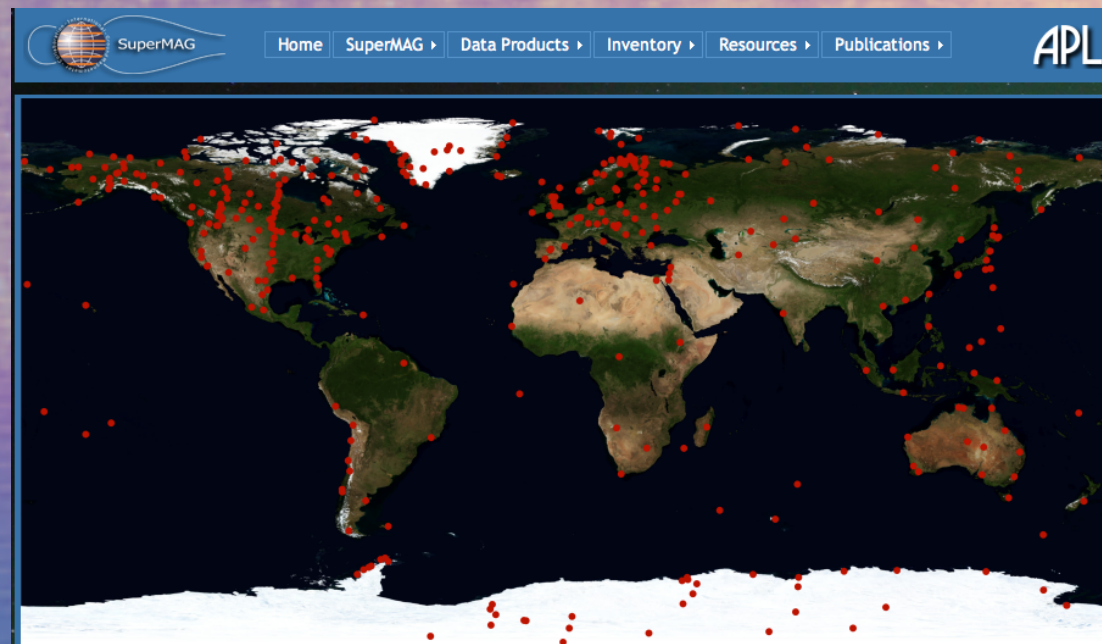
Methodology

- Finally, conductances can be solved from:

$$\sum_H = \frac{(\bar{j} \times \bar{E})_r}{|\bar{E}|^2} \leftarrow \text{Fields need to "line up" well} \quad (3)$$
$$\sum_P = \frac{\bar{j} \cdot \bar{E}}{|\bar{E}|^2} \leftarrow \text{Need to be careful with small electric field amplitudes}$$

Global observations I

- The ground magnetic field observations provided by SuperMag:



Global observations II

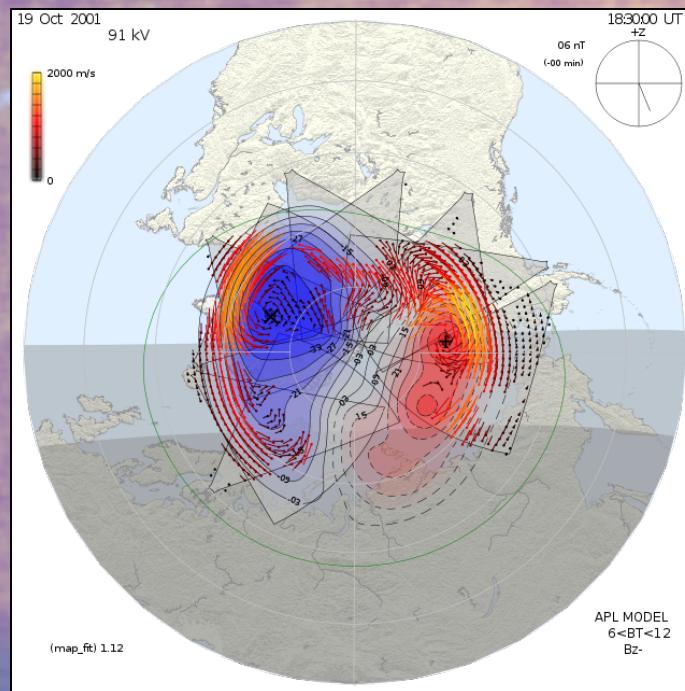
- The field-aligned current densities provided by the Iridium constellation:



>70 spacecraft
in six orbit
planes

Global observations III

- Ionospheric electric field provided by SuperDARN:



Progress

- We have identified a set of steady state periods for which initially lower temporal resolution Iridium data can be applied.
- Software interfaces between SuperMAG, SuperDARN and Iridium data and SECS software generated.
- SECS software package extended to include treatment of divergence-free systems.

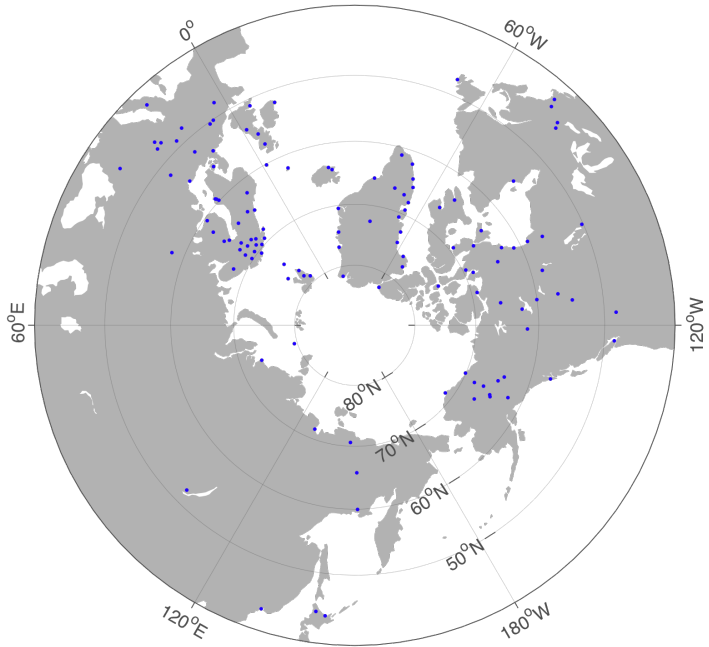
Progress

- Coordinate transformations and mappings between 780 and 110 km required:
 - IDL package for Altitude Adjusted Corrected Geomagnetic (AACGM) coordinates converted to Matlab.
 - New Matlab AACGM package freely available (email antti.a.pulkkinen@nasa.gov).
- DataMap reader for MatLab generated.

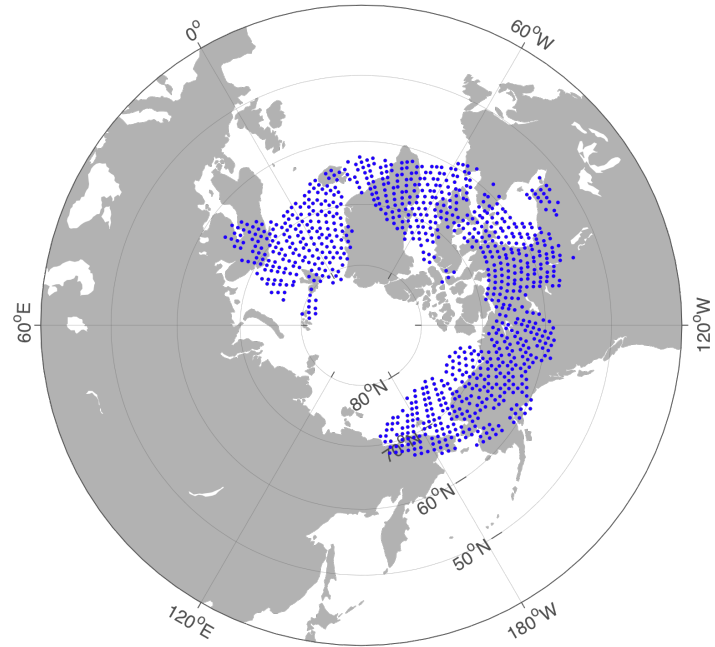
Initial results: 2001-10-09 14:00-15:00Z

In the following 1-min SuperMAG and 2-min SuperDARN data averaged over the given 60-min periods

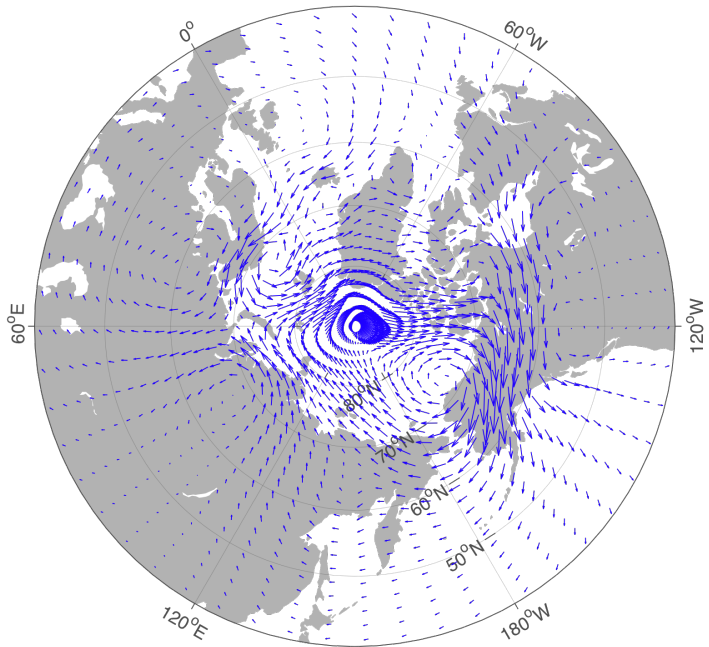
Supermag stations.



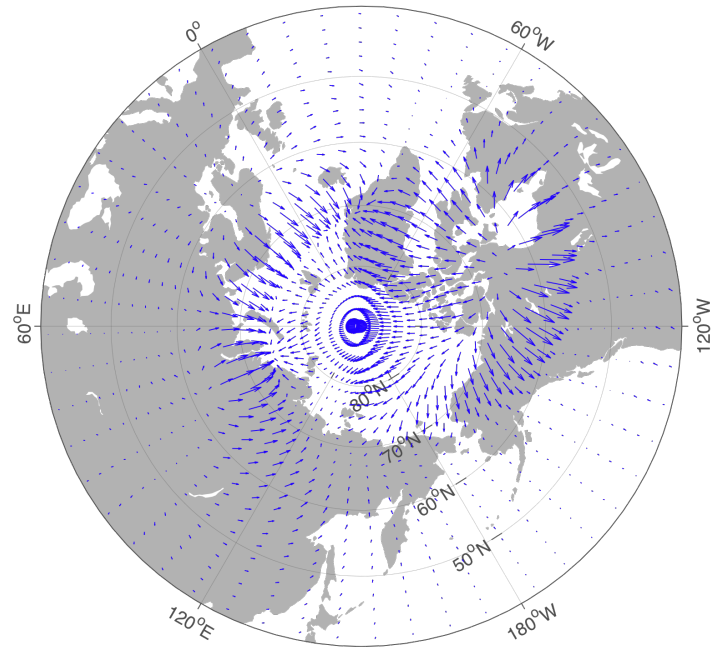
SuperDARN locations with good coverage.



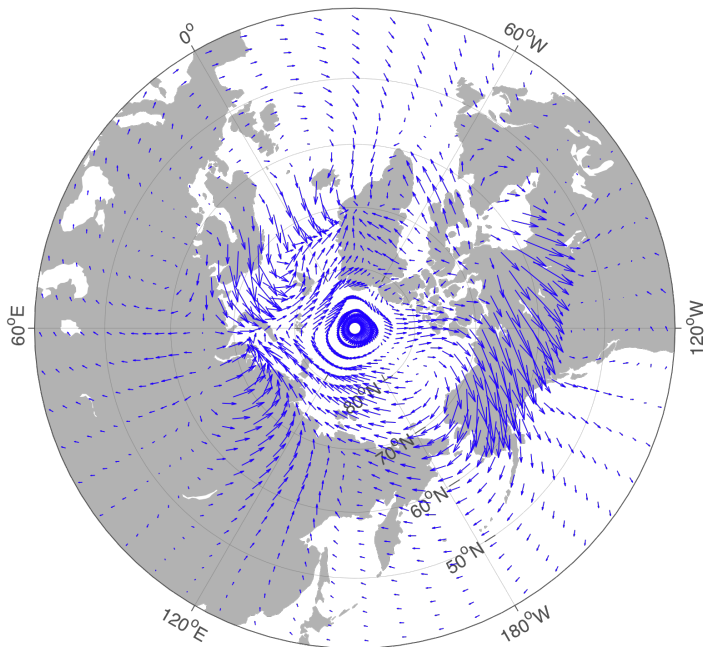
Supermag divergence-free horizontal currents. Max. $|j|$: 0.82 A/m.



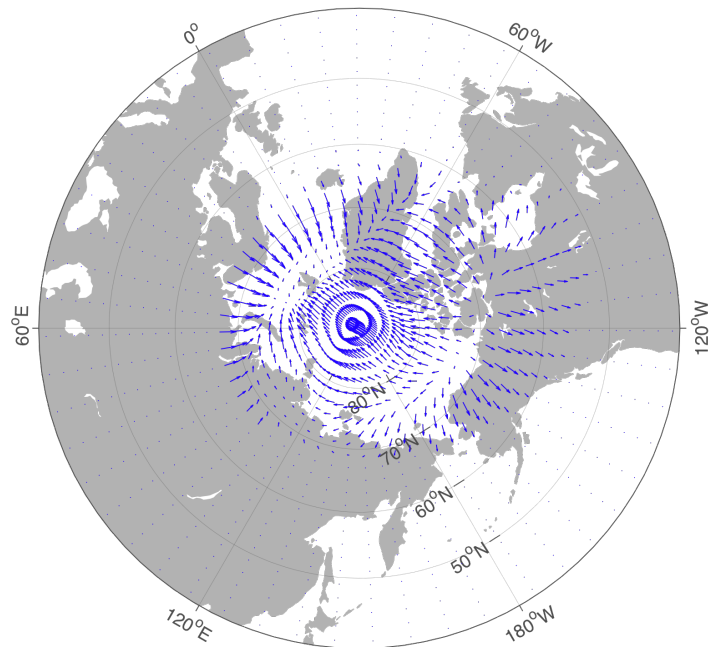
Iridium curl-free horizontal currents. Max. $|j|$: 0.57 A/m.



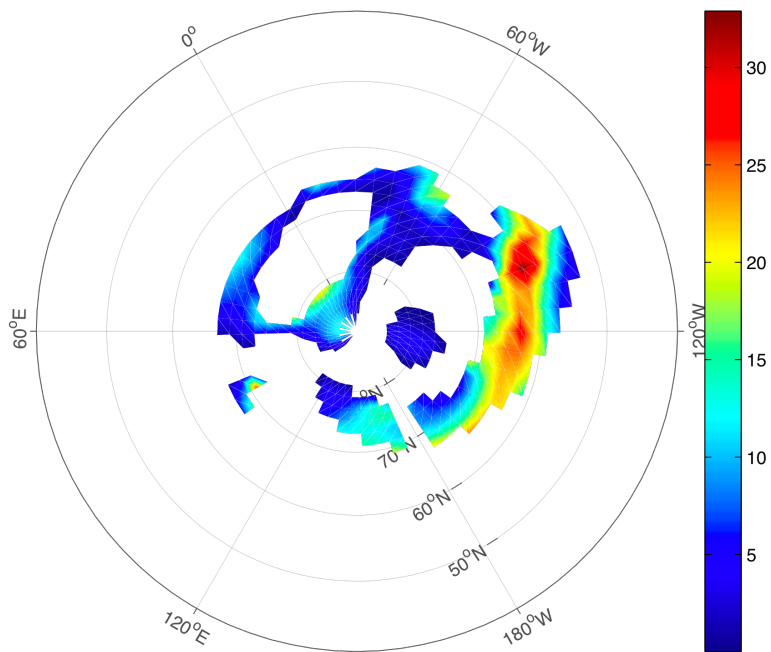
Total horizontal currents. Max. I_{Hj} : 1.04 A/m.



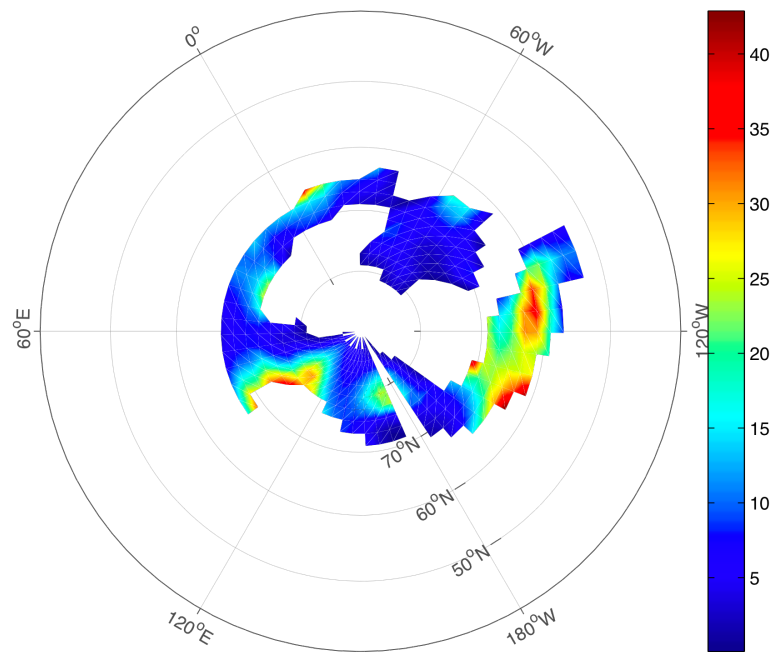
Electric field. Max. $|E|$: 0.05 V/m.



Hall conductance. Max Σ_{Hj} : 32.90 Siemens.



Pedersen conductance. Max Σ_{Pj} : 42.87 Siemens.



Next steps

- Analysis of all 20 quasi-steady state one-hour periods.
- Carry out detailed sciences analyses.
- Use the derived global ionospheric maps in the global geospace model validation.
- Apply the methodology and the generated software with the new AMPERE data.

Acknowledgements

- The work is supported by the NSF grant # AGS-1003513. We also wish to acknowledge SuperMAG PI J. Gjerloev for help with the magnetometer data.
- For the ground magnetometer data we gratefully acknowledge: Intermagnet; USGS, Jeffrey J. Love; Danish Meteorological Institute; CARISMA, PI Ian Mann; CANMOS; The S-RAMP Database, PI K. Yumoto and Dr. K. Shiokawa; The SPIDR database; AARI, PI Oleg Troshichev; The MACCS program, PI M. Engebretson, Geomagnetism Unit of the Geological Survey of Canada; GIMA; MEASURE, UCLA IGPP and Florida Institute of Technology; SAMBA, PI Eftyhia Zesta; 210 Chain, PI K. Yumoto; SAMNET, PI Farideh Honary; The institutes who maintain the IMAGE magnetometer array, PI Eija Tanskanen; PENGUIN; AUTUMN, PI Martin Connors; Greenland magnetometers operated by DTU Space; South Pole and McMurdo Magnetometer, PI's Louis J. Lanzarotti and Alan T. Weatherwax; ICESTAR; RAPIDMAG; PENGUIn; British Antarctic Survey; McMac, PI Dr. Peter Chi; BGS, PI Dr. Susan Macmillan; Pushkov Institute of Terrestrial Magnetism, Ionosphere and Radio Wave Propagation (IZMIRAN).