

ULF wave observations in the inner magnetosphere

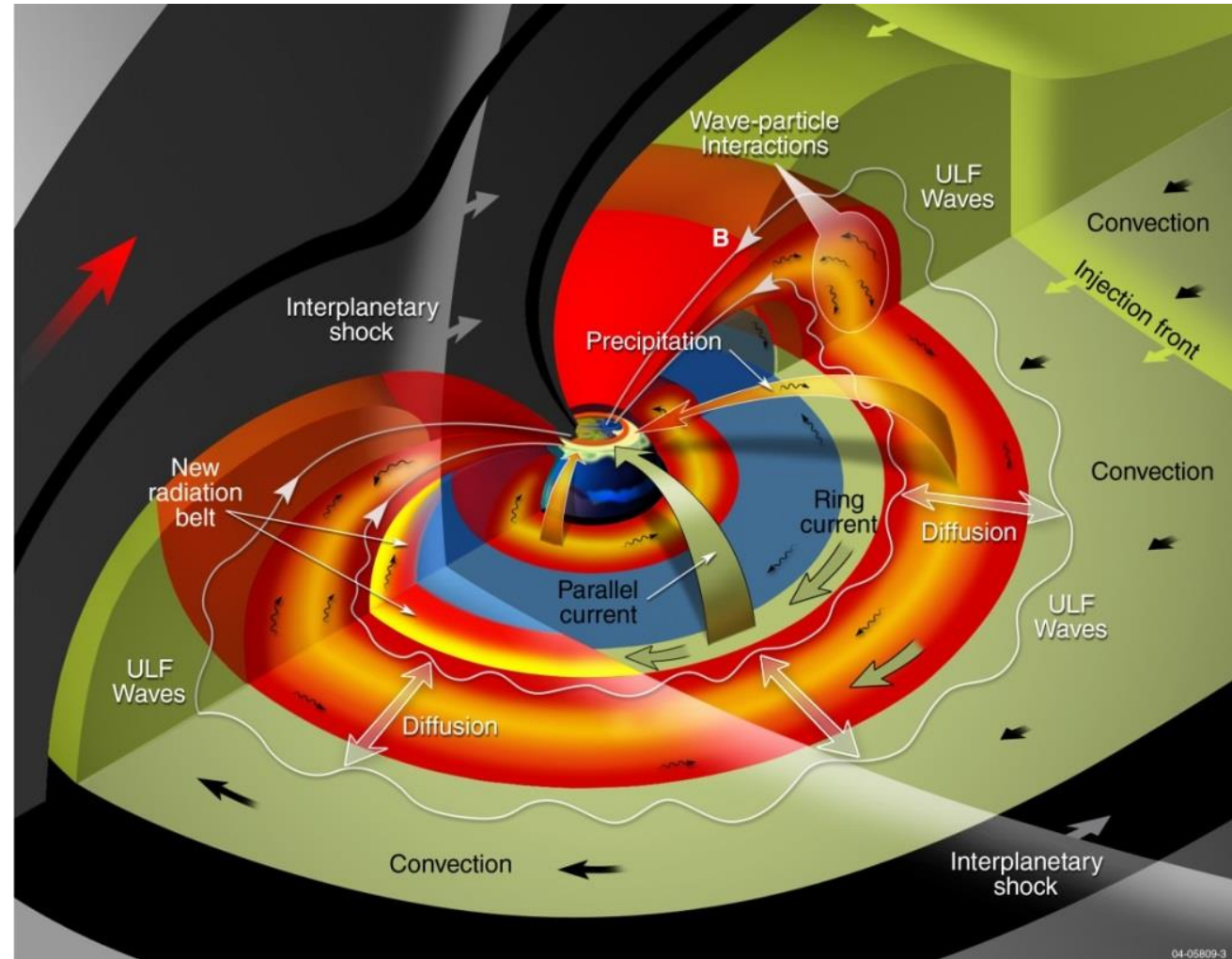
Michael Hartinger^{1,2}

¹Virginia Tech, ²National Institute of Aerospace

Acknowledgements: Ashar Ali, Lauren Blum, Christopher Chaston, Seth Claudepierre, Bob Clauer, Scot Elkington, Mark Engebretson, Alexa Halford, Allison Jaynes, Hyomin Kim, David Malaspina, Toshi Nishimura, Kazue Takahashi, Maria Usanova, Zhonghua Xu, Colin Waters, ...

Ultra Low Frequency Waves

- Ultra Low Frequency (ULF) waves: plasma waves in the Earth's magnetosphere with absolute frequency less than 5 Hz (Jacobs et al., 1964)
- Largest wavelengths and lowest frequencies in system
- Space weather impacts including radiation belt/ring current interactions



[Mauk et al., 2012]

“...the subject of hydromagnetic waves in the magnetosphere (or magnetospheres) is highly developed...” but

“...the subject is far from played out as a research field...”

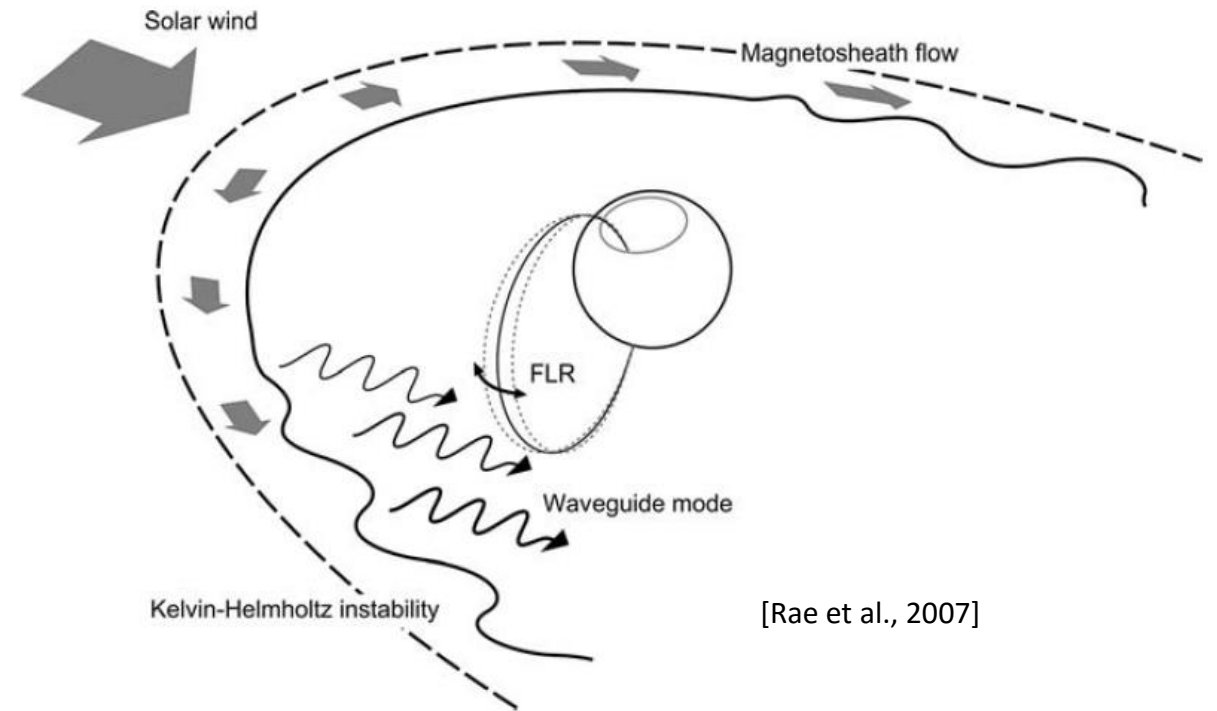
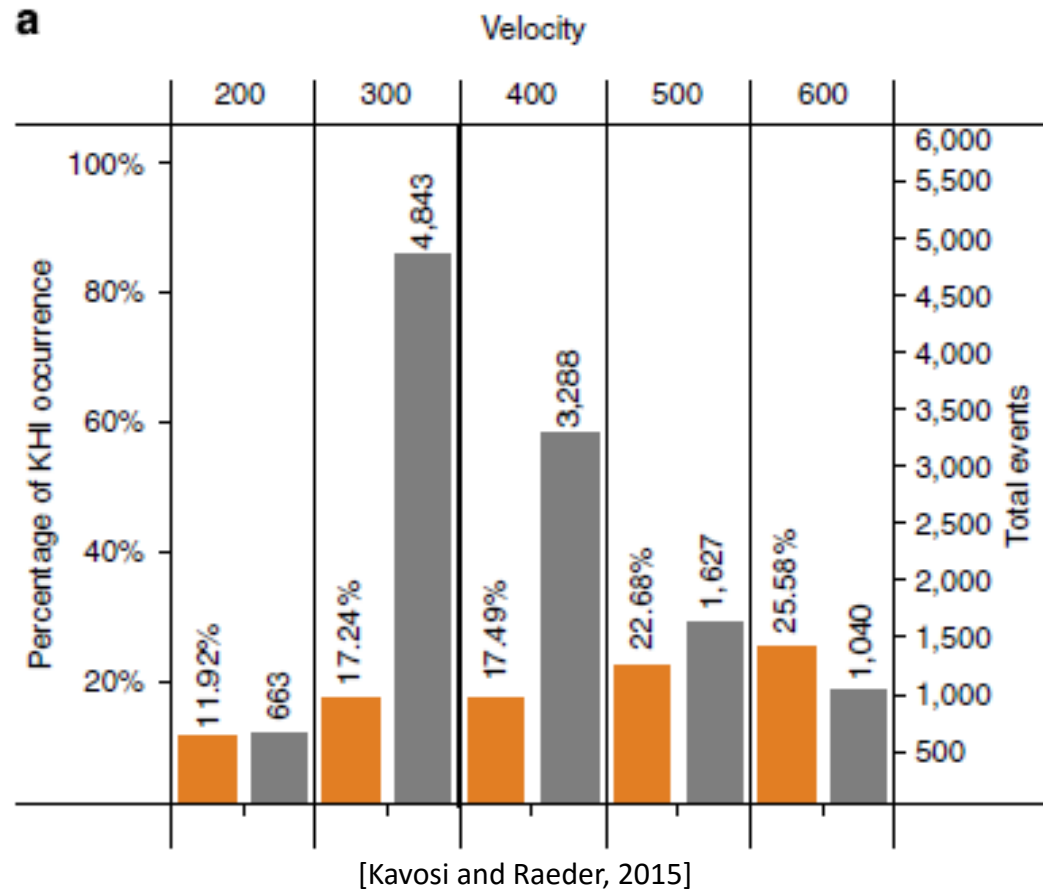
Southwood and Hughes, [1983]

Outline

- Focus on recent advances, challenges, and open questions
- Focus on research related to over-arching questions:
 - What excites ULF waves?
 - How do ULF waves couple to the plasmasphere? Ring current? Radiation belts?
 - What is the role of ULF waves in Magnetosphere-Ionosphere Coupling?
- Introducing the new GEM focus group, “ULF wave modeling, effects, and applications”

What excites inner magnetosphere ULF waves?

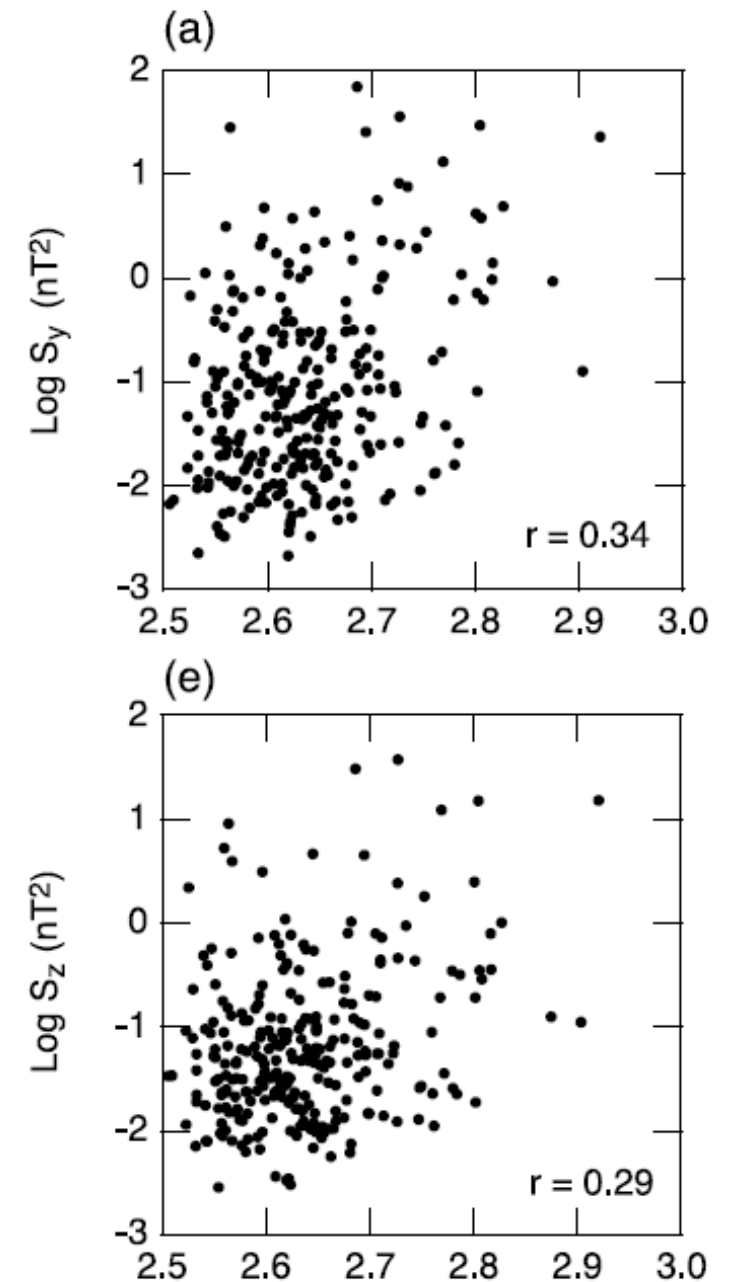
Example: Magnetopause surface waves



- Increasing availability of spacecraft observations near the magnetopause
- Statistics on magnetopause surface waves related to the Kelvin-Helmholtz instability - increased occurrence with increased flow shear
- These waves couple to inner magnetospheric ULF waves

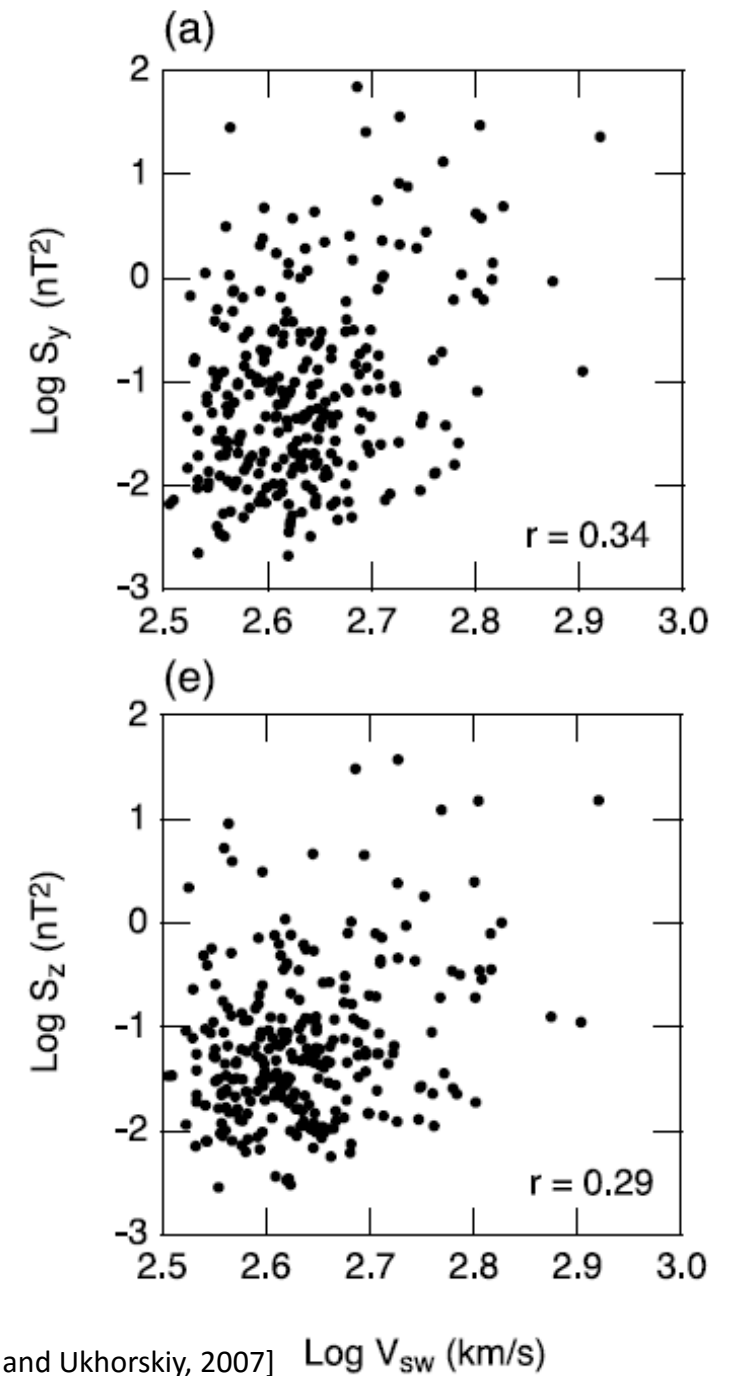
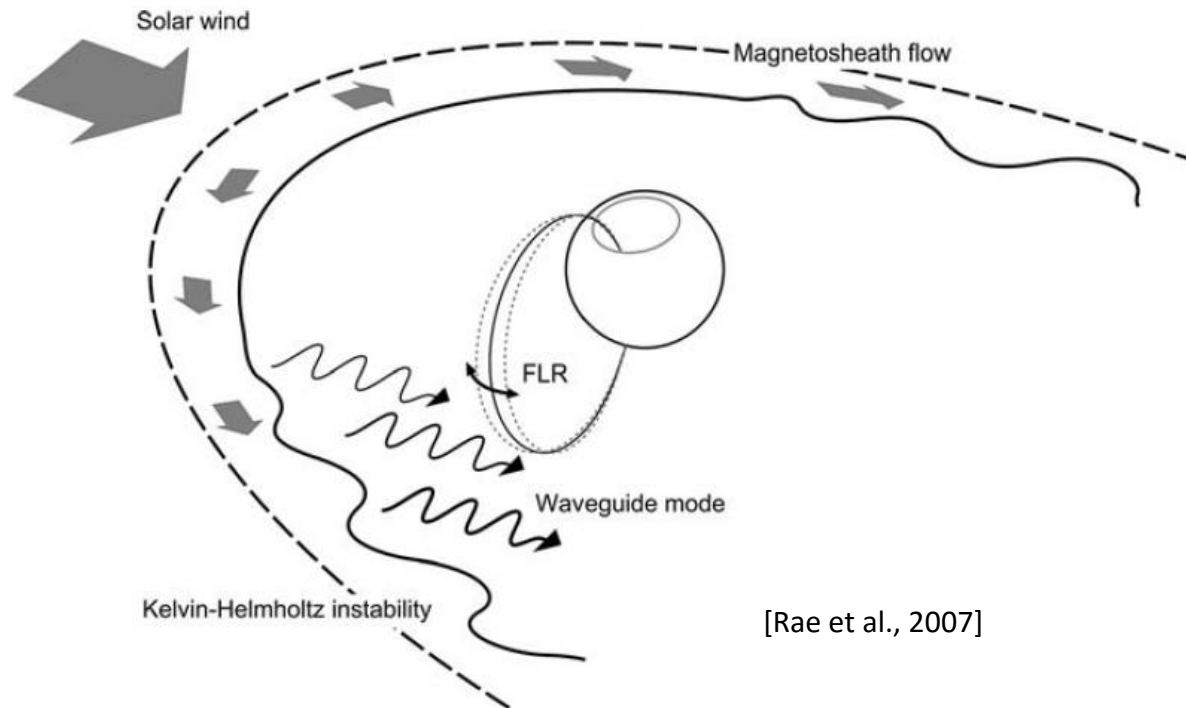
What excites inner magnetosphere ULF waves? Example: Magnetopause surface waves

- Inner magnetospheric Pc5 ($\sim 2\text{-}7$ mHz) ULF wave power and occurrence rate generally increase with solar wind flow speed in the flank magnetosphere
- Large amount of scatter in the trend

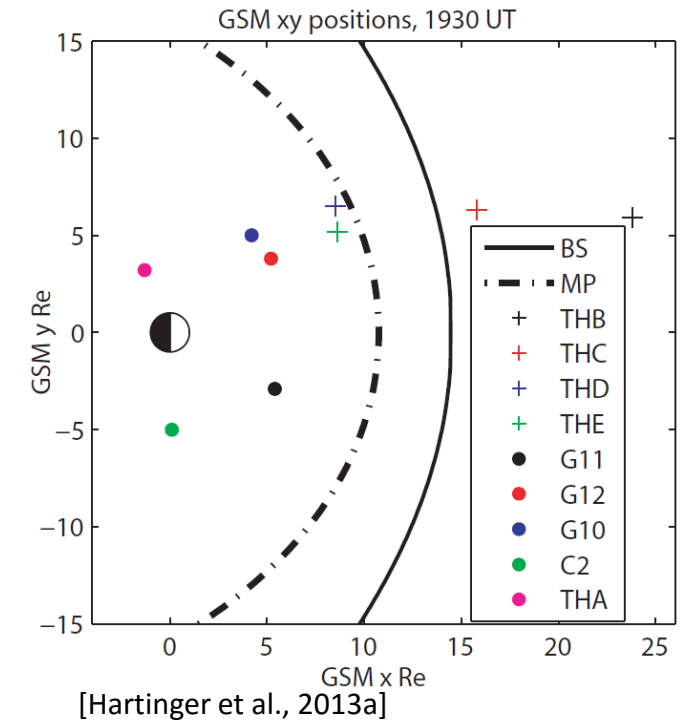
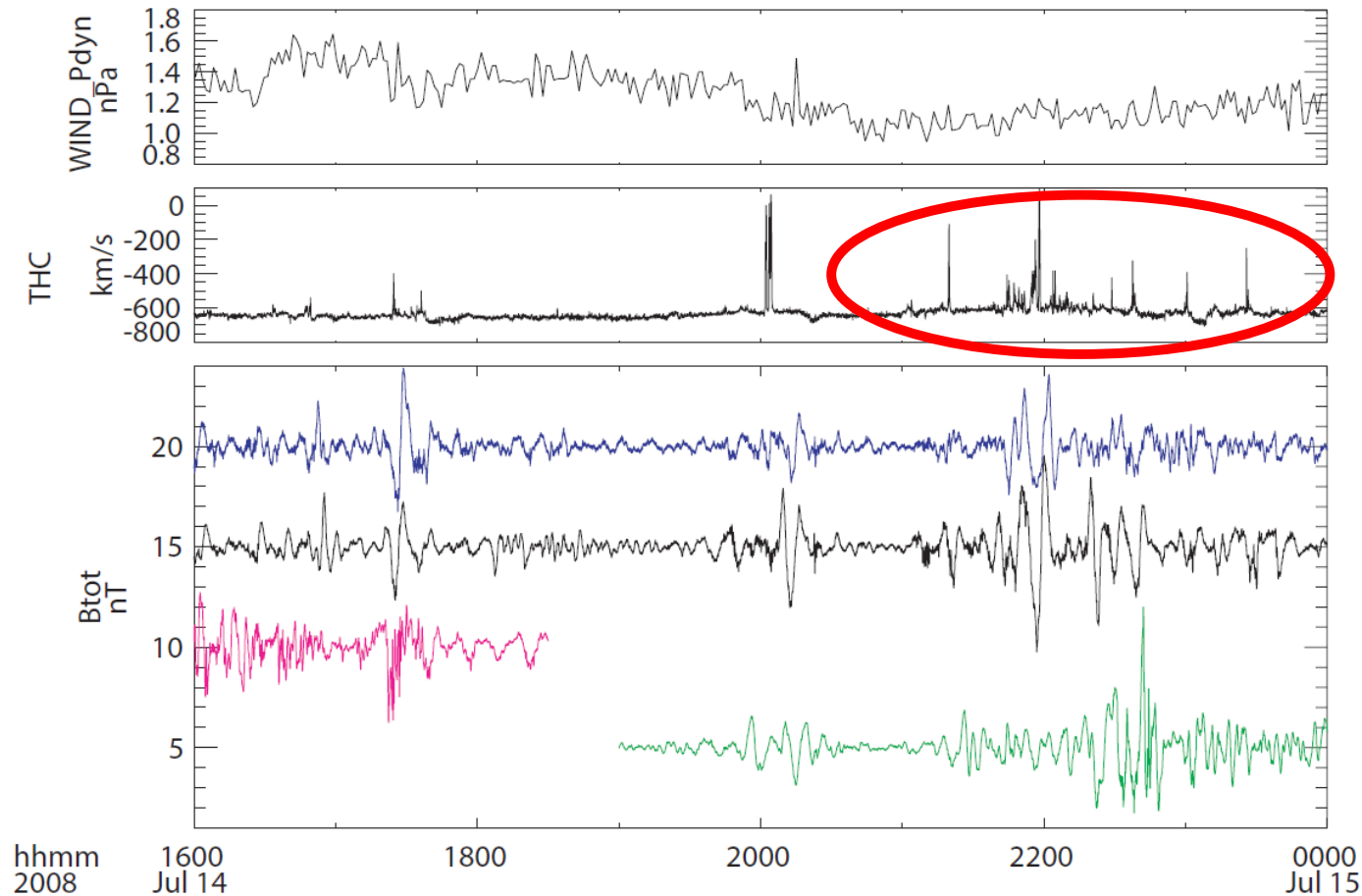


What excites inner magnetosphere ULF waves? Example: Magnetopause surface waves

- Inner magnetospheric Pc5 ($\sim 2\text{--}7$ mHz) ULF wave power and occurrence rate generally increase with solar wind flow speed in the flank magnetosphere
- Large amount of scatter in the trend
- Need more observations to explain scatter: role of seed fluctuations, magnetospheric Alfvén speed and other factors

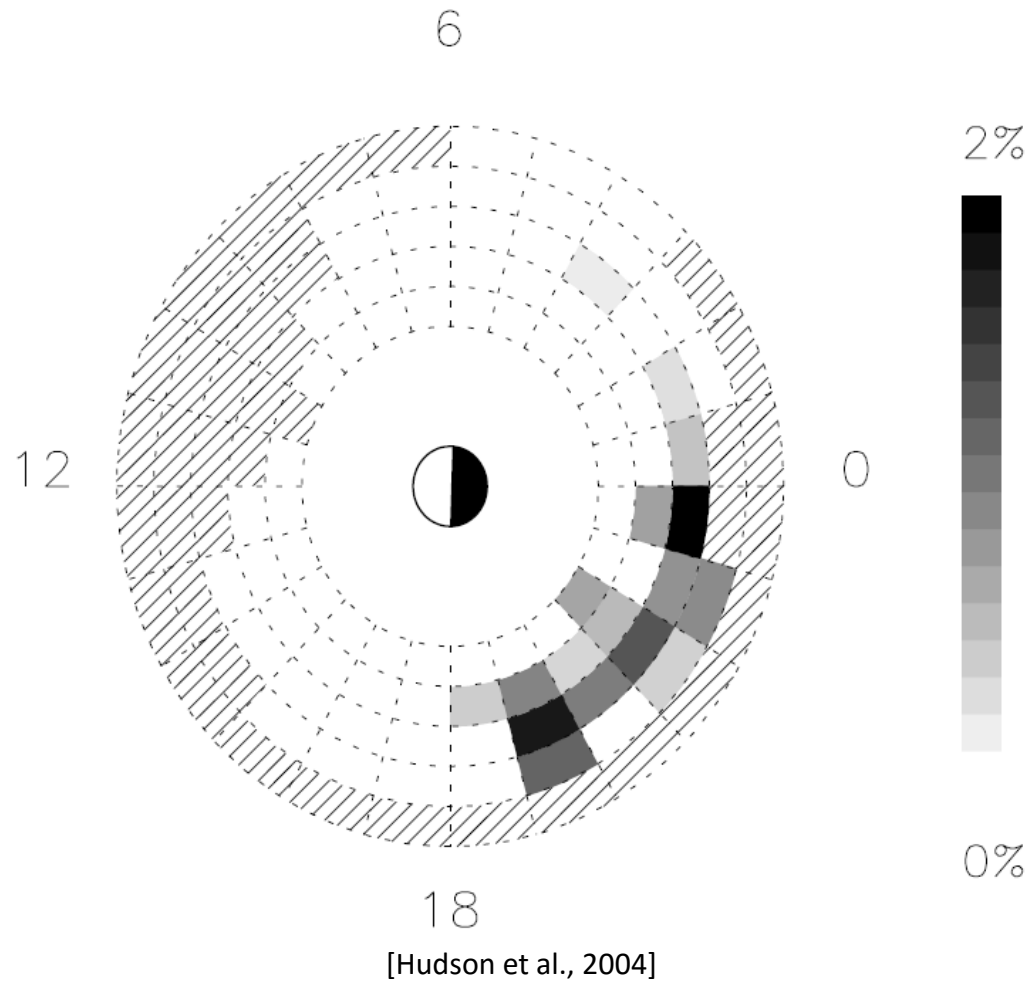


What excites inner magnetosphere ULF waves? Example: Magnetopause surface waves



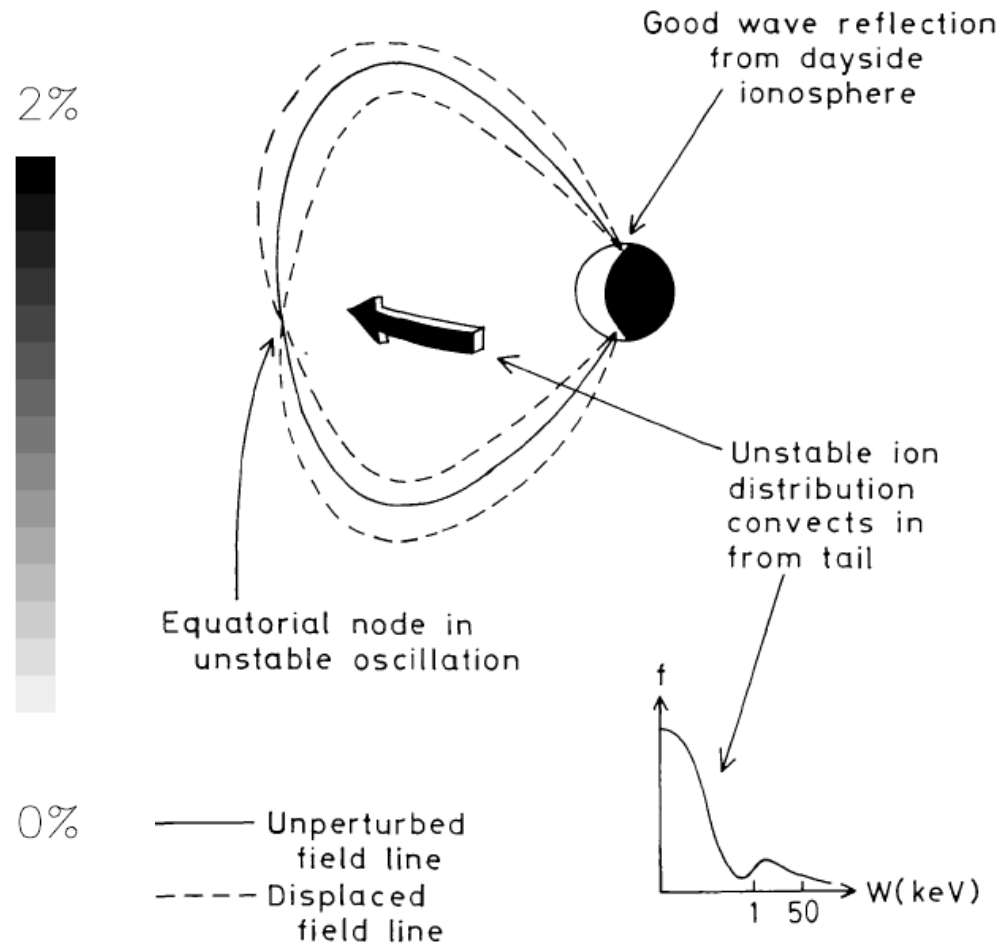
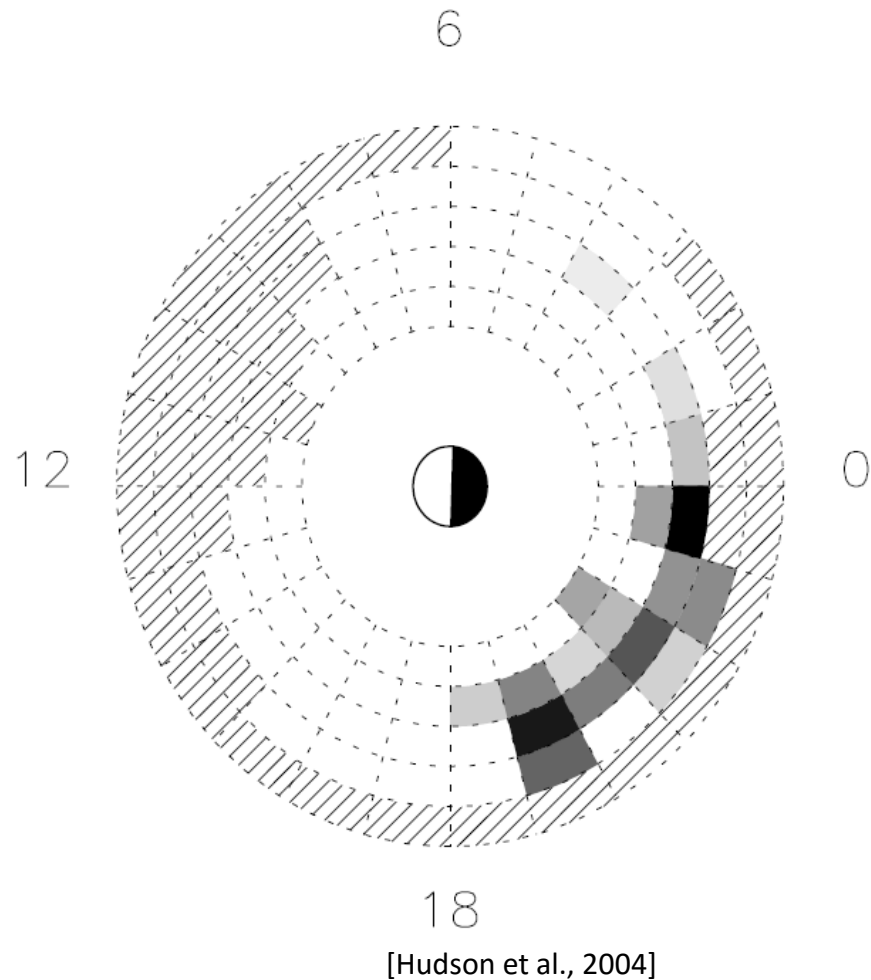
- The ion foreshock can provide sporadic and/or nearly continuous impulses to seed the growth of magnetopause surface waves during high speed solar wind

What excites inner magnetosphere ULF waves? Example: Drift and Drift-bounce resonance



- Inner magnetosphere second harmonic poloidal mode standing Alfvén waves occur in the pre-midnight sector

What excites inner magnetosphere ULF waves? Example: Drift and Drift-bounce resonance

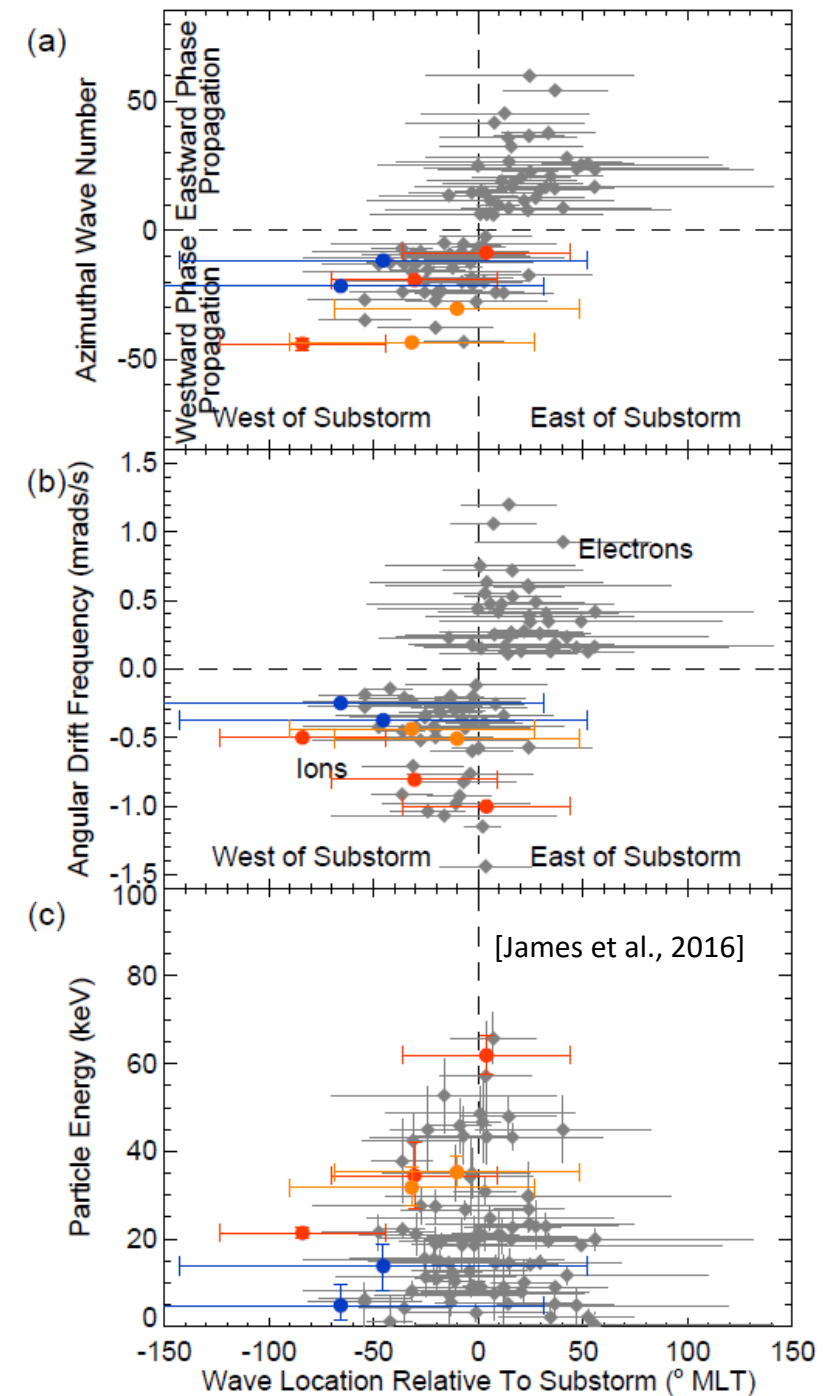
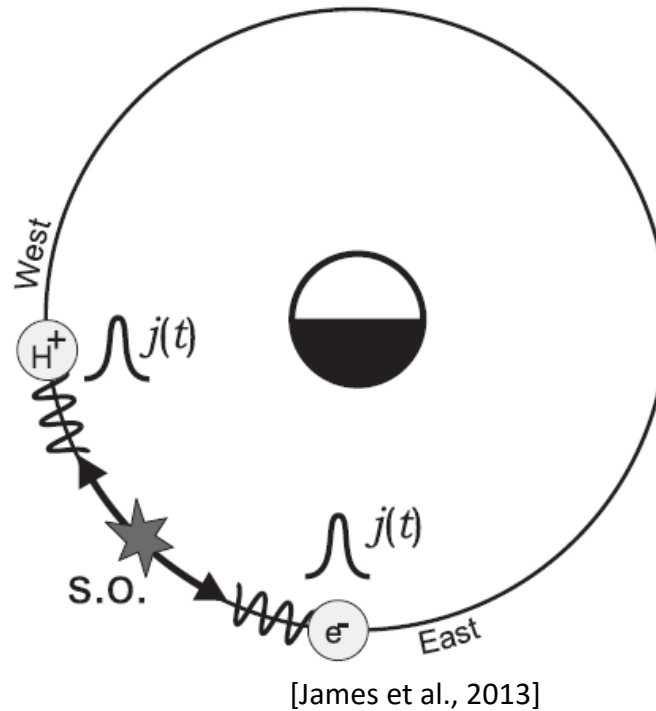


[Southwood and Hughes, 1983]

- Inner magnetosphere second harmonic poloidal mode standing Alfvén waves occur in the pre-midnight sector
- Generally associated with ion injections
- Wave generation mechanisms include drift-bounce resonance

What excites inner magnetosphere ULF waves? Example: Drift and Drift-bounce resonance

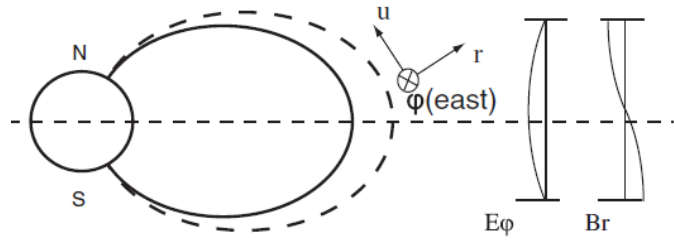
- Increasingly, sunward propagating ULF waves are observed in the post-midnight sector (e.g., Eriksson et al., 2008)
- Ground-based SuperDARN radar observations have provided excellent statistics on wave properties



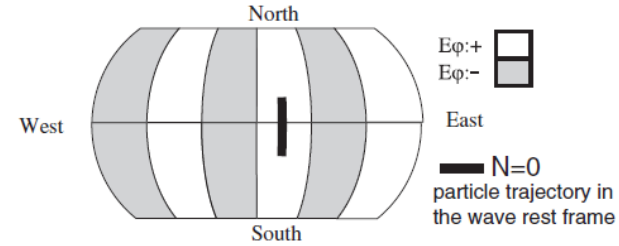
How do ULF waves couple to the ring current?

Example: Drift and Drift-bounce resonance

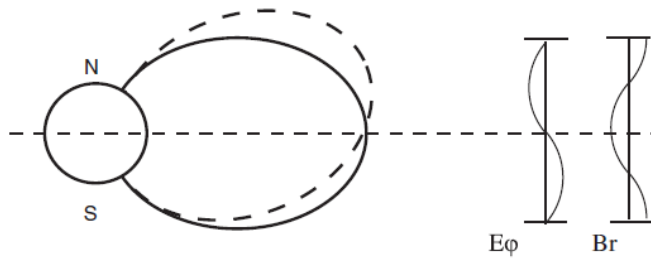
a) Fundamental poloidal mode



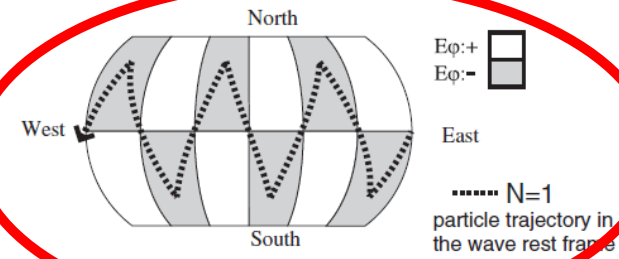
b) drift resonance with a fundamental mode



Second-harmonic poloidal mode



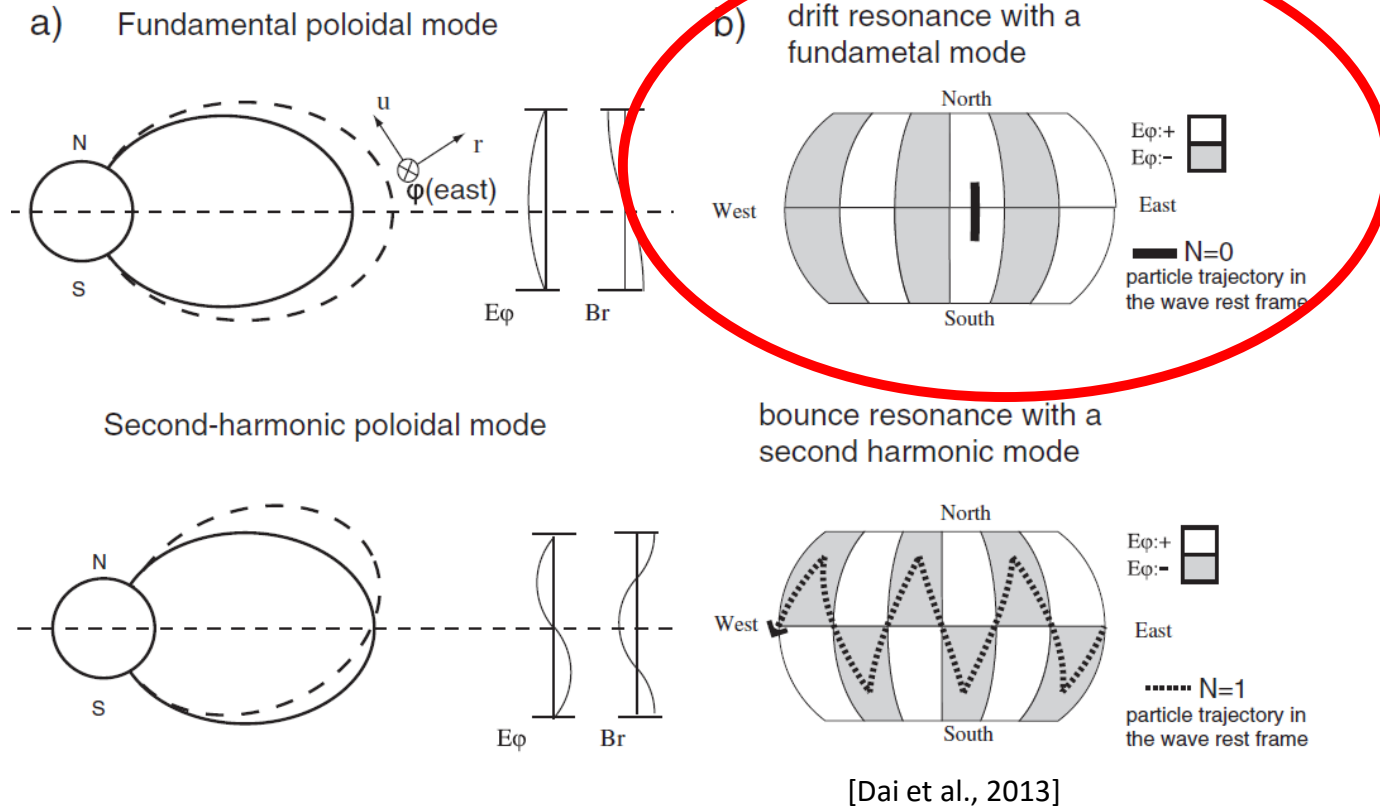
bounce resonance with a second harmonic mode



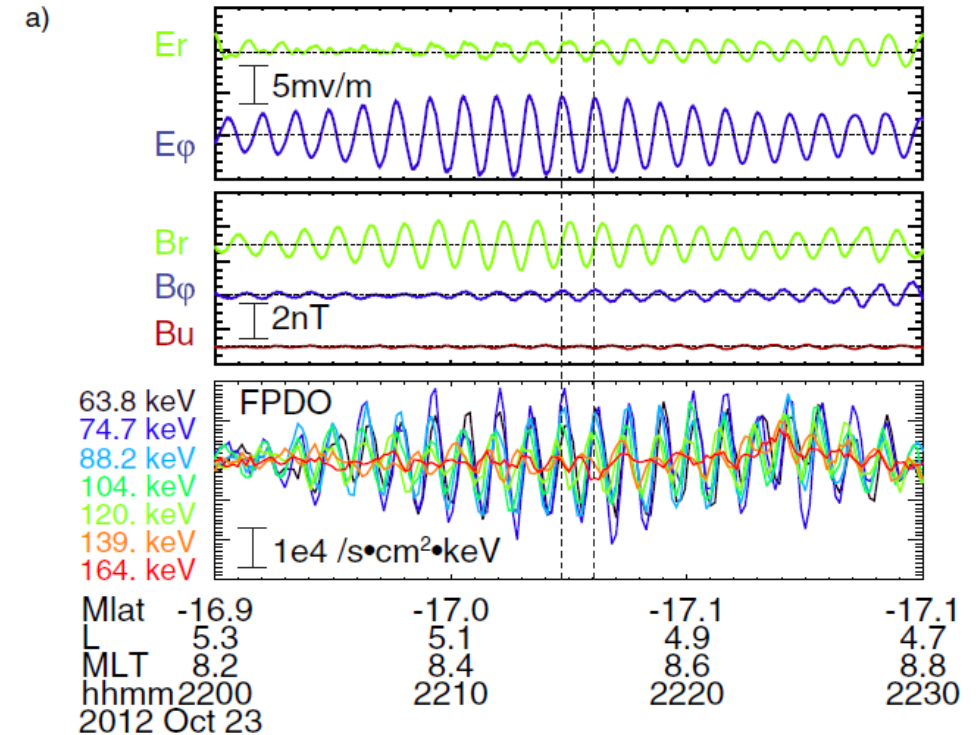
[Dai et al., 2013]

How do ULF waves couple to the ring current?

Example: Drift and Drift-bounce resonance



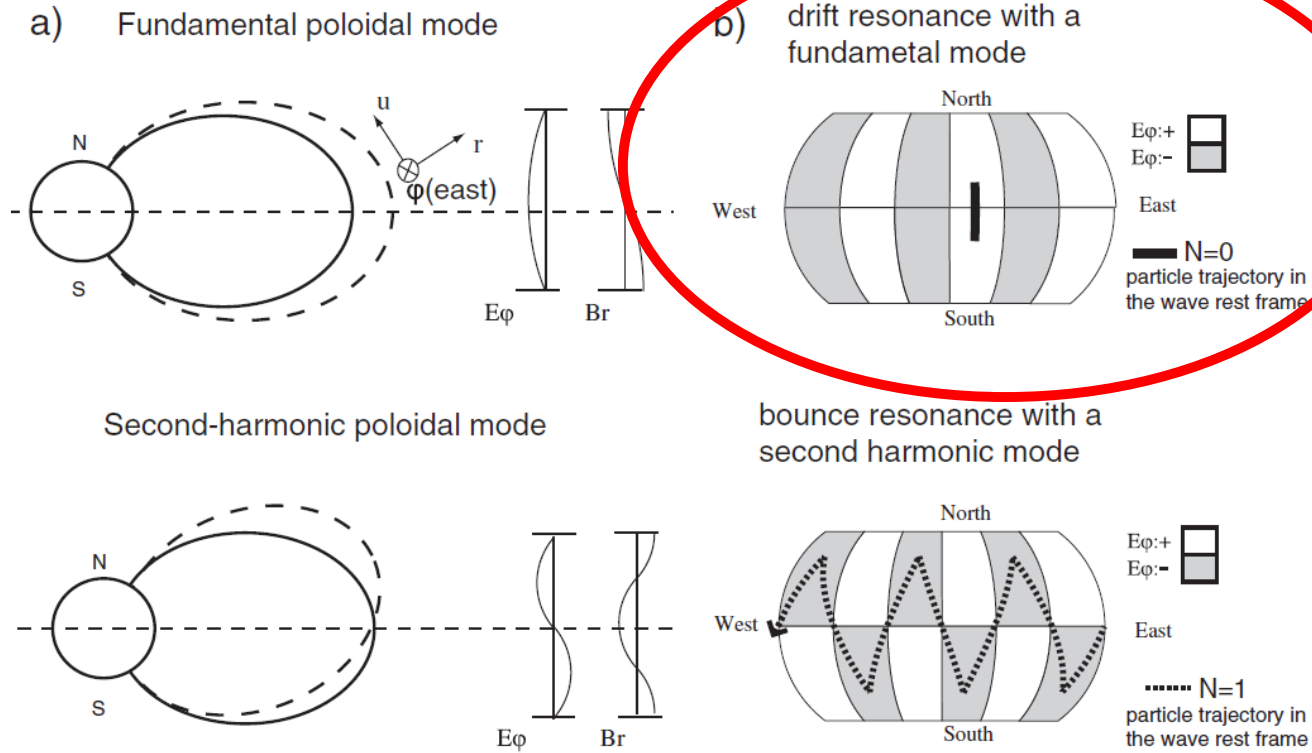
- In situ particle data from Van Allen Probes have confirmed the operation of drift resonance with ions → excitation of fundamental poloidal mode



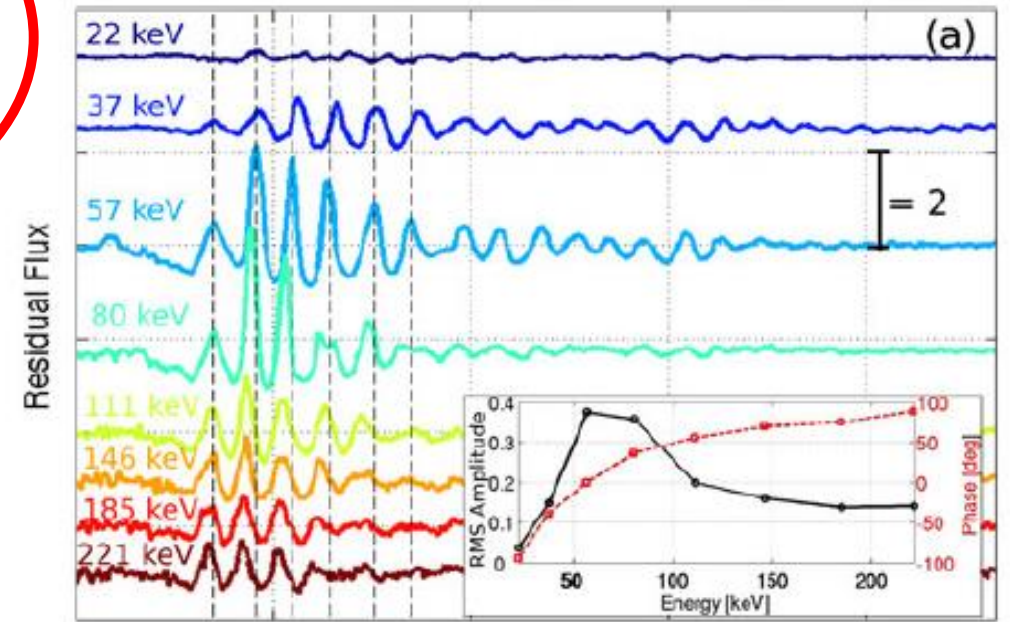
[Dai et al., 2013]

How do ULF waves couple to the ring current?

Example: Drift and Drift-bounce resonance



[Dai et al., 2013]



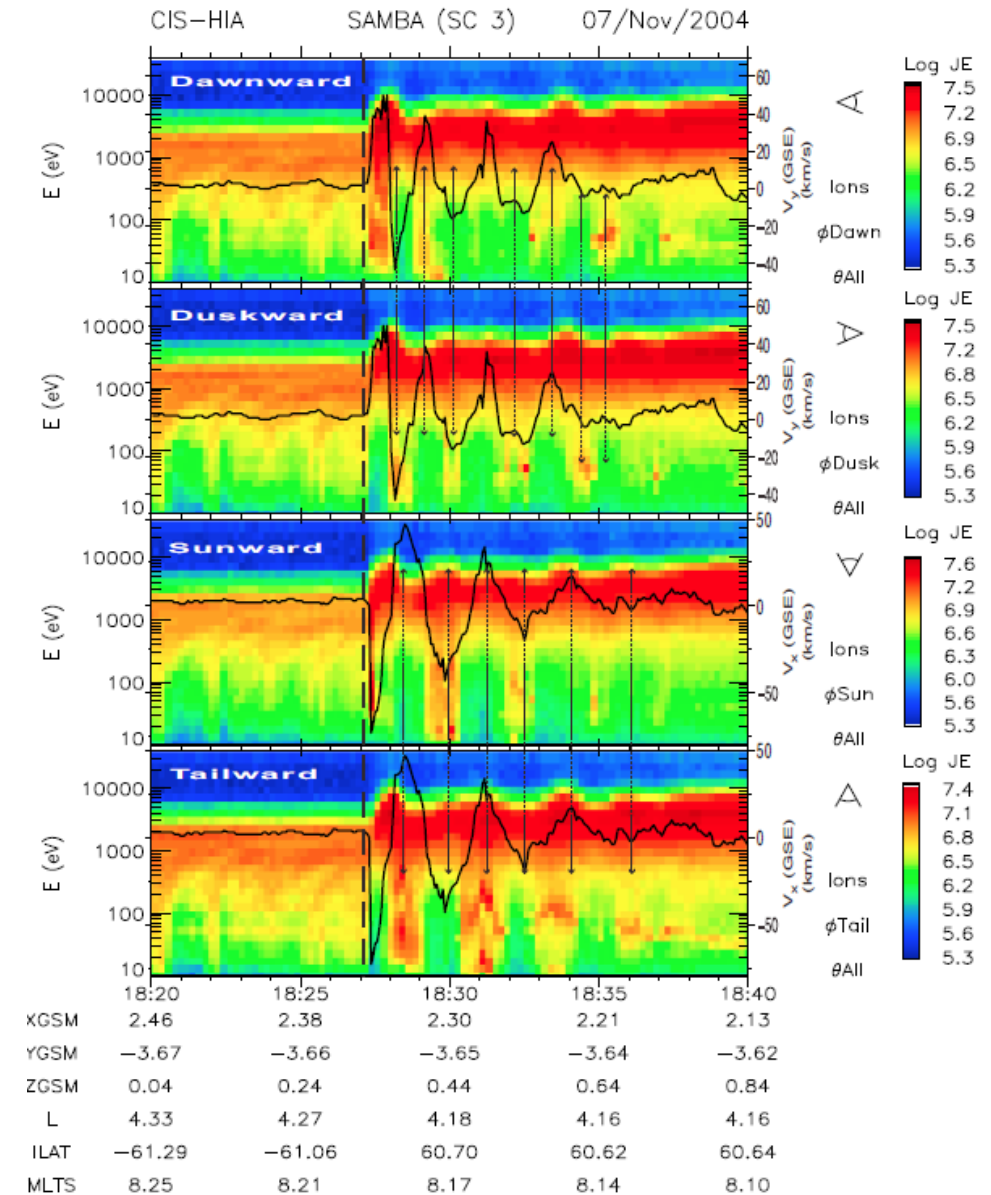
[Claudepierre et al., 2013]

- In situ particle data from Van Allen Probes have confirmed the operation of drift resonance with electrons \rightarrow excitation of fundamental poloidal mode
- Associated with solar wind pressure increase

How do ULF waves couple to the ring current?

Example: Drift and Drift-bounce resonance

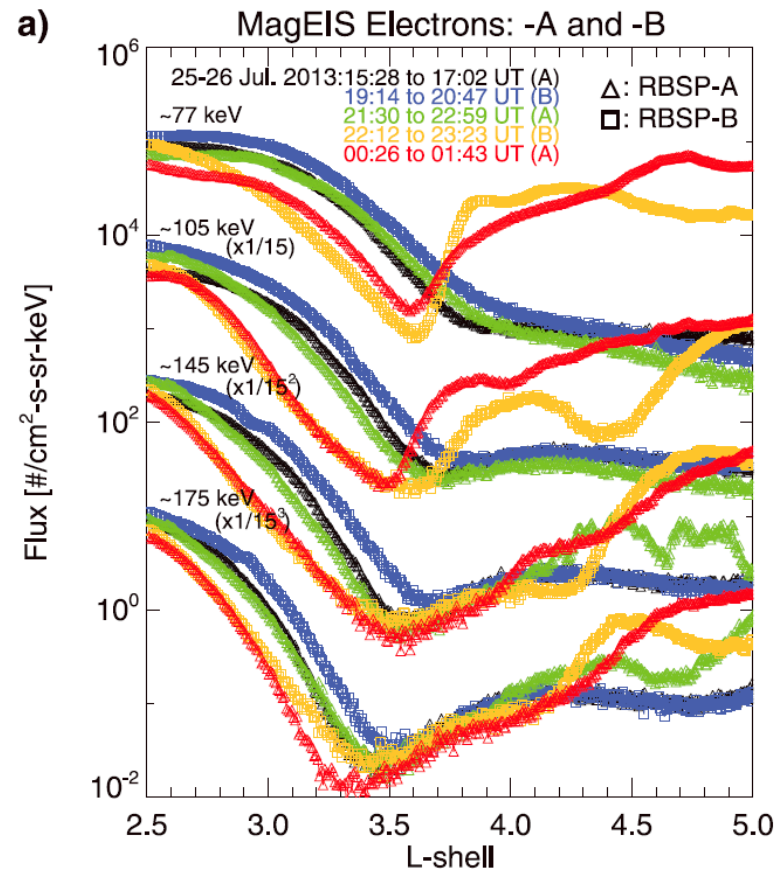
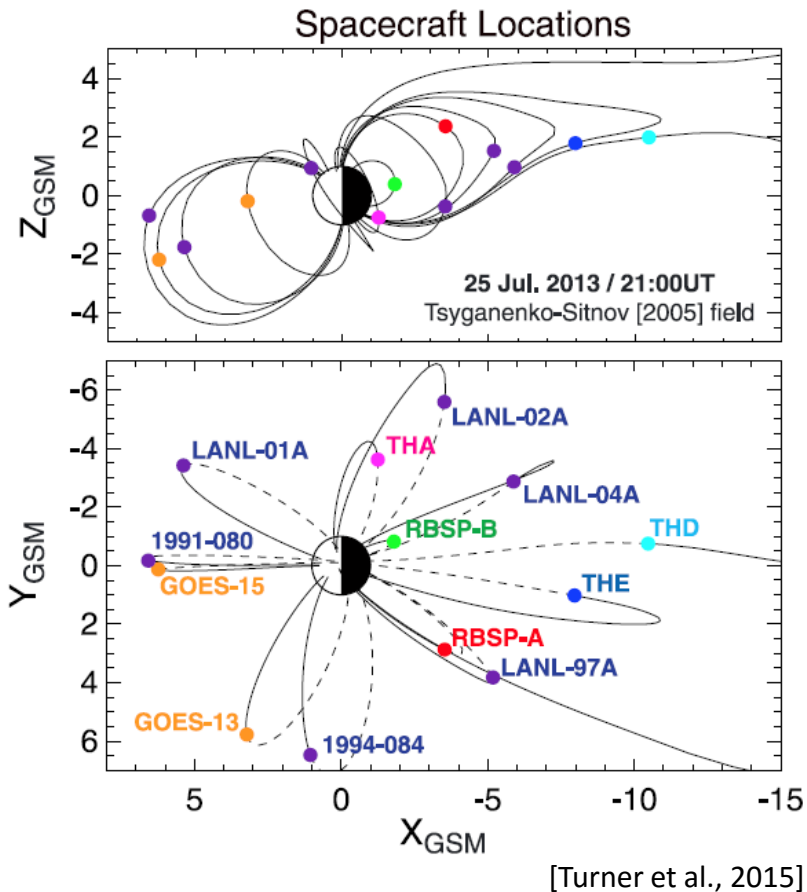
- These waves can affect the ring current
- Example: poloidal mode standing Alfvén waves excited by interplanetary shocks rapidly accelerate hydrogen and oxygen ions [Zong et al., 2012]



How do ULF waves couple to the ring current?

Example: Particle injections

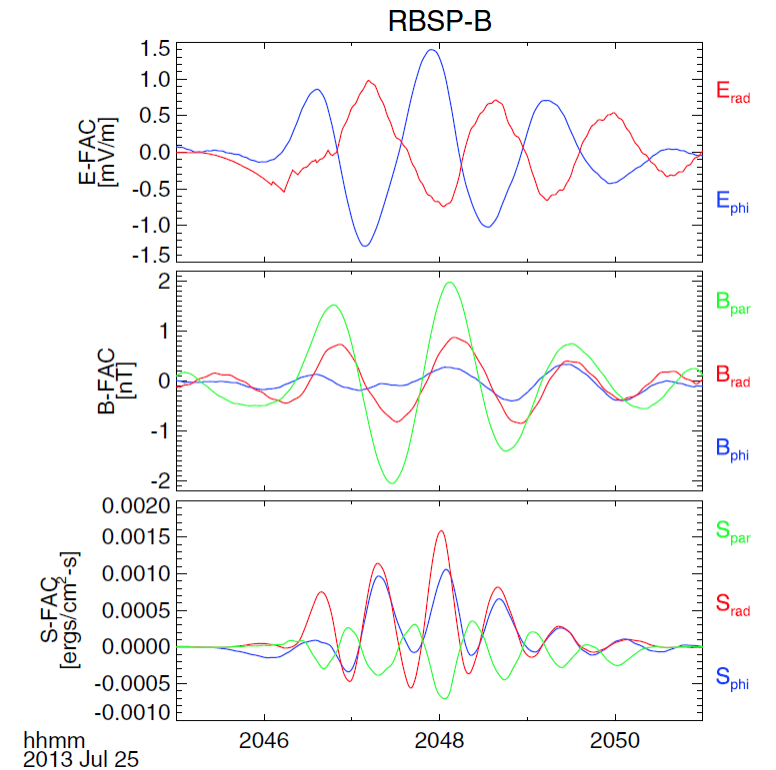
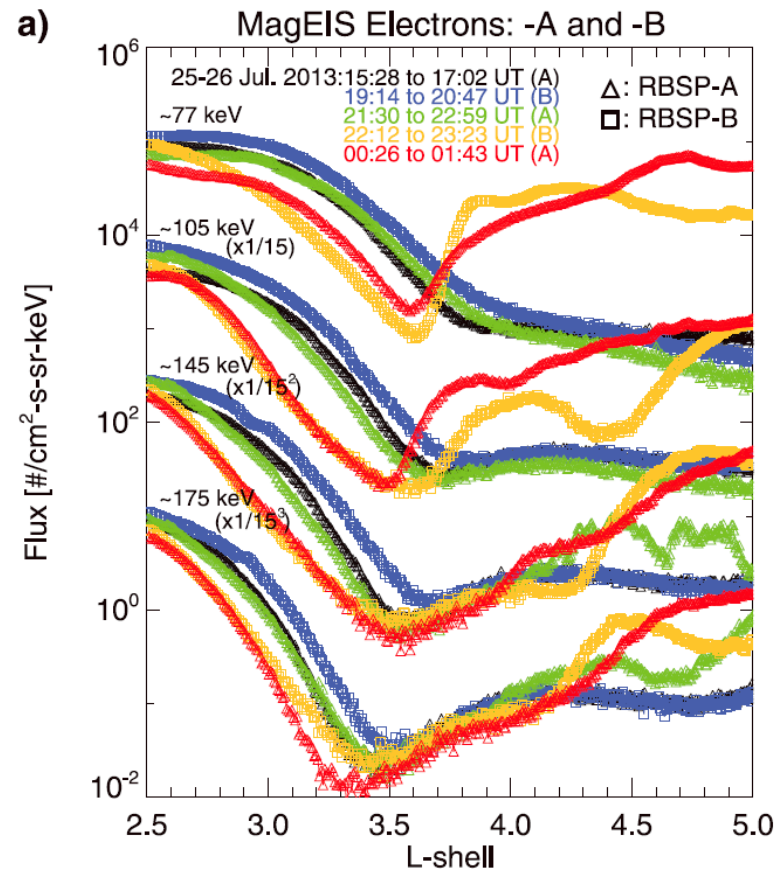
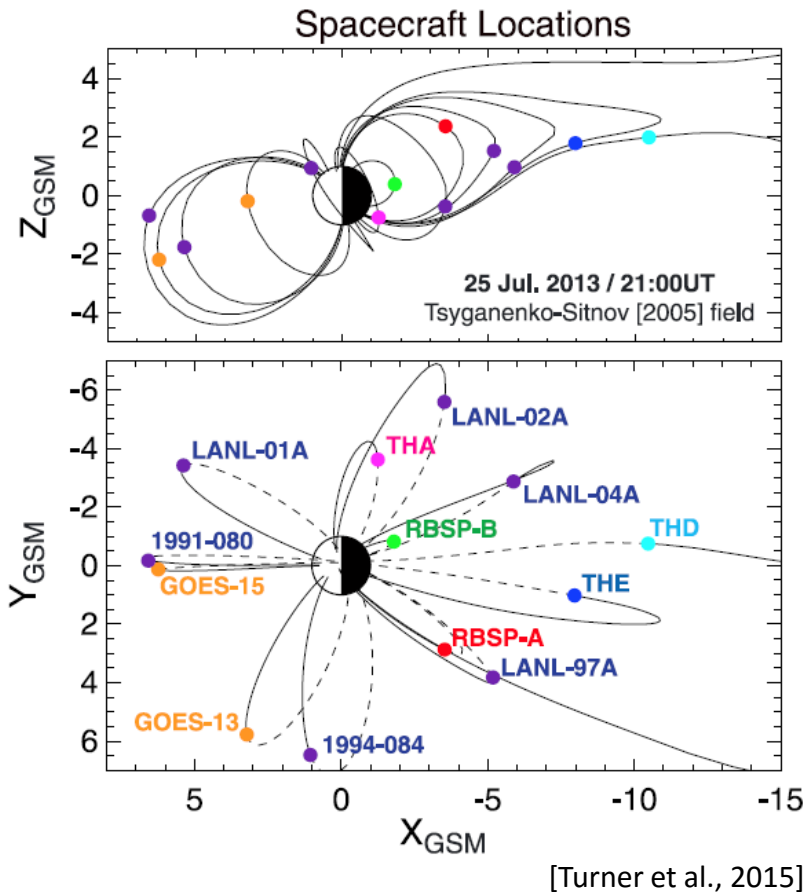
- During electron injection events, particles can be transported to extremely low L values



How do ULF waves couple to the ring current?

Example: Particle injections

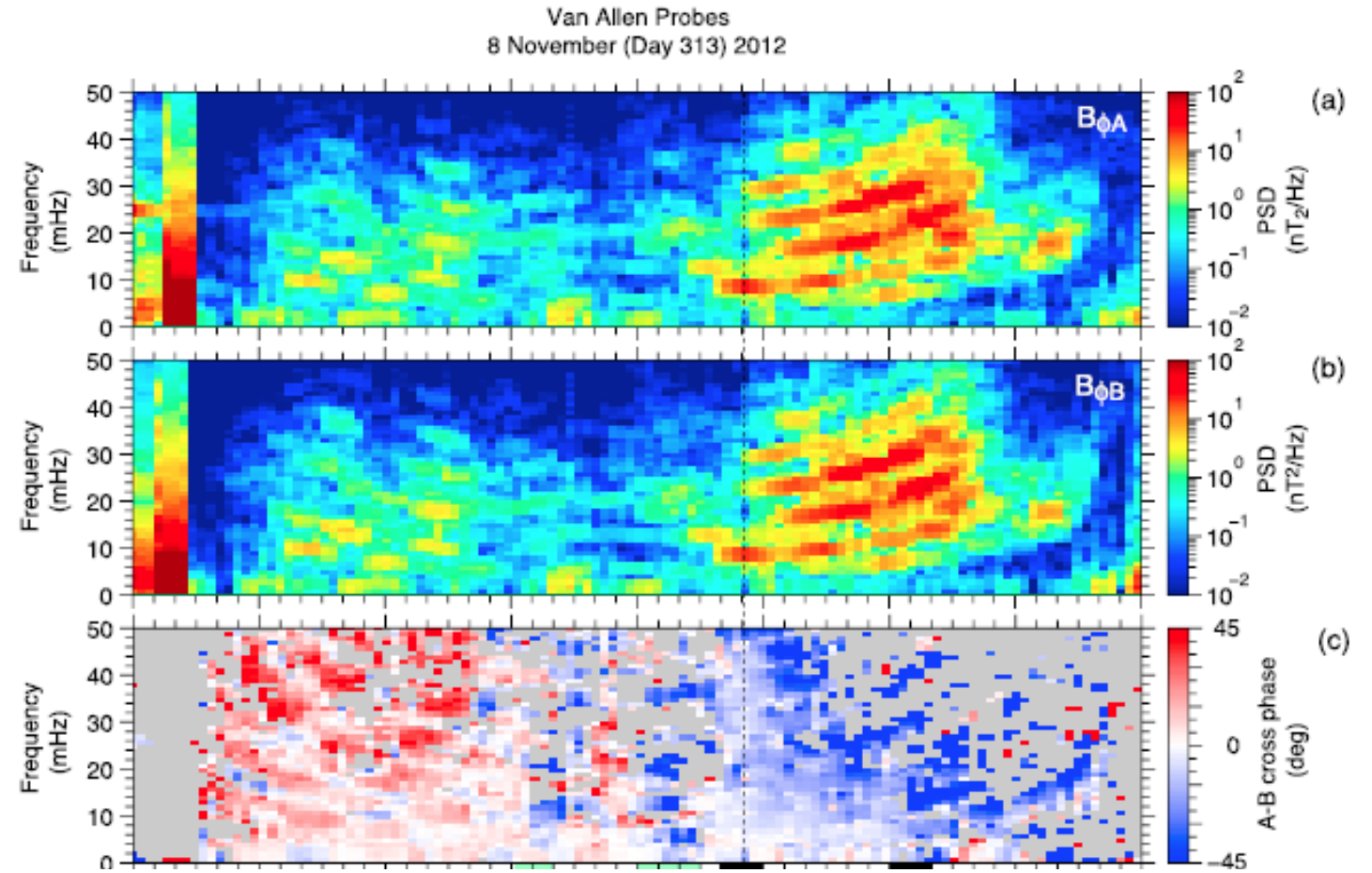
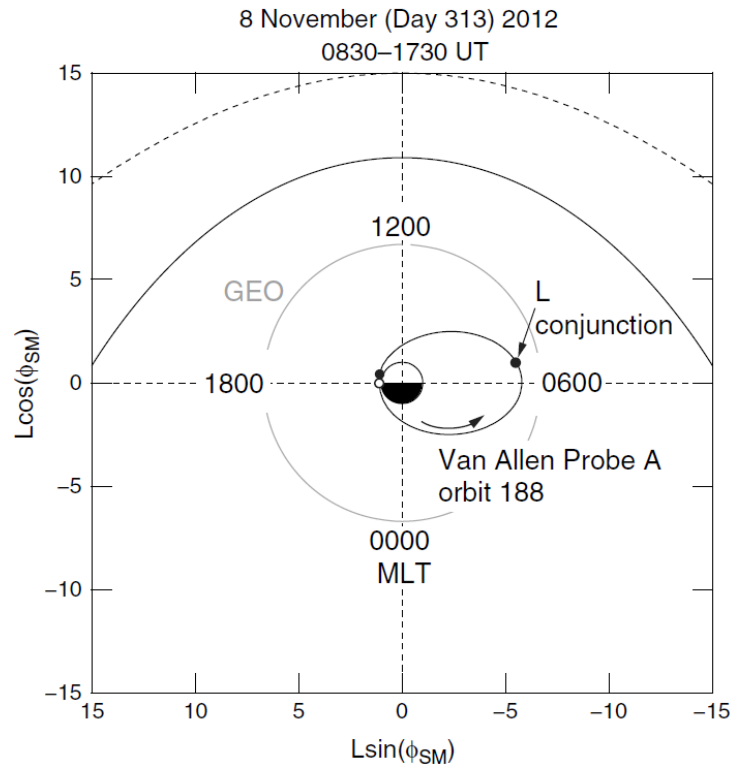
- During electron injection events, particles can be transported to extremely low L values
- Possible mechanism: rapid diffusion via plasmaspheric cavity mode waves



How do ULF waves couple to the plasmasphere?

Example: Standing Alfvén waves

- Multiple spacecraft distinguish between different wave modes
- Diagnostic for wave propagation, wave coherence scales in plasmasphere/outside plasmasphere

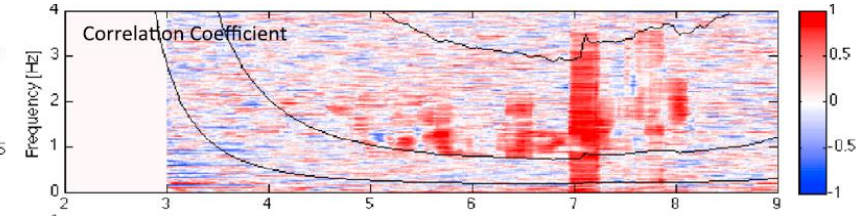
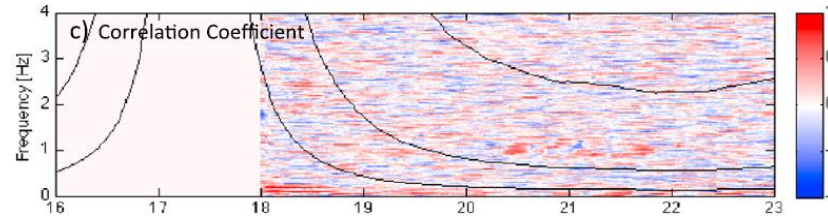


[Takahashi et al., 2015]

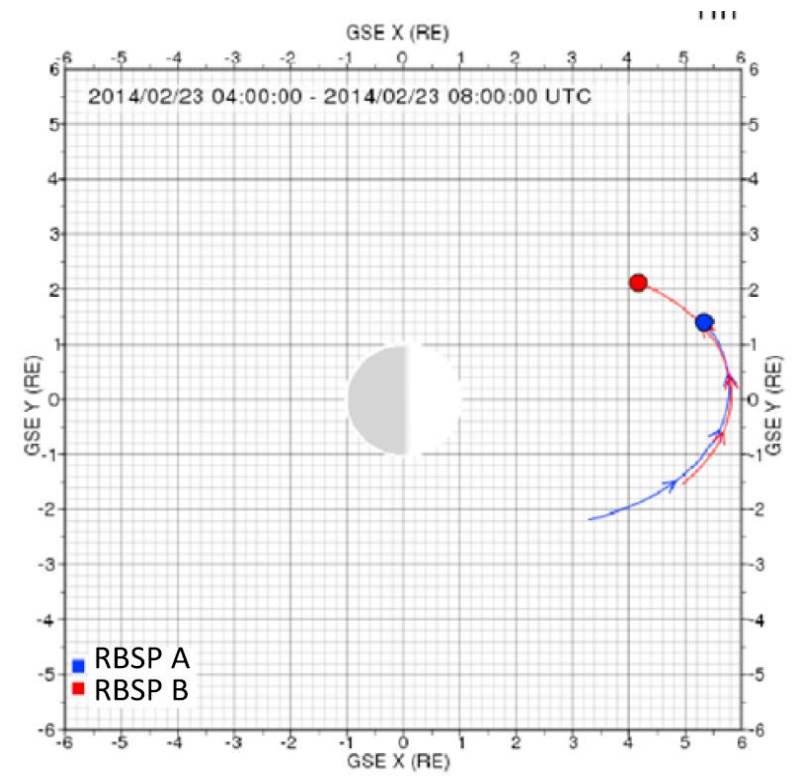
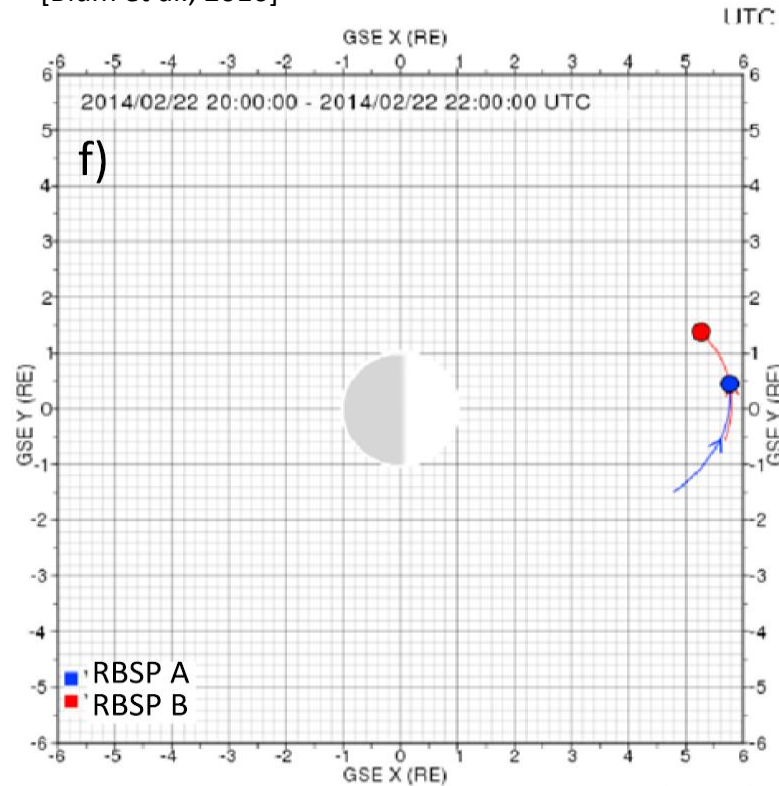
How do ULF waves couple to the plasmasphere?

Example: EMIC waves

- Use multiple spacecraft to detect coherence scales for EMIC waves

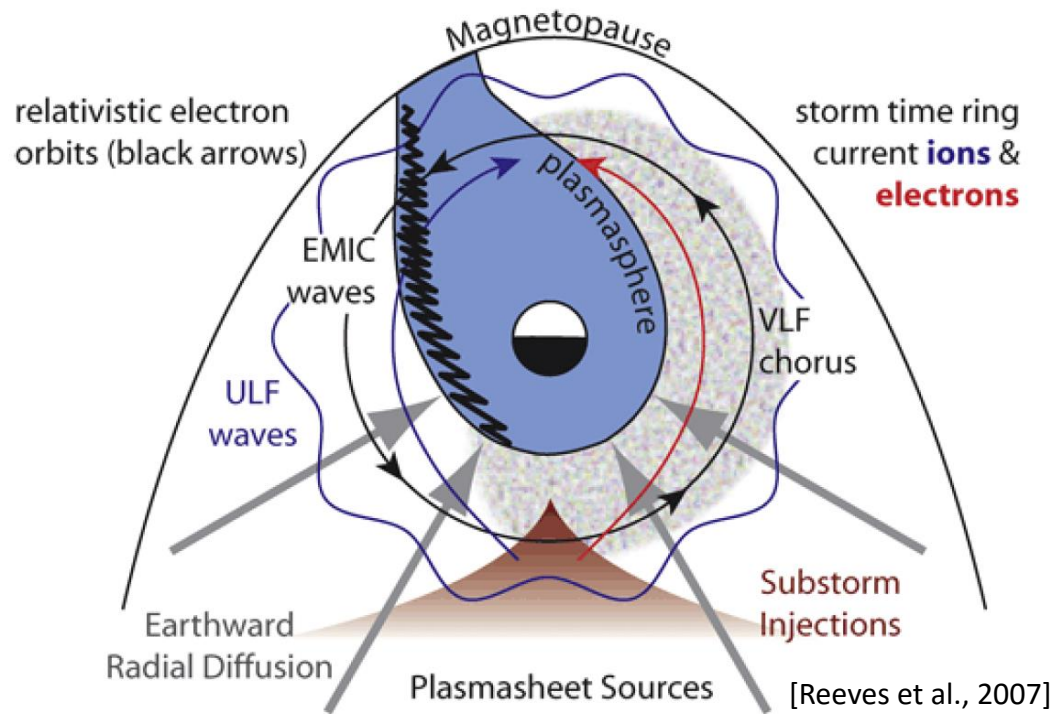


[Blum et al., 2016]



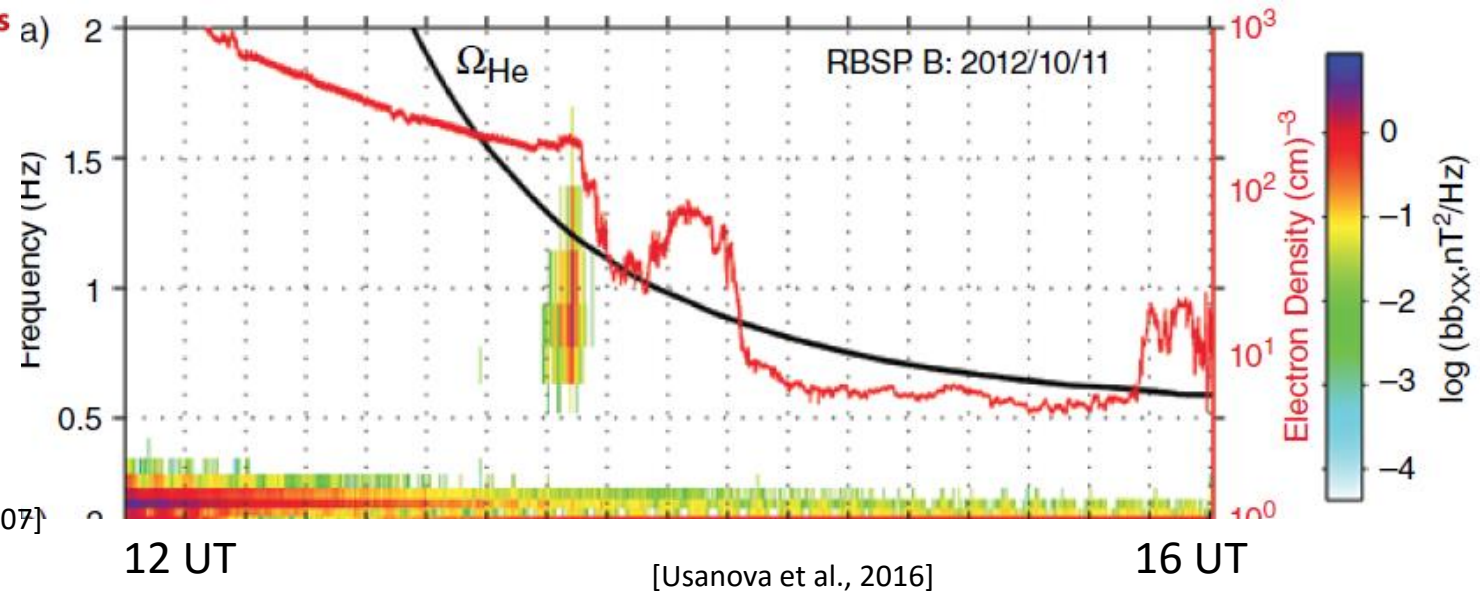
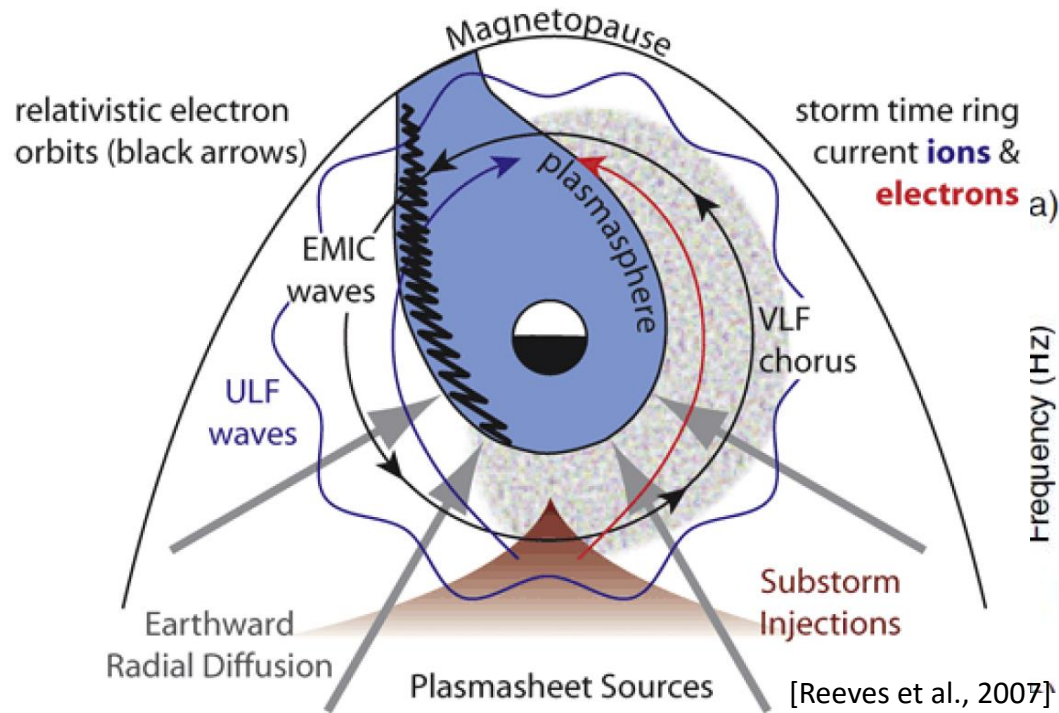
How do ULF waves couple to the plasmasphere?

Example: EMIC wave growth



How do ULF waves couple to the plasmasphere?

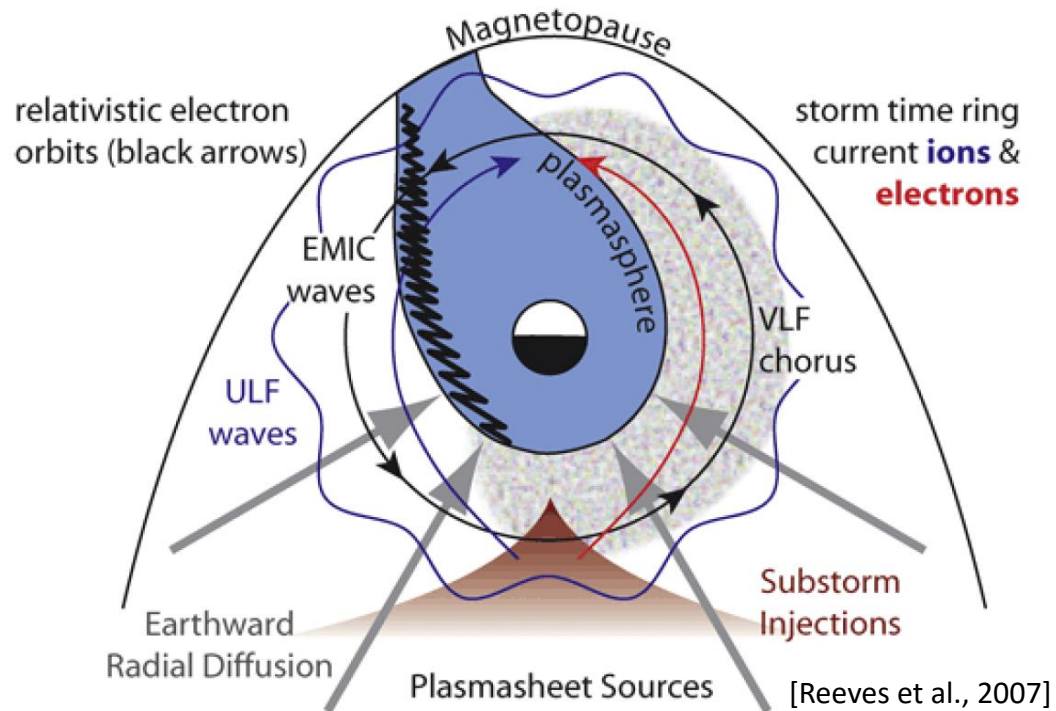
Example: EMIC wave growth



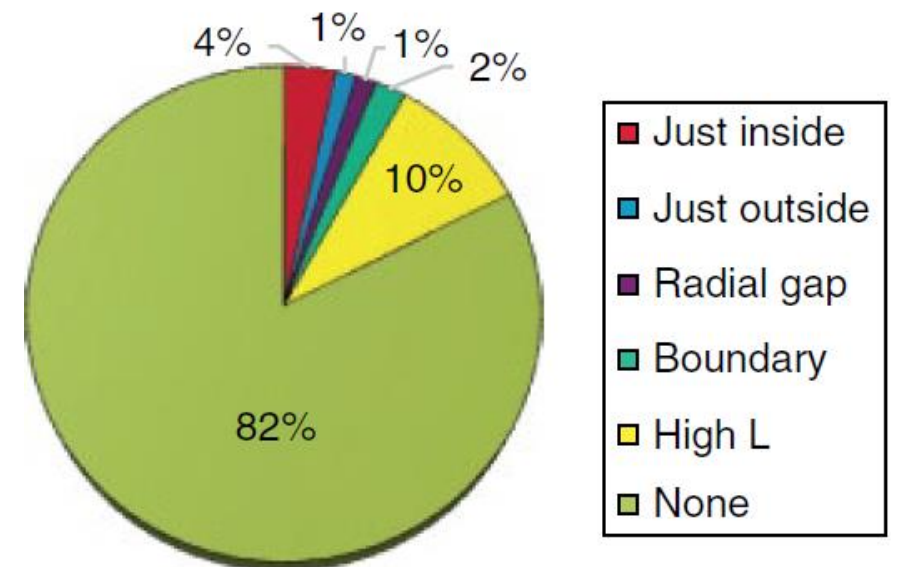
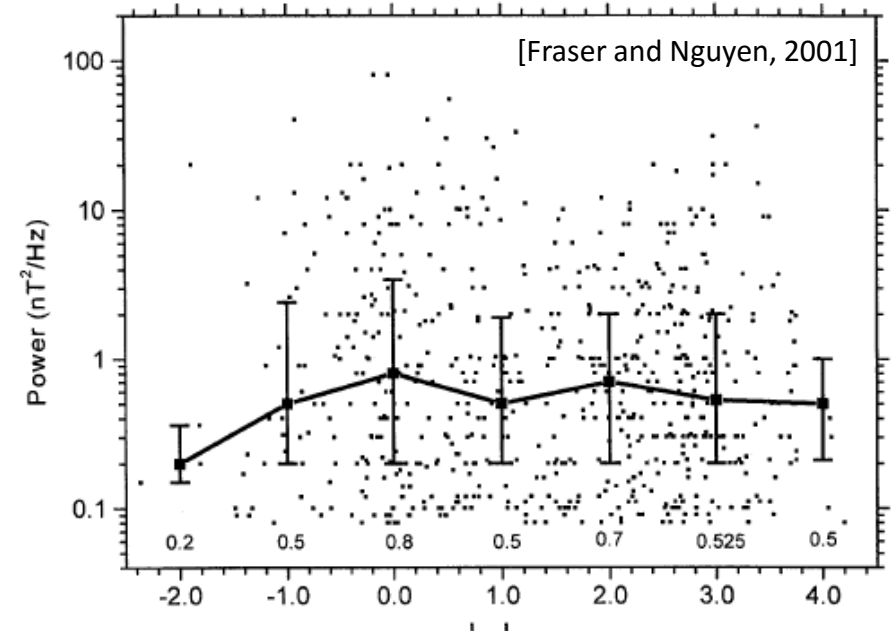
- There is correspondence in some cases

How do ULF waves couple to the plasmasphere?

Example: EMIC wave growth



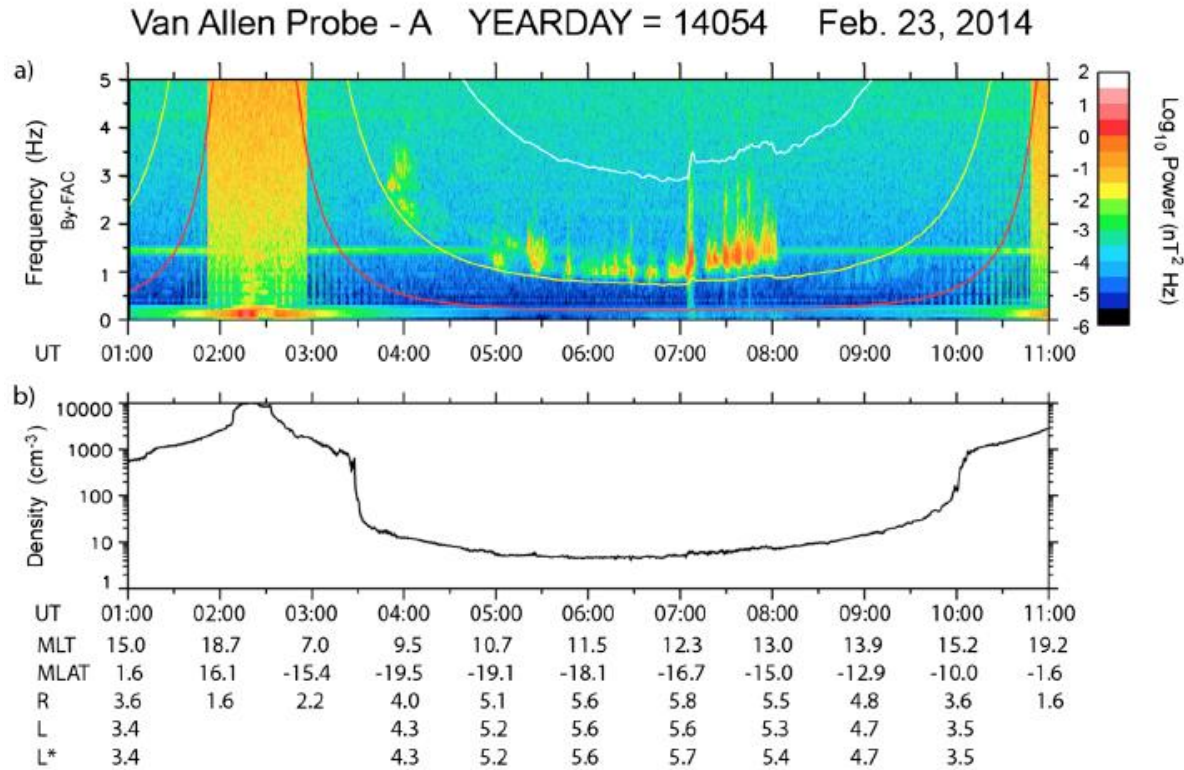
- There is correspondence in some cases
- However, EMIC wave occurrence is not as closely tied to the plasmopause as previously thought



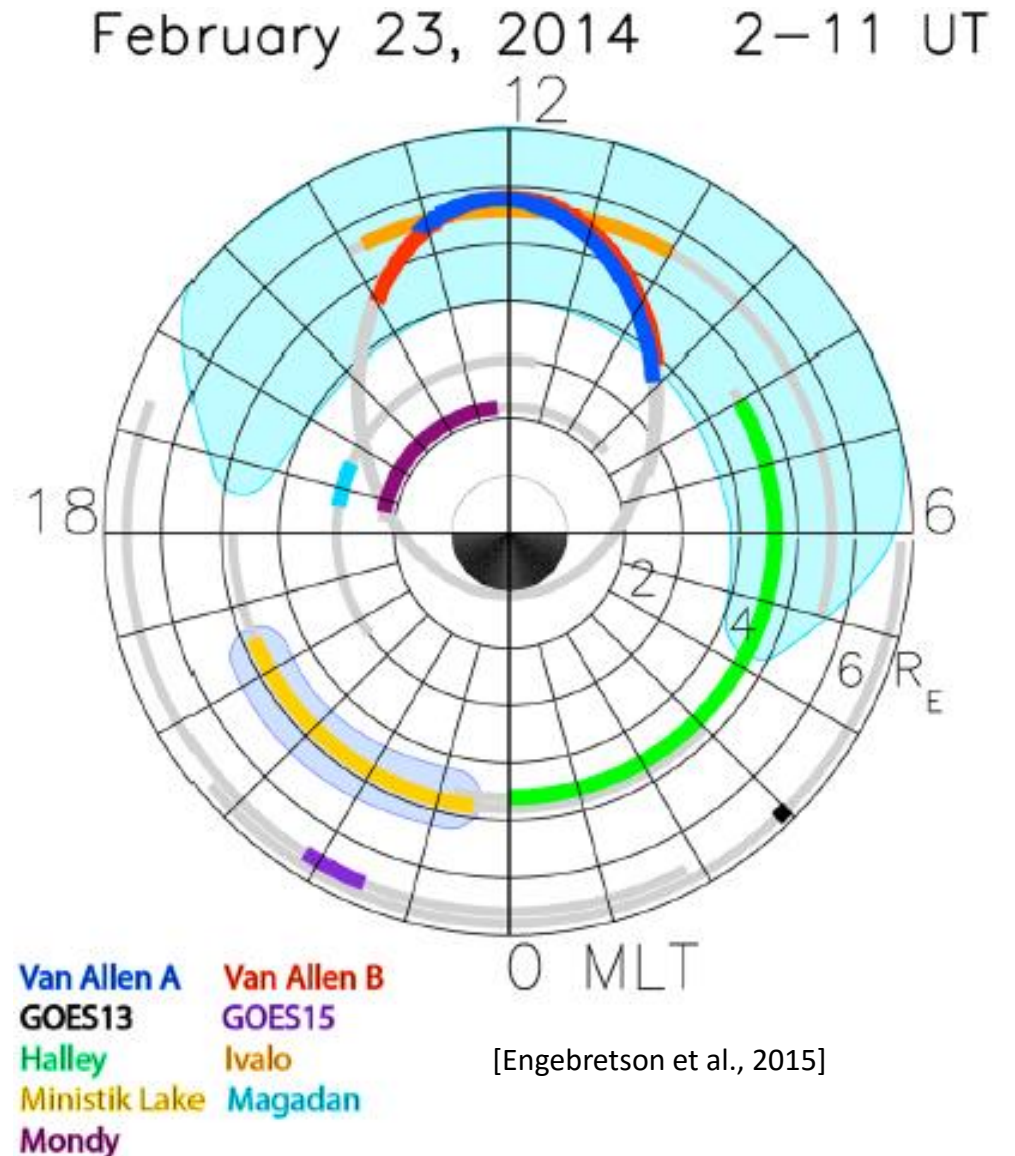
[Usanova et al., 2016]

How do ULF waves couple to the plasmasphere?

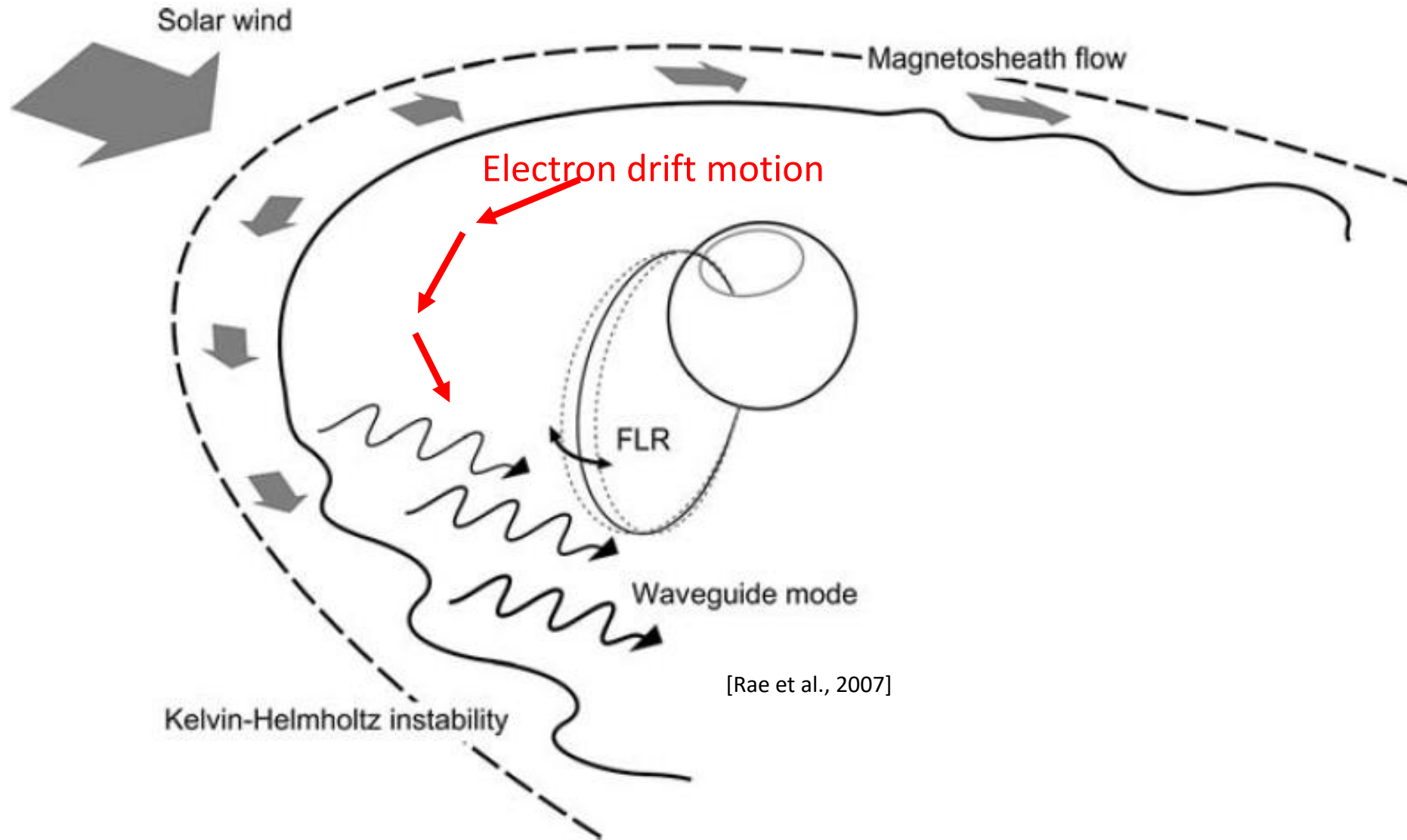
Example: EMIC wave growth



- In some cases, global observations show the plasmapause location matters very little for EMIC wave growth



How do ULF waves couple to the radiation belts?

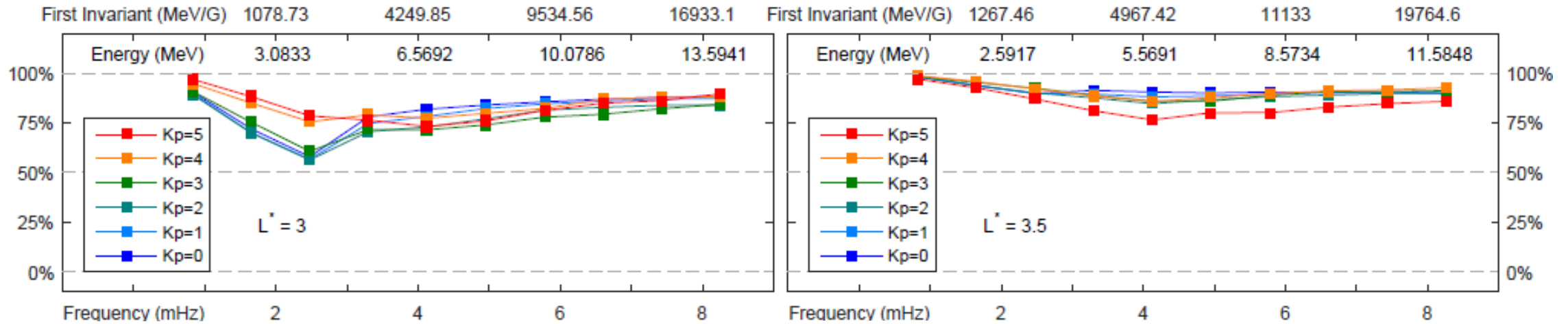


How do ULF waves couple to the radiation belts?

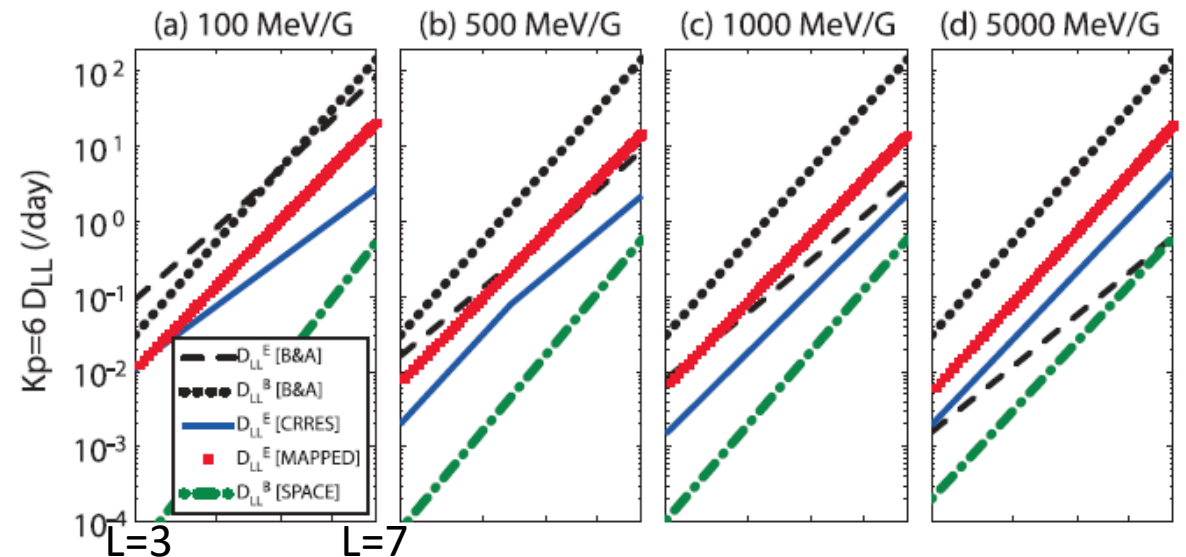
Example: New diffusion coefficients (1)

The Relative Contribution of the Electric Diffusion Coefficient D_{LL}^E/D_{LL}^{Total}

[Ali et al., 2016]



- New results from in situ measurements and ground-based magnetometers suggest the electric field diffusion coefficient is much more important than the magnetic field coefficient
- Contrary to previous results
- What will the effect be on models of diffusive transport via ULF waves

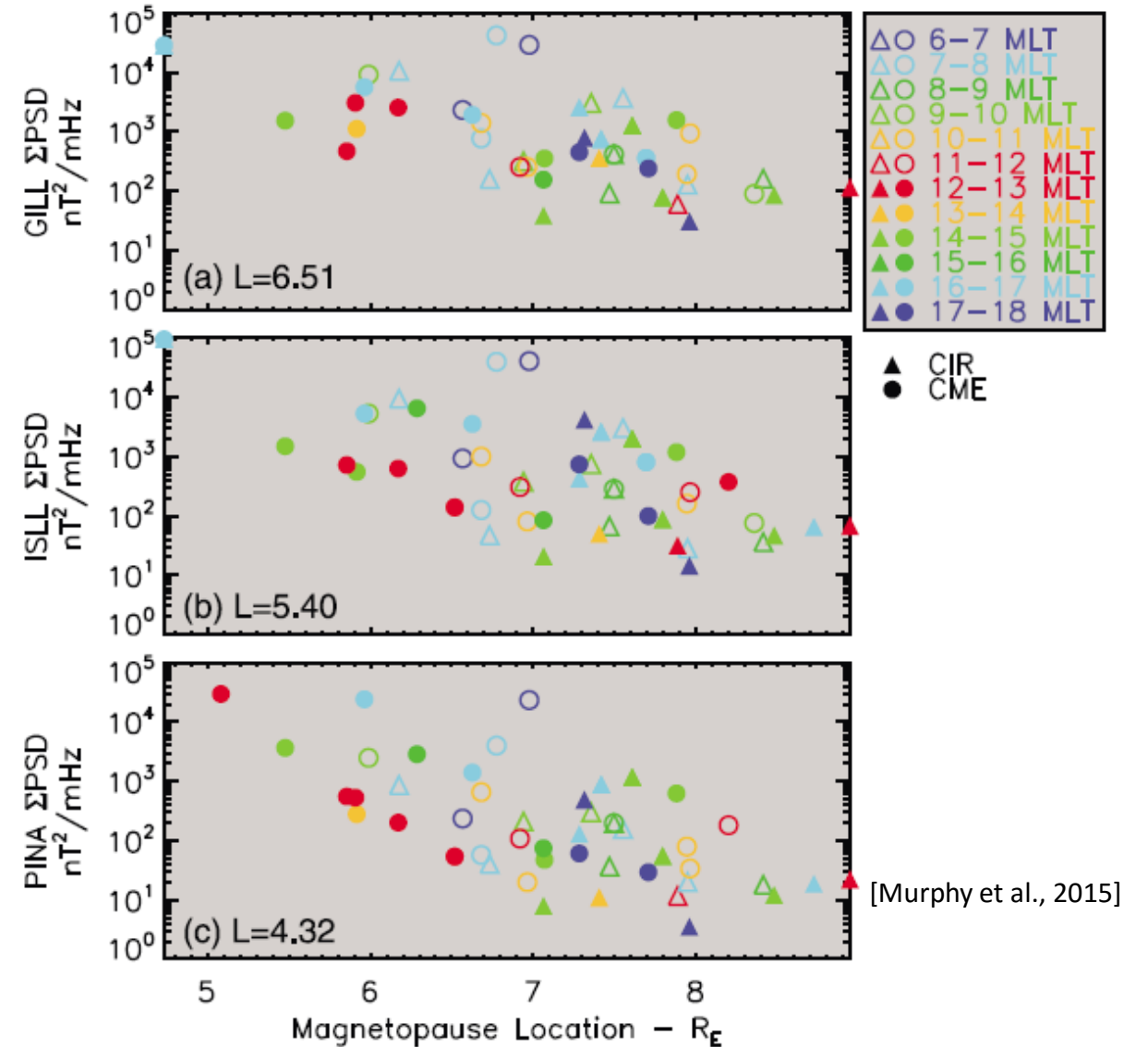


[Ozeke et al., 2012]

How do ULF waves couple to the radiation belts?

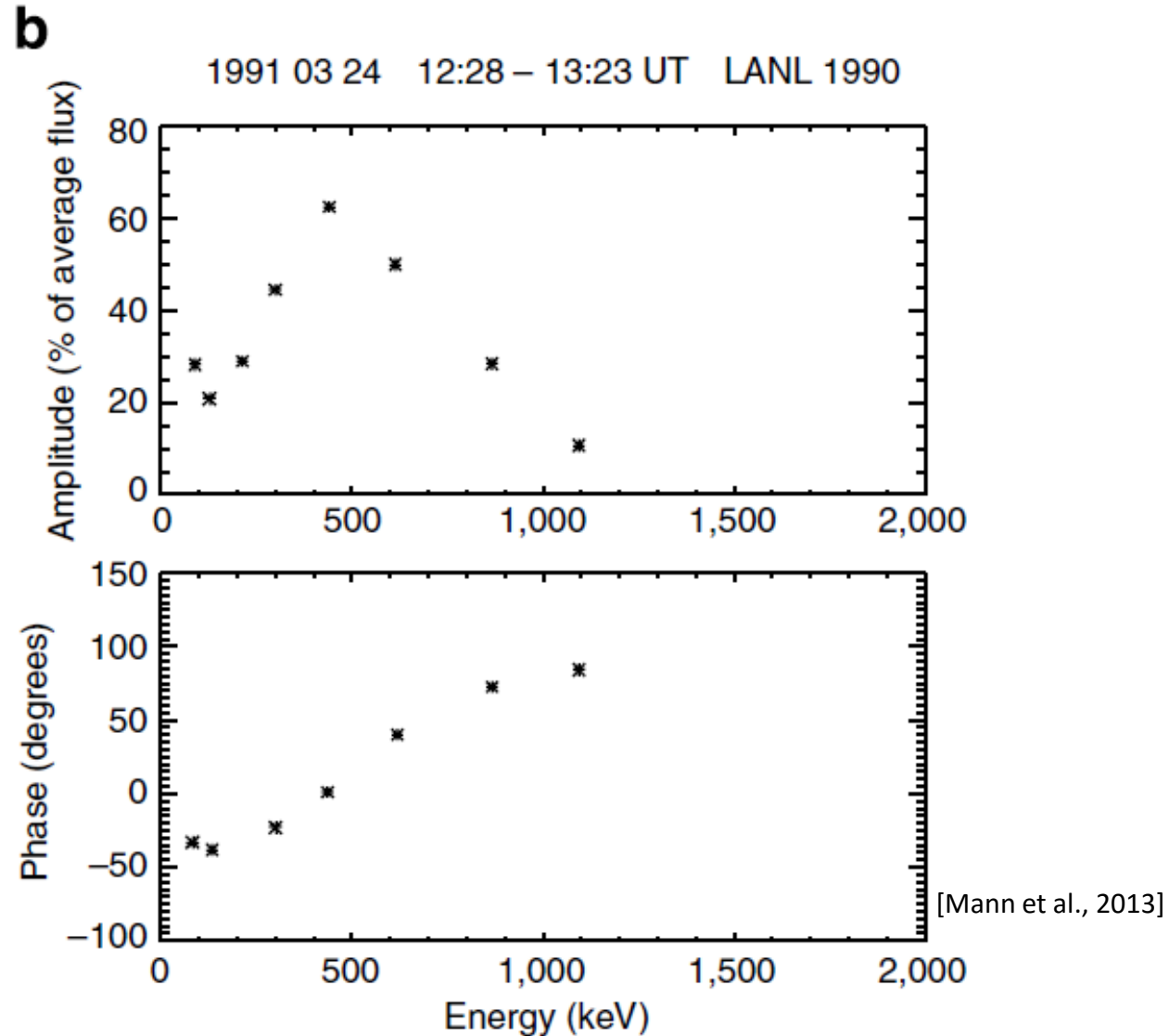
Example: New diffusion coefficients (2)

- ULF wave amplitudes during storms depends strongly on position relative to magnetopause
- Diffusion coefficients during periods when the magnetosphere is compressed – e.g., during CIR/CME storms – can be substantially different
- Need parameterizations beyond Kp (e.g., Dimitrakoudis et al., 2015)



How do ULF waves couple to the radiation belts?

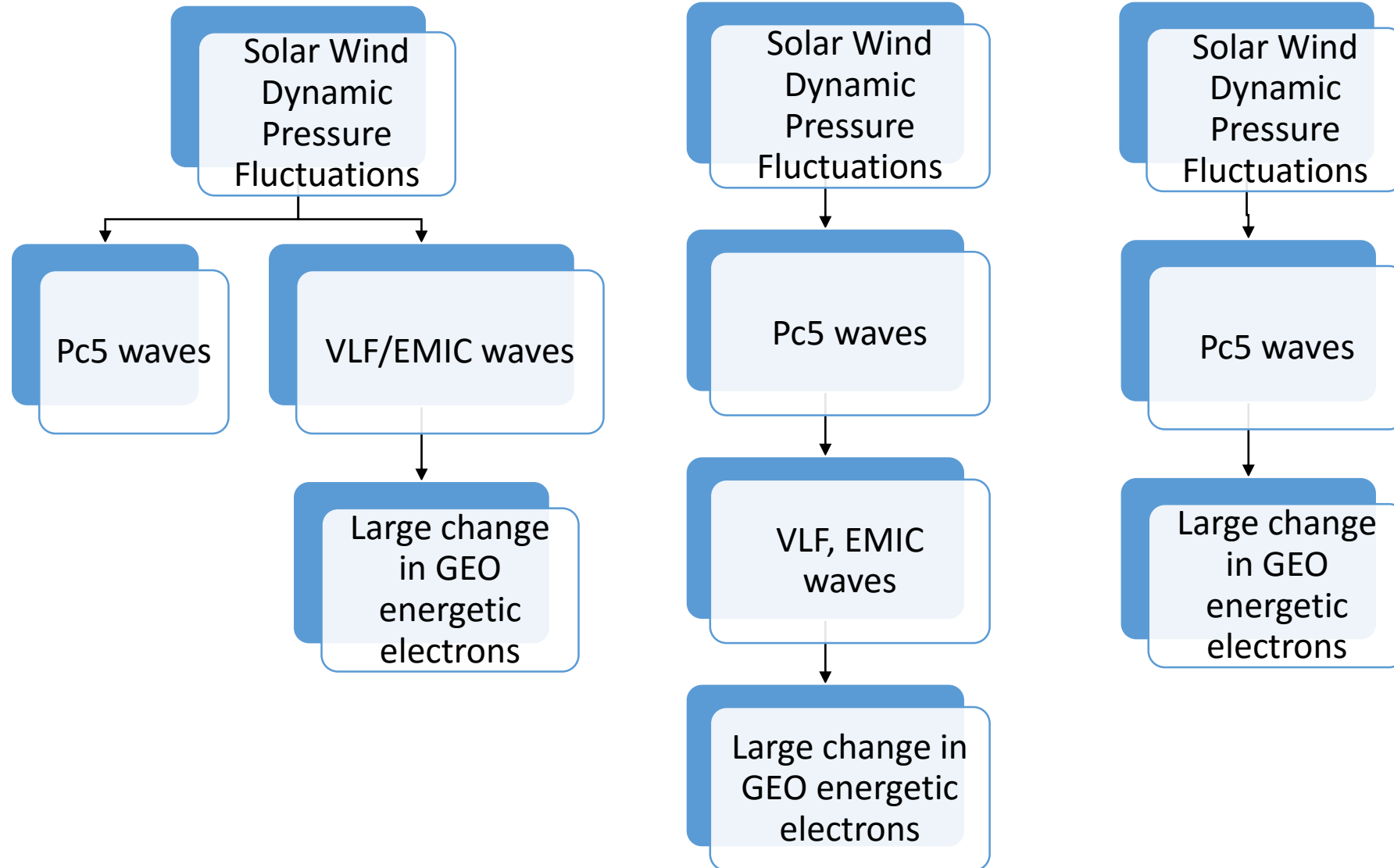
Example: Drift resonance via Pc5 waves



- Drift resonance with electron energies ~ 500 keV
- Leads to rapid transport at select energy ranges

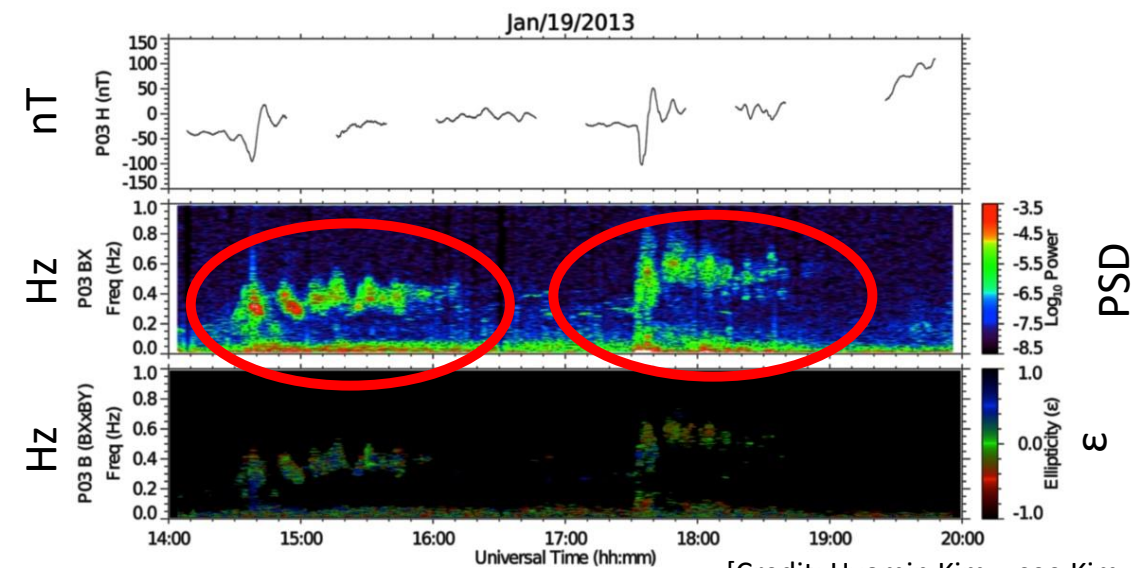
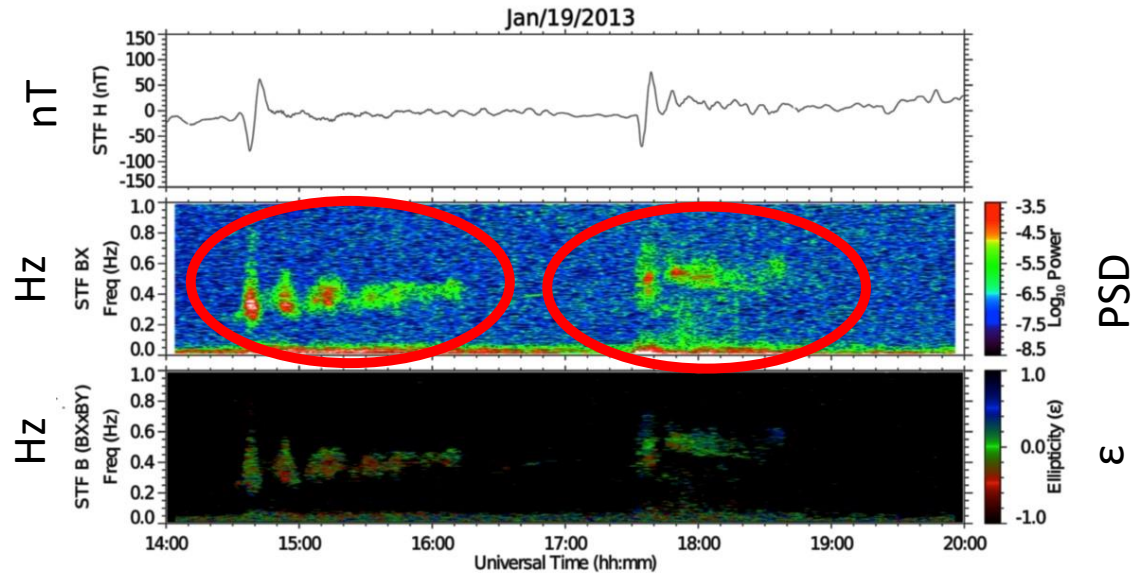
How do ULF waves couple to the radiation belts?

Example: When/where/how do Pc5 waves affect radiation belt dynamics



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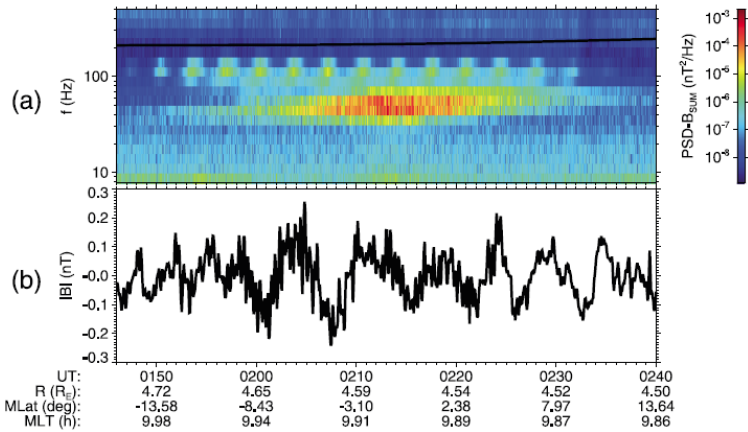
- Pc5 (2-7 mHz) wave amplitudes are anti-correlated with EMIC wave amplitudes in many cases (e.g., Loto'aniu et al., 2009 and previous studies)

[Credit: Hyomin Kim – see Kim et al., 2015]

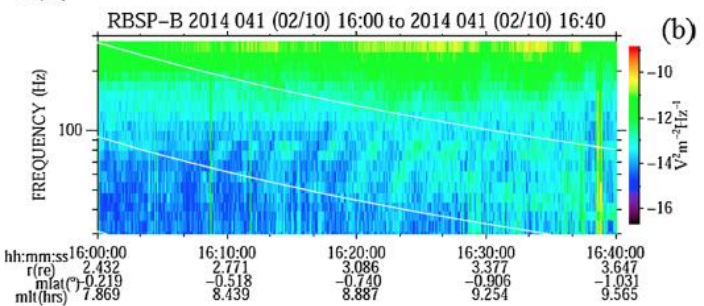
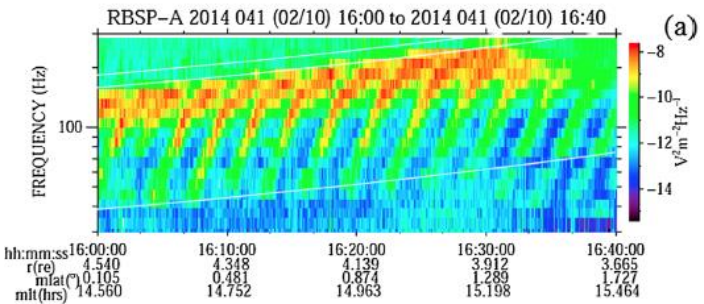
How do ULF waves couple to the radiation belts?

Example: When/where/how do Pc5 waves affect radiation belt dynamics

Equatorial noise/magnetosonic



[Nemec et al., 2015]



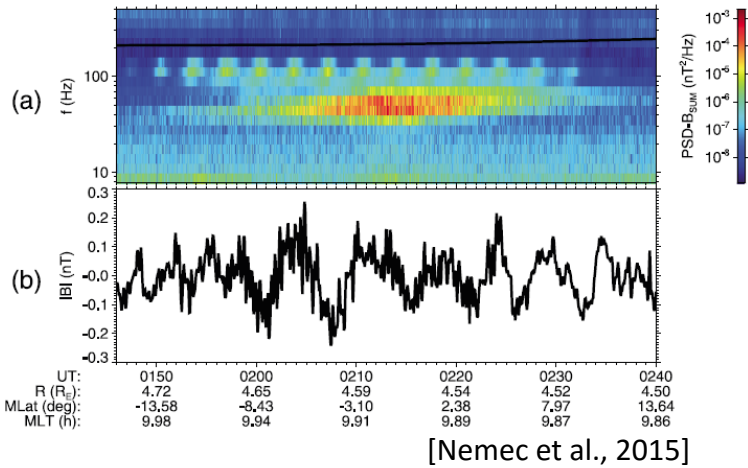
[Boardsen et al., 2014]

- In situ observations are shedding new light on previous ground-based observations/theory of ULF modulated VLF waves (QP)
- Many VLF wave modes have amplitudes that peak/dip at ULF frequencies – not all are necessarily linked to ULF waves, and more than one ULF wave mode is implicated

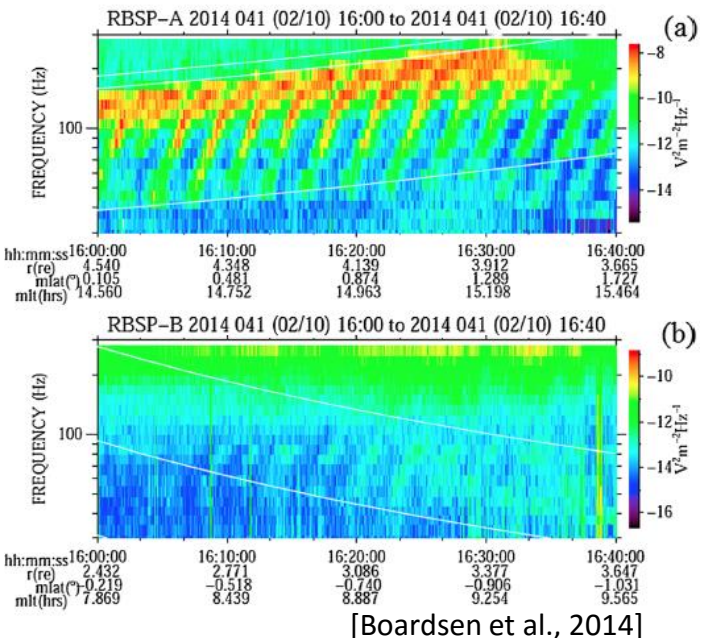
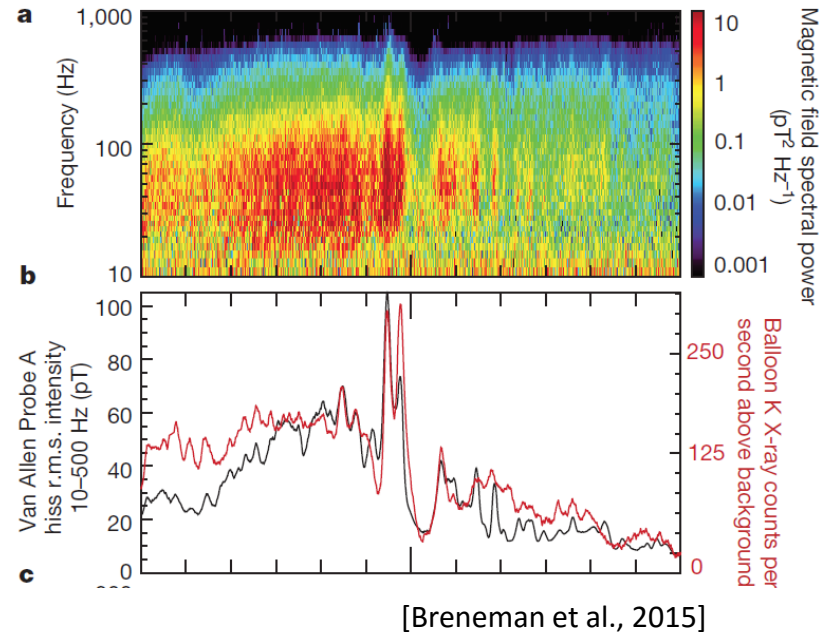
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Equatorial noise/magnetosonic



Plasmaspheric Hiss

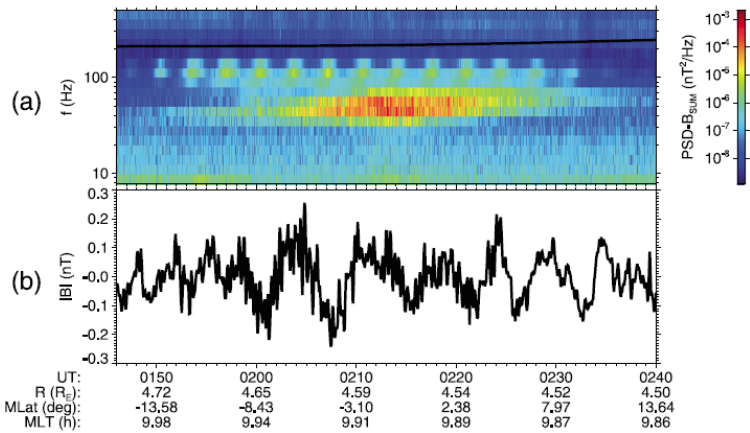


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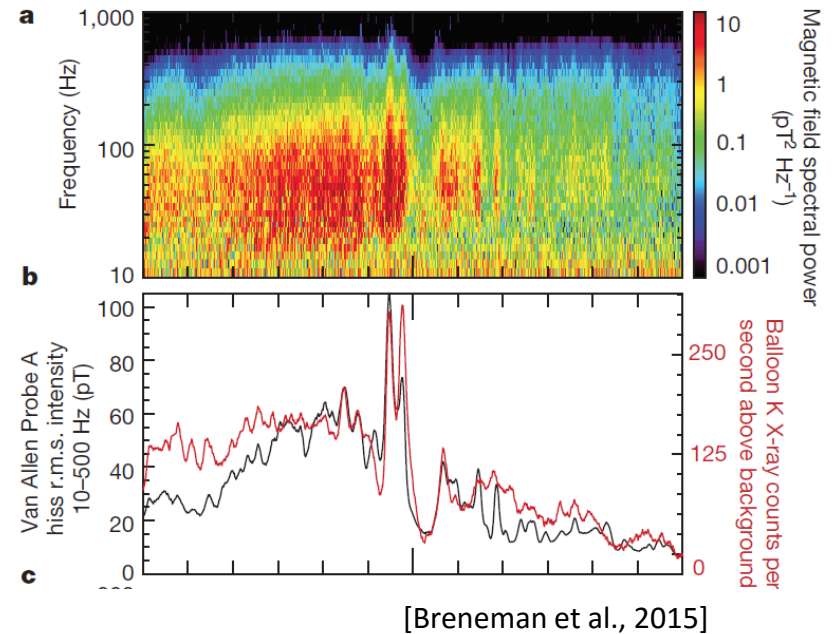
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Example: When/where/how do Pc5 waves affect radiation belt dynamics

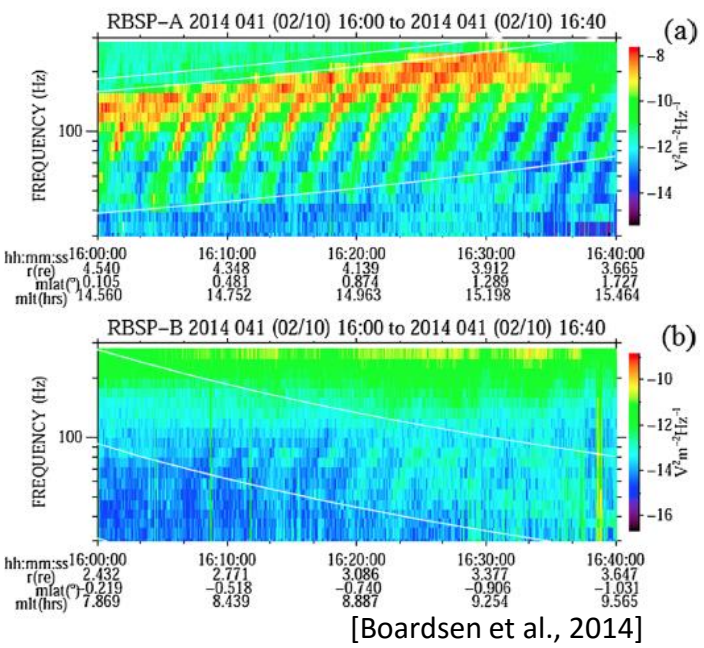
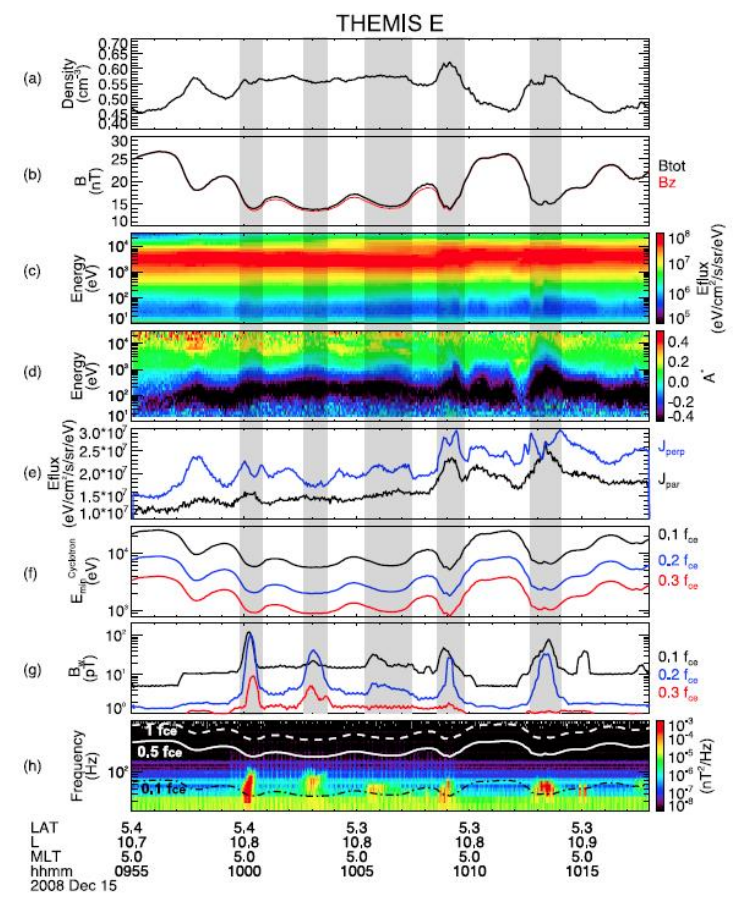
Equatorial noise/magnetosonic



Plasmaspheric Hiss



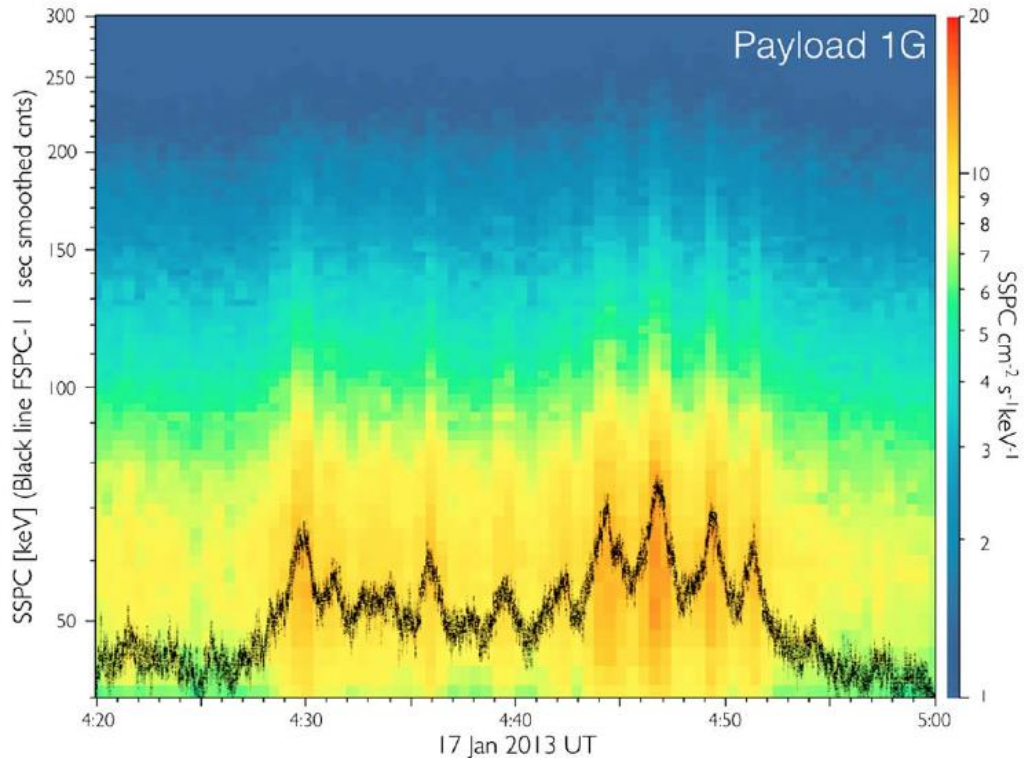
Whistler-mode Chorus



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How do ULF waves couple to the radiation belts?

Example: When/where/how do Pc5 waves affect radiation belt dynamics

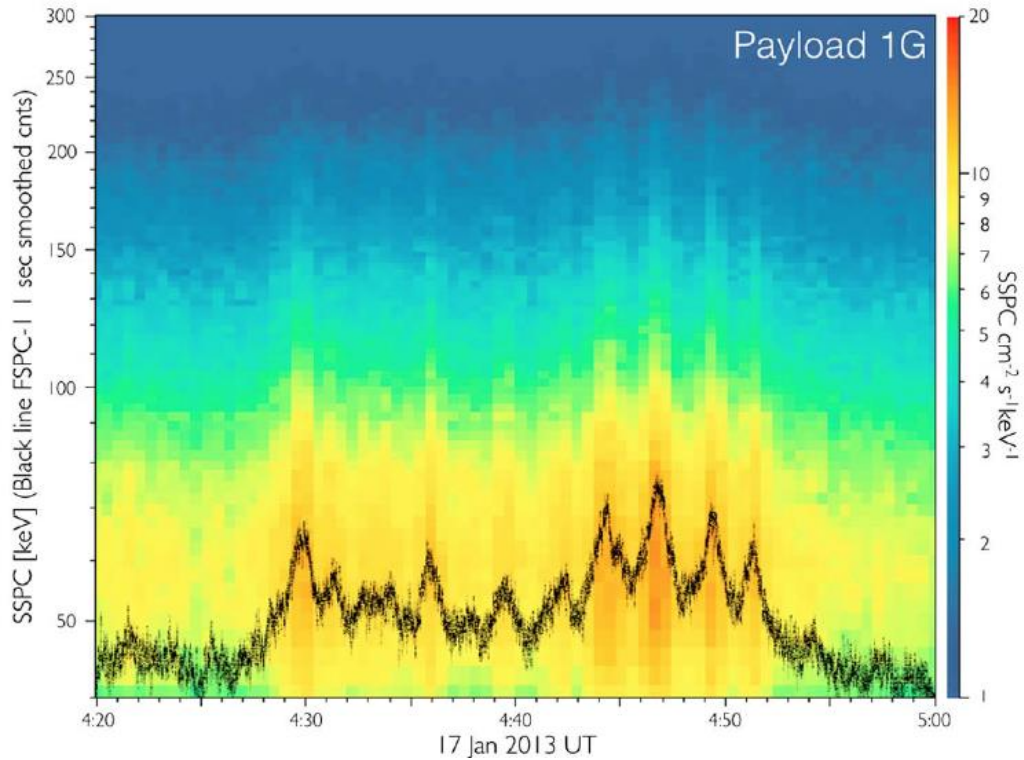


[Brito et al., 2015]

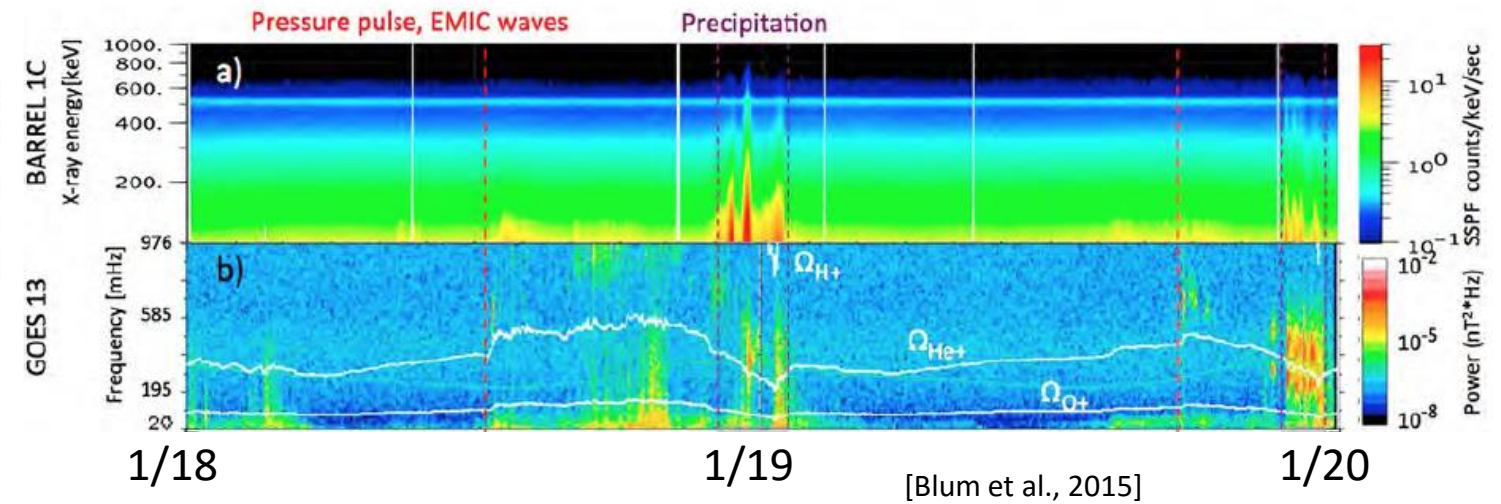
- ULF waves with frequencies <10 mHz have been linked to relativistic electron precipitation in the absence of VLF/EMIC waves (above, Brito et al., 2015)

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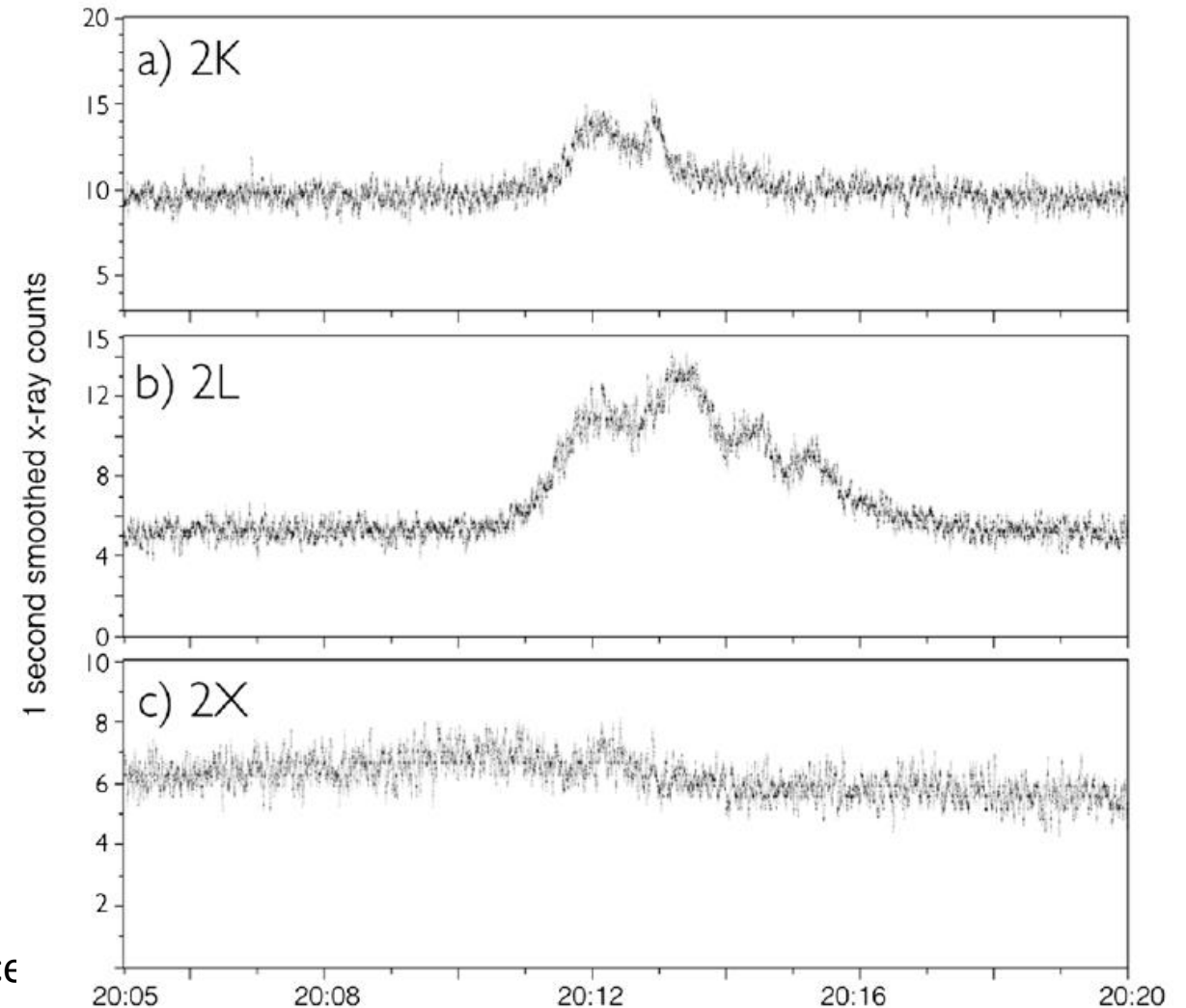
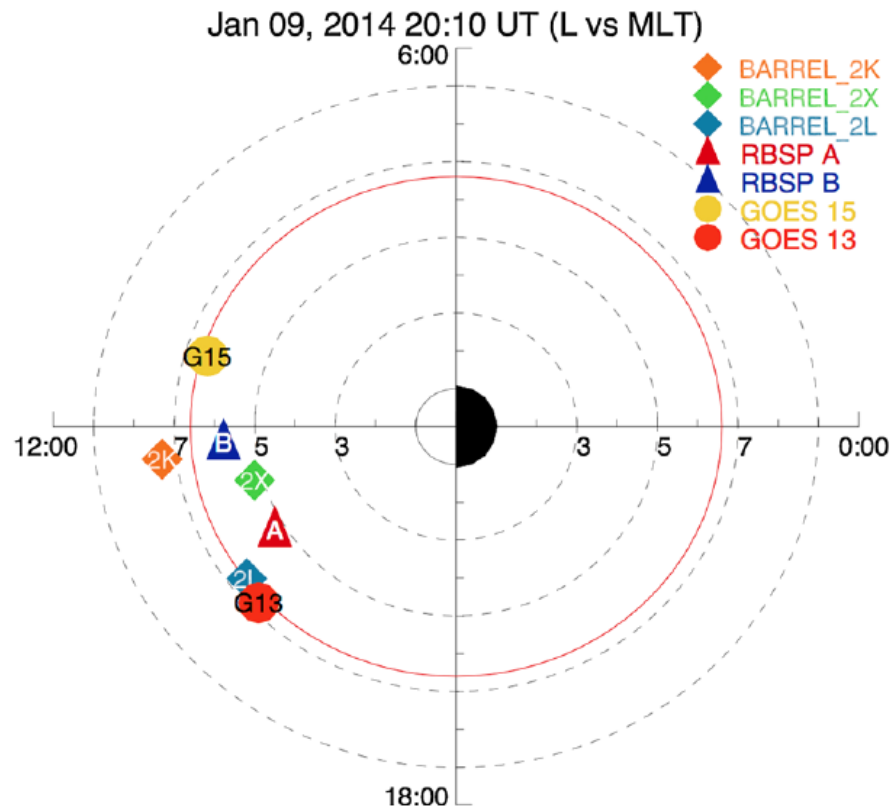


[Blum et al., 2015]

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- Recent BARREL observations are also providing strong evidence for direct links between EMIC waves and precipitation

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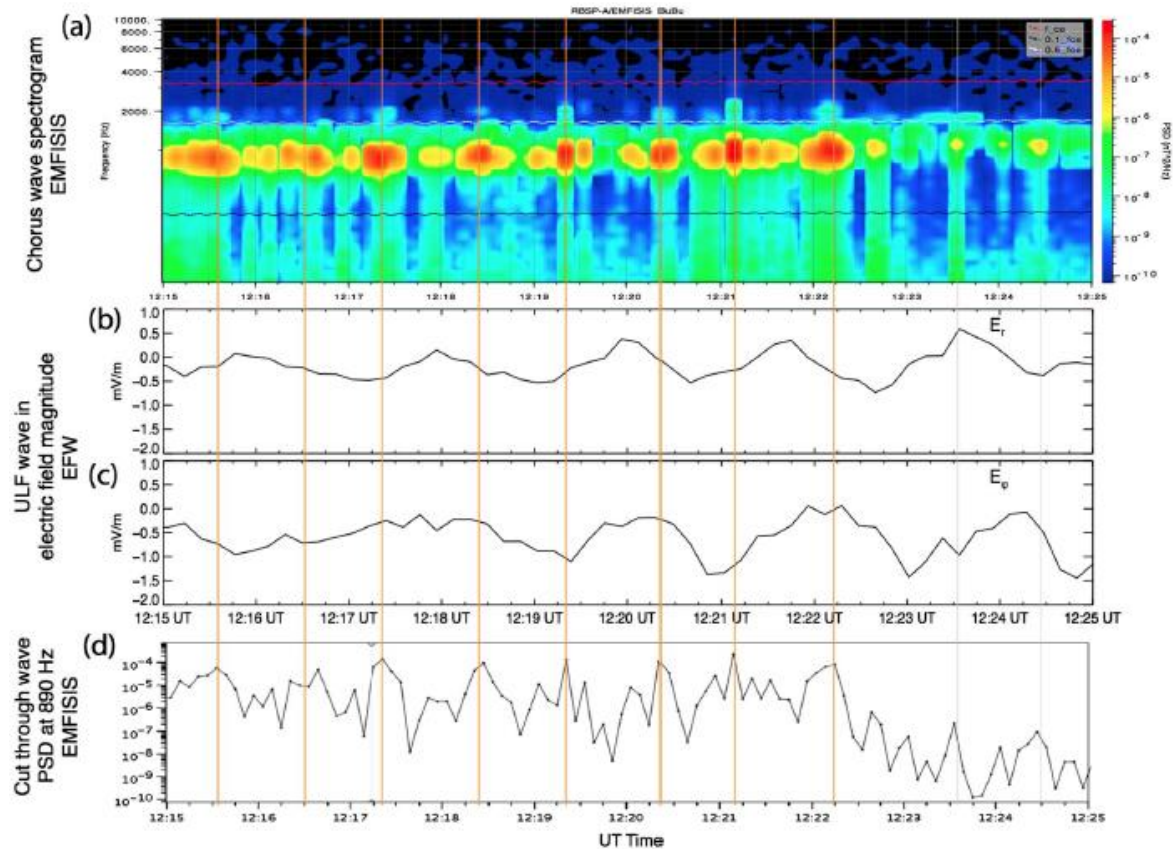
- Electron precipitation via VLF waves modulated by ULF waves (as well as electric field impulses)
- Recent BARREL observations are also providing strong evidence for direct links between ULF modulated-VLF waves and precipitation

9 January 2014

[Halford et al., 2015]

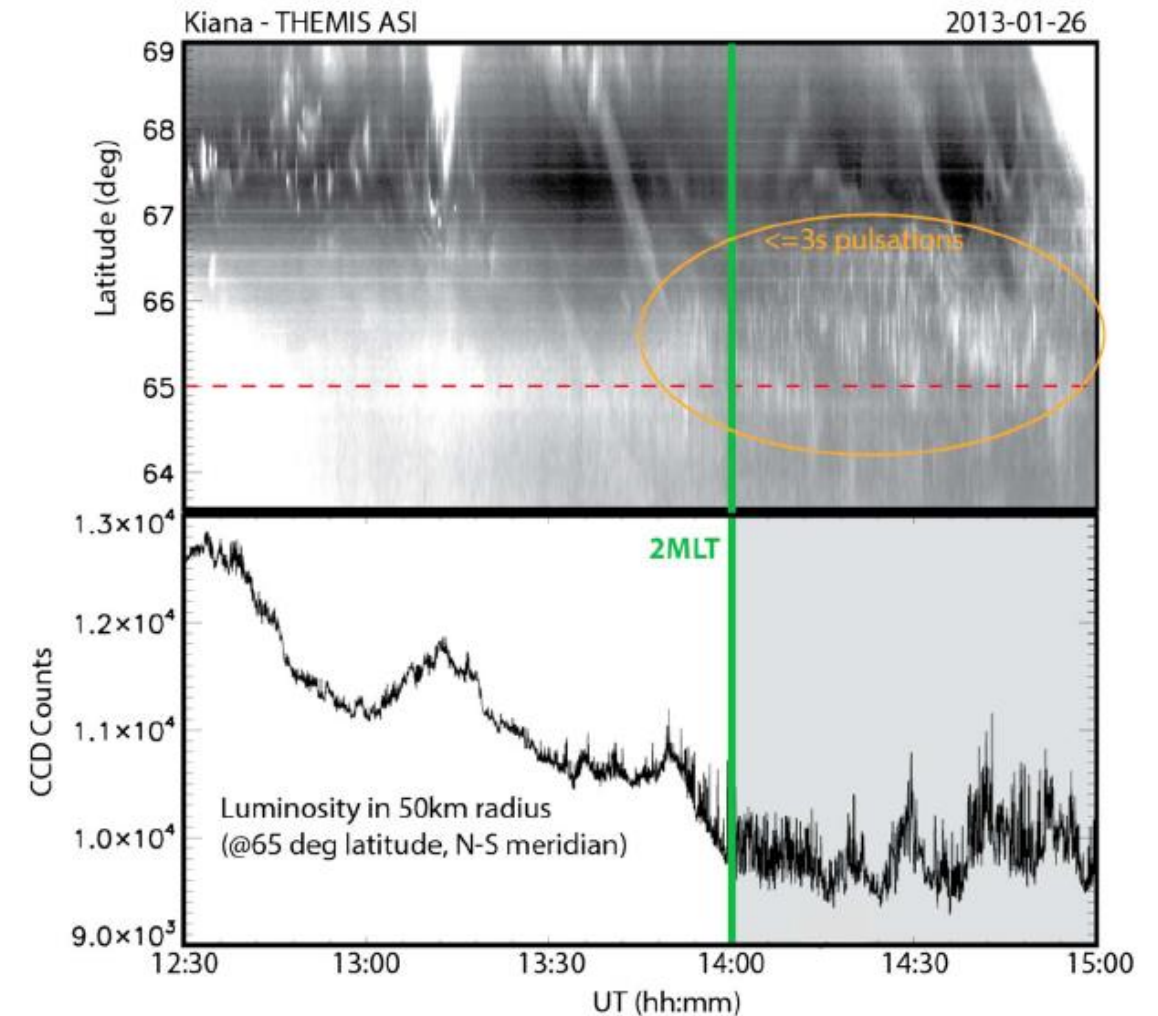
What is the role of ULF waves in Magnetosphere-Ionosphere Coupling?

Example: ULF modulation of auroral precipitation (1)



[Jaynes et al., 2015]

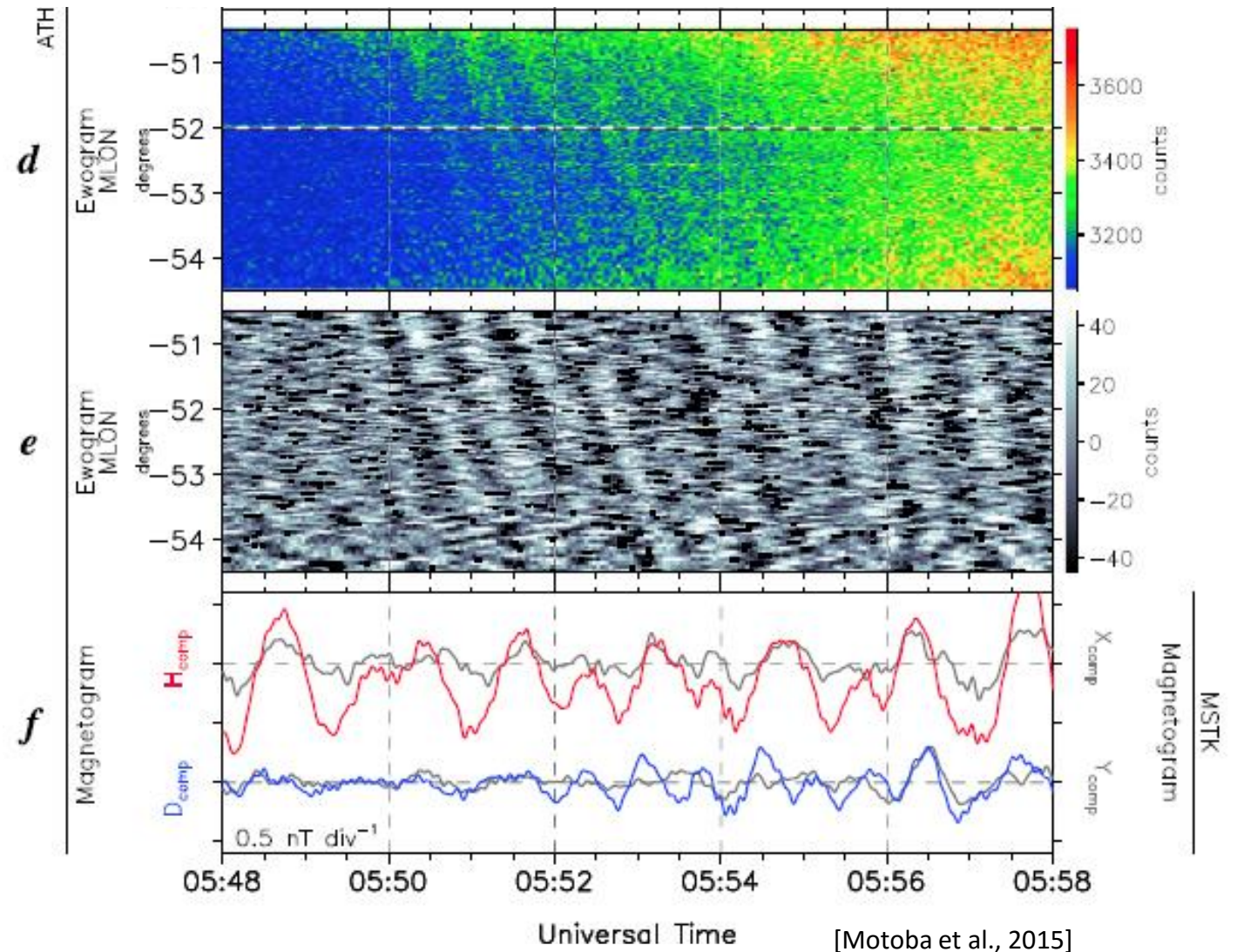
- Simultaneous observations of ULF wave properties, VLF wave properties, and auroral modulation
- ULF waves grow due to substorm injection, modulate chorus waves that in turn cause precipitation as they grow



What is the role of ULF waves in Magnetosphere-Ionosphere Coupling?

Example: ULF modulation of auroral precipitation (2)

- ULF modulation of VLF waves leading to auroral structures
- Direct interaction with ULF waves can also cause auroral modulation

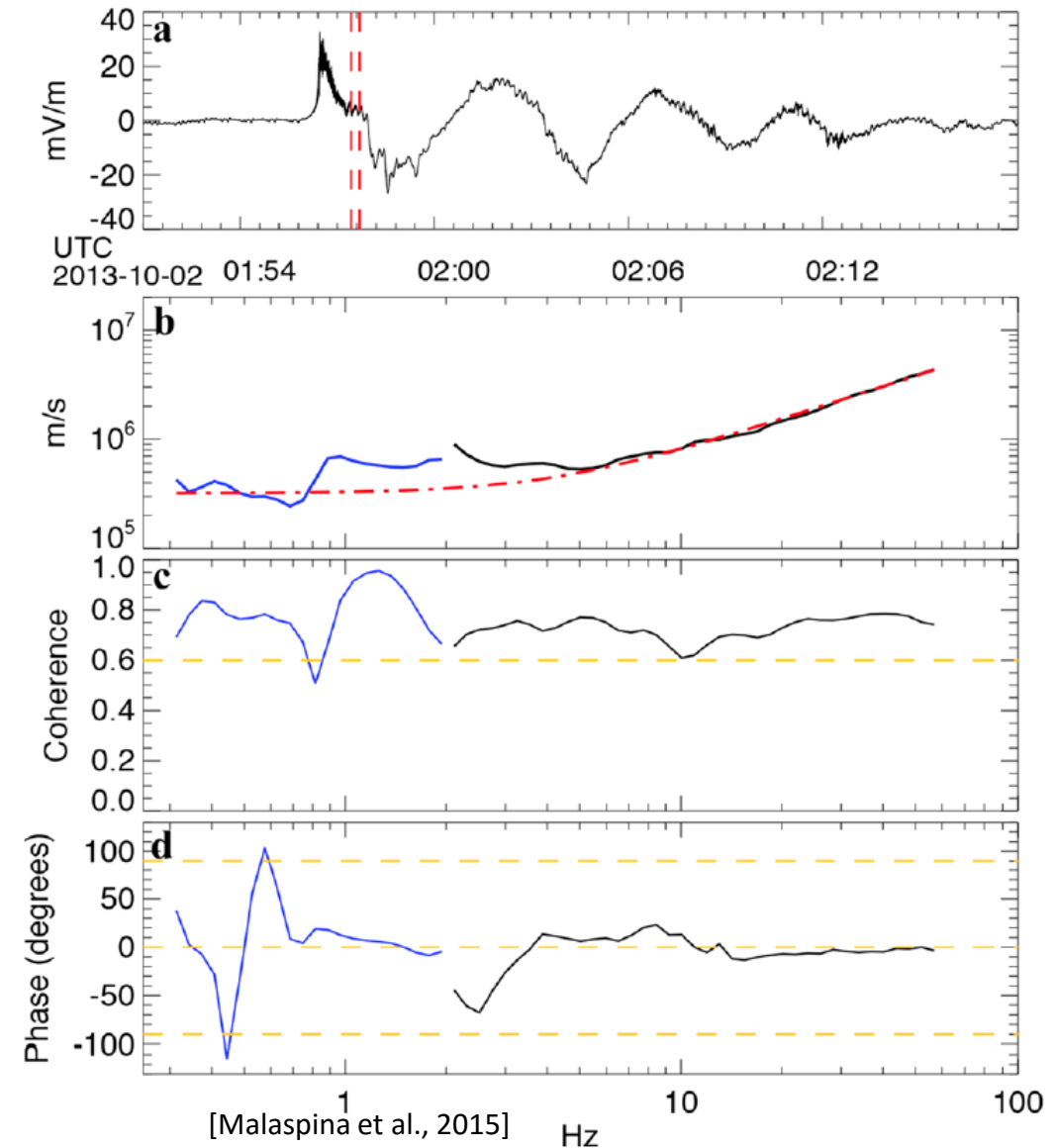


[Motoba et al., 2015]

What is the role of ULF waves in Magnetosphere-Ionosphere Coupling?

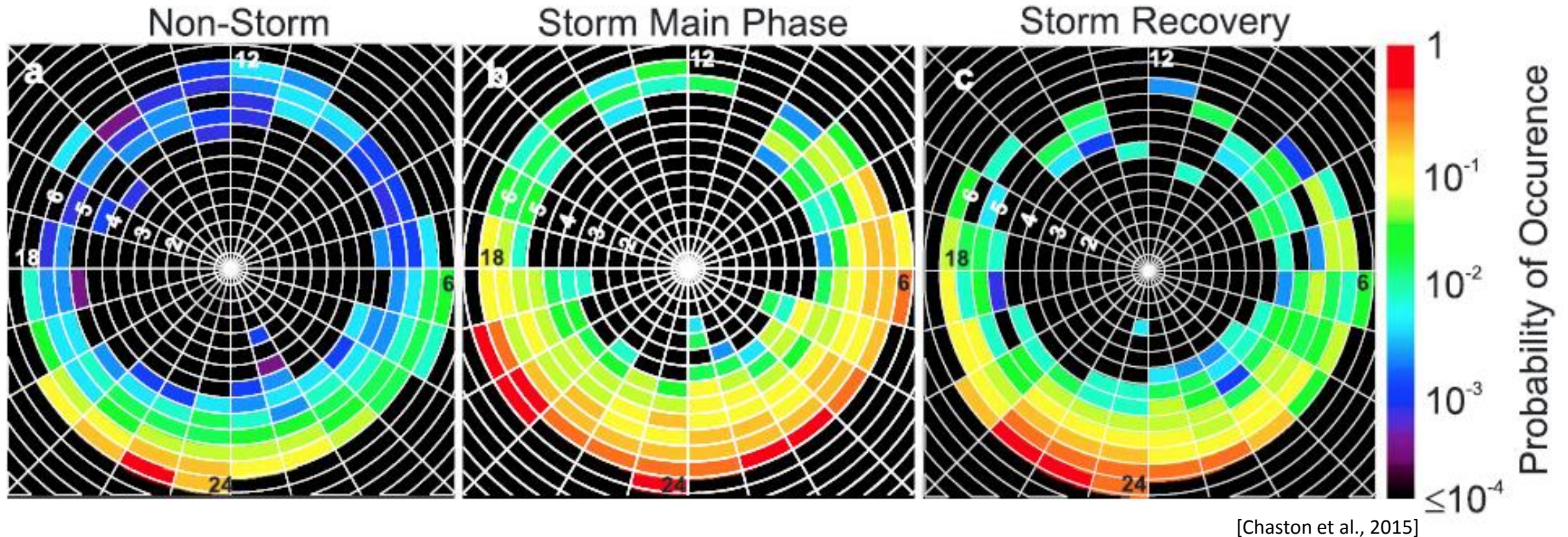
Example: Small scale, dispersive Alfvén waves (1)

- There are an increasing number of observations of small scale, dispersive Alfvén waves embedded within large wave fields – e.g., standing Alfvén waves with Pc5 frequencies
- They are associated with parallel electric fields that cause particle acceleration
- They are part of the energy budget for the large scale structure



What is the role of ULF waves in Magnetosphere-Ionosphere Coupling?

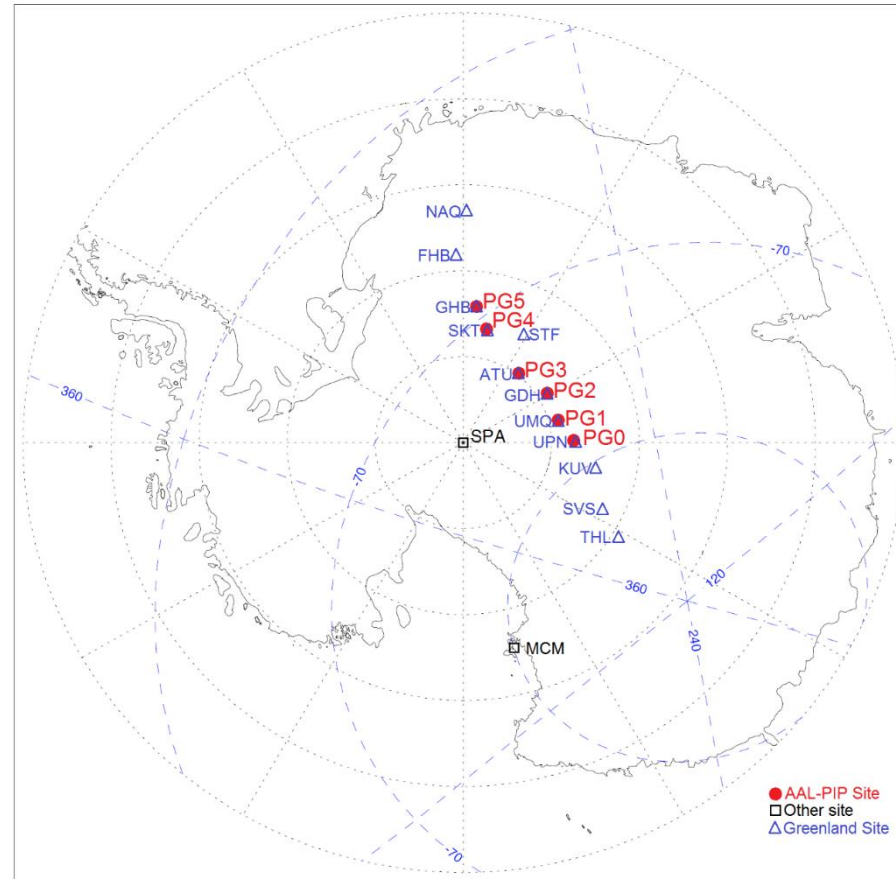
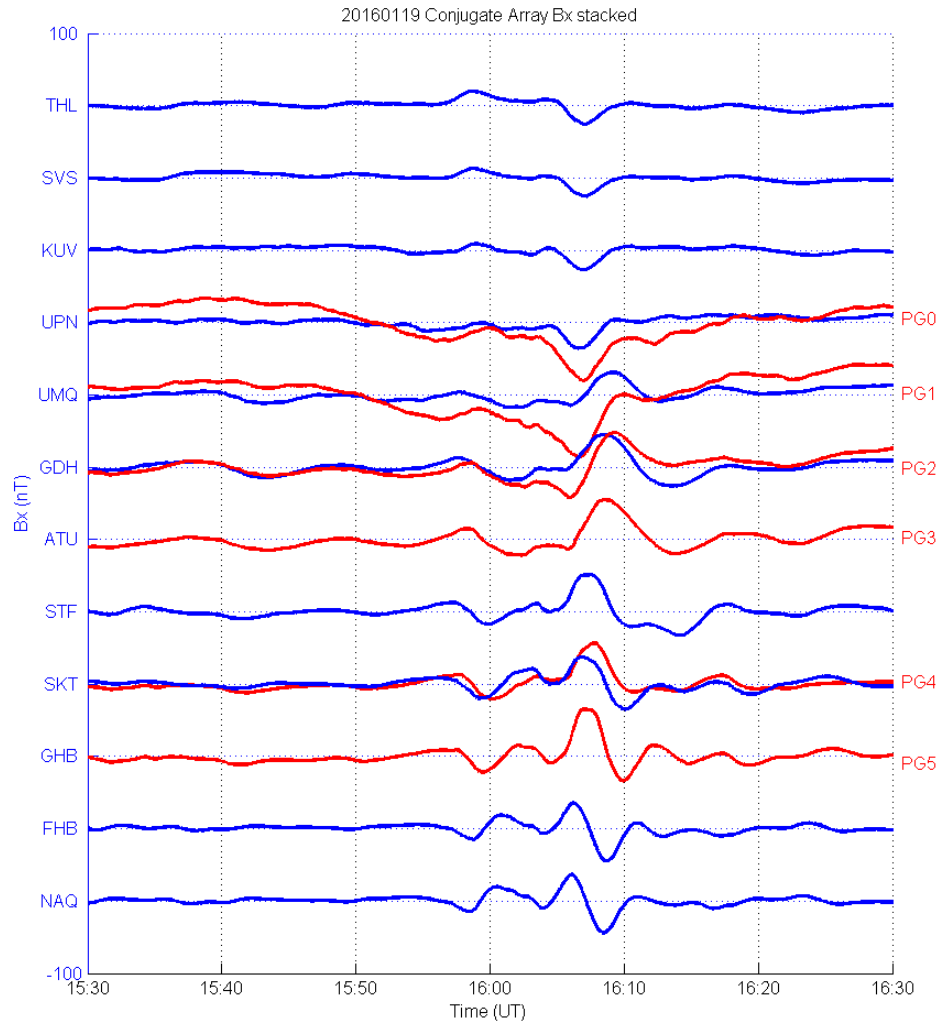
Example: Small scale, dispersive Alfvén waves (3)



- Dispersive Alfvén waves are increasingly observed in the inner magnetosphere and may also affect higher energy radiation belt particles

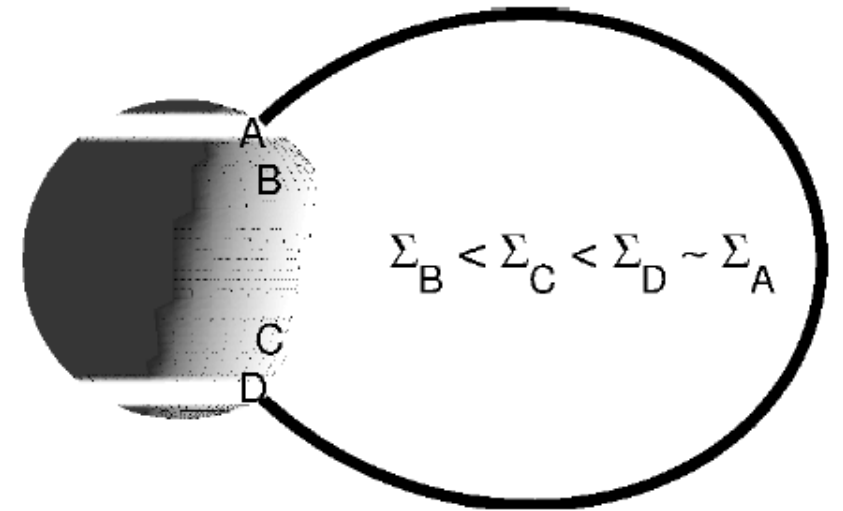
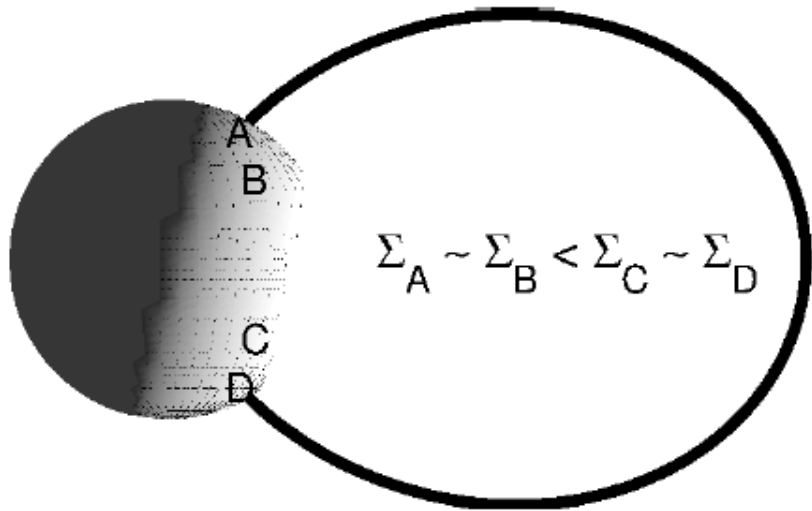
What is the role of ULF waves in Magnetosphere-Ionosphere Coupling?

Example: Explaining North-south asymmetries



- ULF wave properties often differ between the north-south hemisphere at latitudes corresponding to the auroral oval/open-closed field line boundary

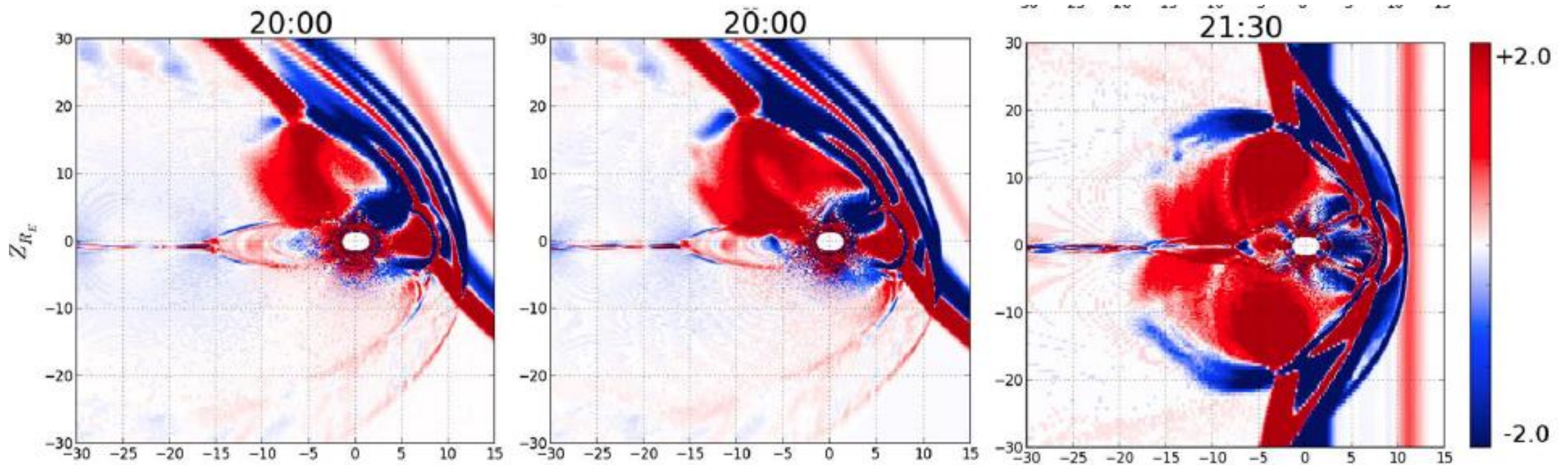
What is the role of ULF waves in Magnetosphere-Ionosphere Coupling? Example: Explaining North-south asymmetries



- Internal asymmetries affect ULF wave properties
- Large scale Alfvén waves should vary according to the ionospheric conductivity and conductivity gradients

What is the role of ULF waves in Magnetosphere-Ionosphere Coupling?

Example: Explaining North-south asymmetries



[Oliveira and Raeder, 2014, 2015]

- External asymmetries affect ULF wave properties
- The solar wind, ion foreshock, and magnetosheath energy sources for ULF wave excitation are not necessarily symmetric with respect to the magnetic equatorial plane
- Need more observations in southern hemisphere to explain north-south asymmetries

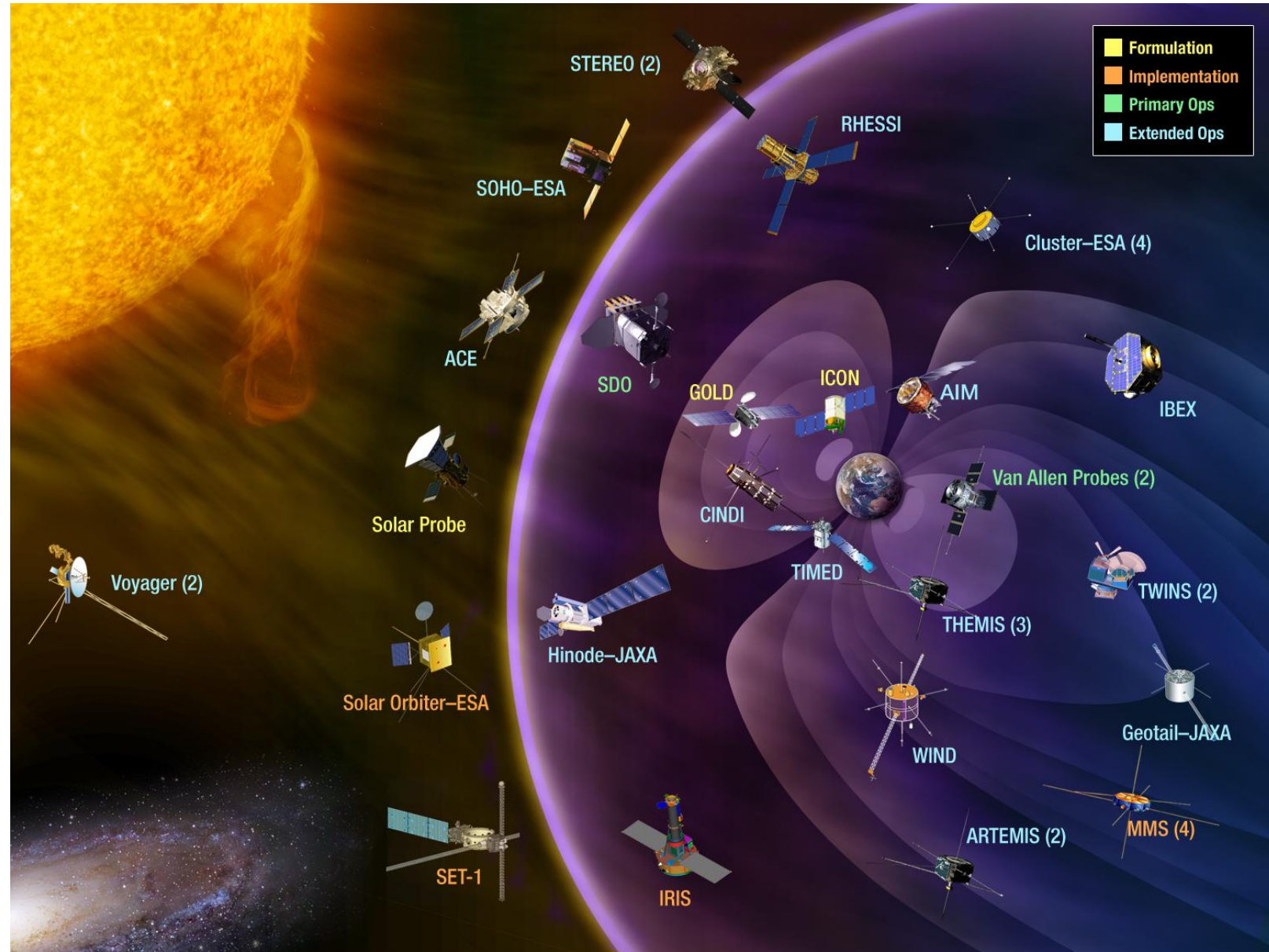
Summary of a few recent studies and new questions

- What excites ULF waves?
 - Magnetopause surface waves
 - Drift/drift-bounce resonance
- How do ULF waves couple to the ring current?
 - Drift/drift-bounce resonance
 - Particle injections to small radial distances
- How do ULF waves couple to the plasmasphere?
 - Standing Alfvén waves
 - EMIC waves/EMIC wave growth
- How do ULF waves couple to the radiation belts?
 - New radial diffusion coefficients
 - Non-diffusive transport/drift resonance
 - Pc5 modulation of VLF waves and EMIC waves – when/where/how do Pc5 waves affect radiation belt dynamics?
- What is the role of ULF waves in MI coupling?
 - ULF modulation of auroral precipitation
 - Relation between large scale ULF waves and small scale, dispersive Alfvén waves that can cause precipitation
 - Role of north-south asymmetries

New Observations and New Opportunities

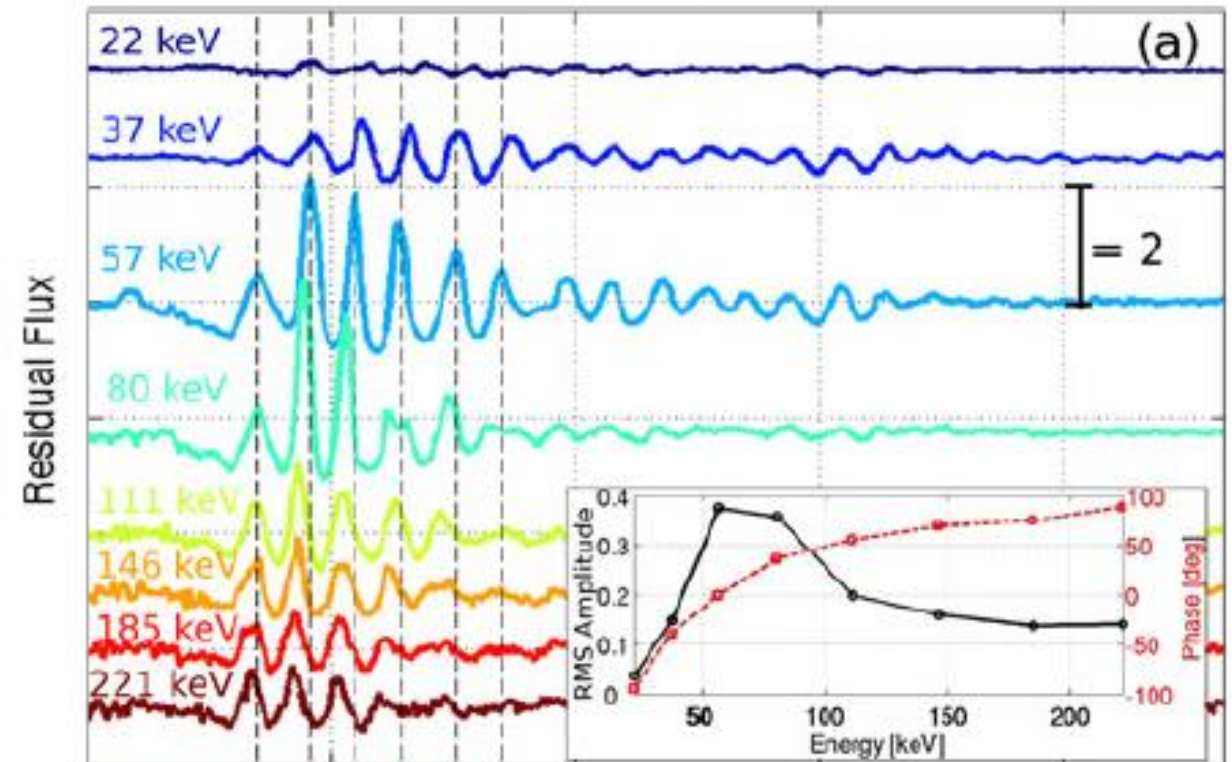
- Multiple satellite constellations
 - → Global wave properties

[NASA]



New Observations and New Opportunities

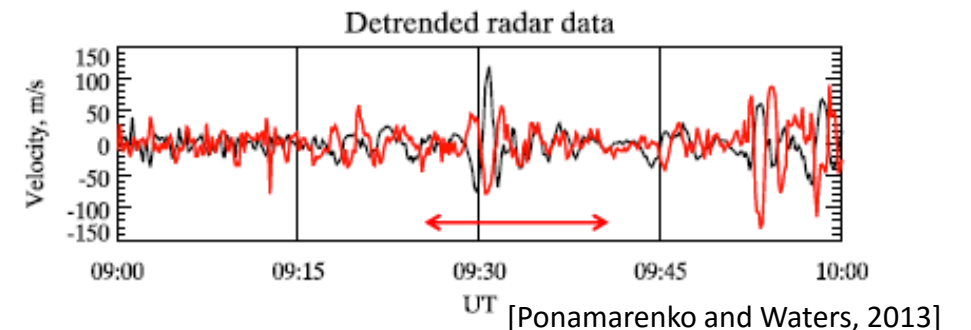
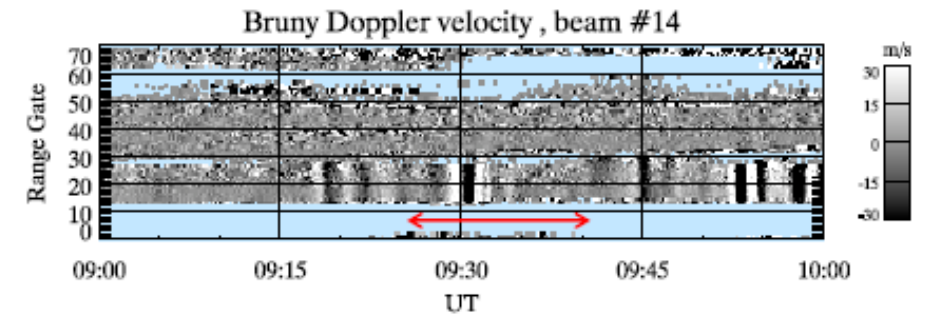
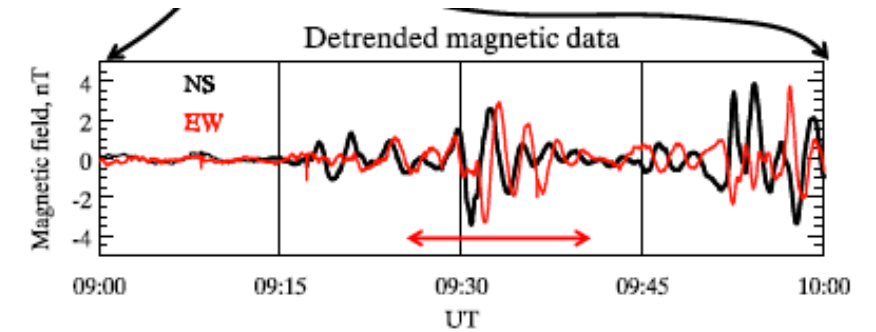
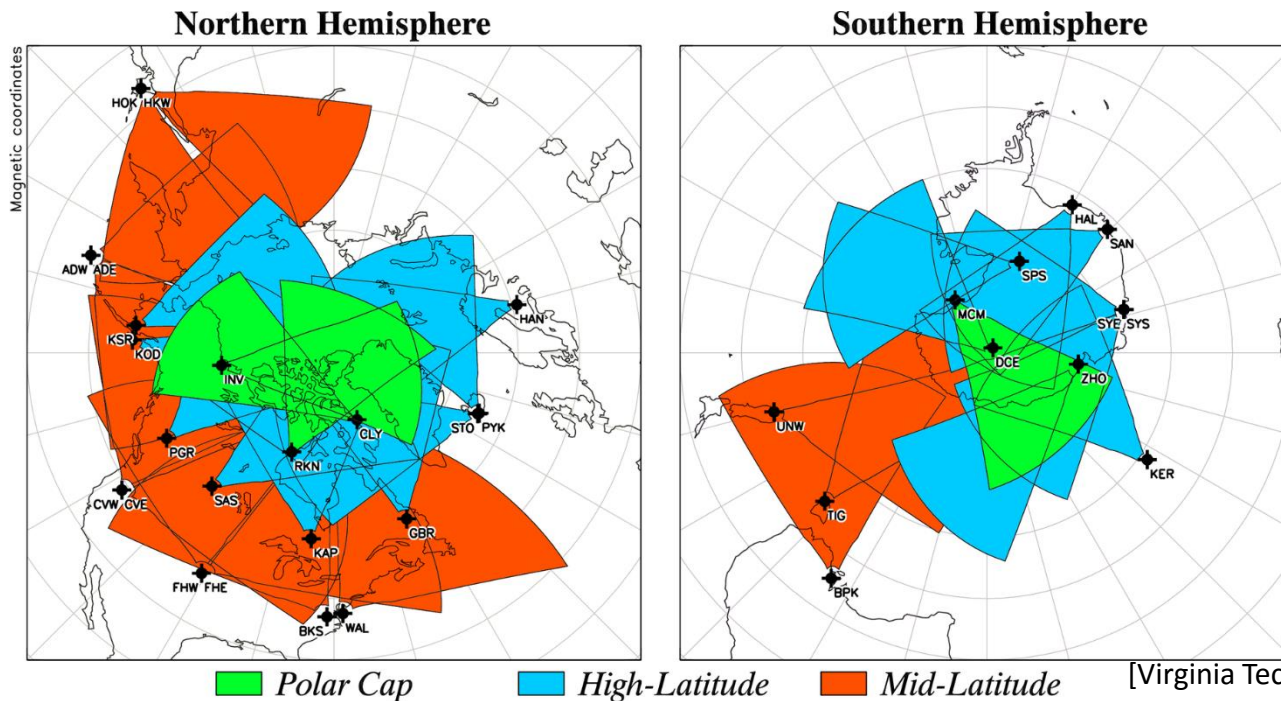
- Multiple satellite constellations
 - → Global wave properties
- High resolution particle and fields instruments
 - → Drift/drift-bounce resonance



[Claudepierre et al., 2013]

New Observations and New Opportunities

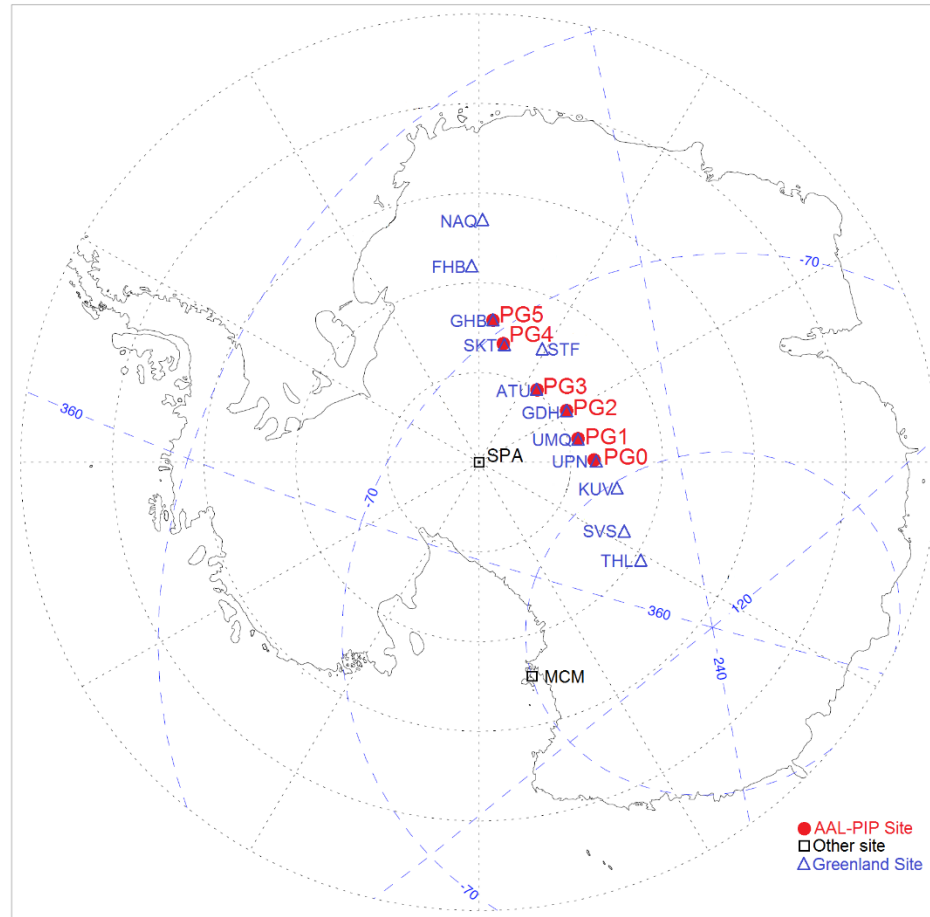
- Multiple satellite constellations
 - → Global wave properties
- High resolution particle and fields instruments
 - → Drift/drift-bounce resonance
- Expanded ground-based arrays
 - → New information about ULF wave excitation and dissipation



[Ponamarenko and Waters, 2013]

New Observations and New Opportunities

- Multiple satellite constellations
 - → Global wave properties
- High resolution particle and fields instruments
 - → Drift/drift-bounce resonance
- Expanded ground-based arrays
 - → New information about ULF wave excitation and dissipation
- Coordinated instrument operation and data analysis
 - → North-south asymmetries, global wave properties and their impact on the ring current/radiation belts



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Introducing the UMEA focus group: ULF wave modeling, effects, and applications

- ULF wave research spans many different research areas and focus groups
- UMEA aims to bring researchers in different areas together to address these questions:
 - What excites ULF waves?
 - How do ULF waves couple to the plasmasphere/ring current/radiation belts?
 - What is the role of ULF waves in Magnetosphere-Ionosphere coupling?
- This is an ideal time for such an effort: unprecedented availability of multi-point in situ and ground-based observations, high quality measurements of electric/magnetic fields and particles, and improved modeling capabilities

“...the subject of hydromagnetic waves in the magnetosphere (or magnetospheres) is highly developed...”
but

“...the subject is far from played out as a research field...”
Southwood and Hughes, [1983]

Thank you!