

# System Understanding of Radiation Belt Particle Dynamics through Multi-Spacecraft and Ground-Based Observations and Modeling

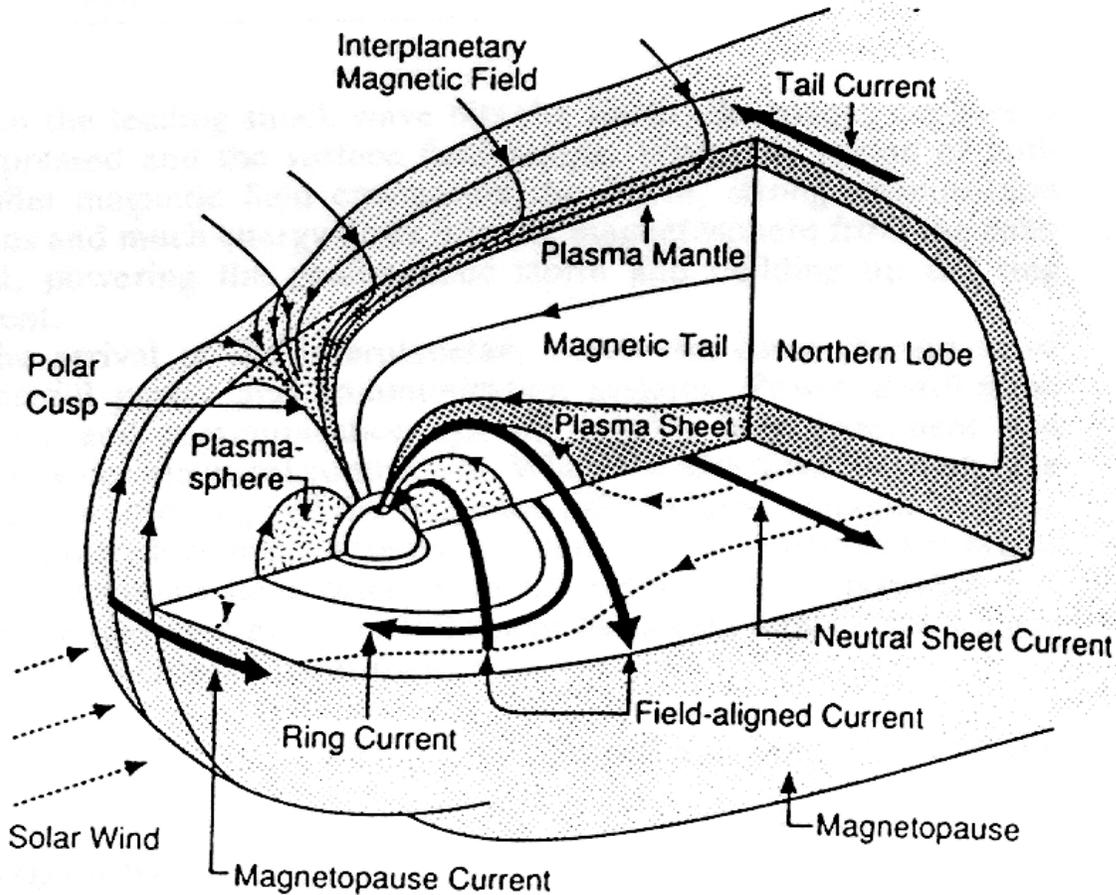
Hong Zhao

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# Outline

- Introduction of Radiation Belt Particle Dynamics
- Recent Advances in Radiation Belt Studies based on Van Allen Probes observations
- System Understanding of Radiation Belt Particle Dynamics through Multi-Spacecraft and Ground-Based Observations and Modeling
- Summary

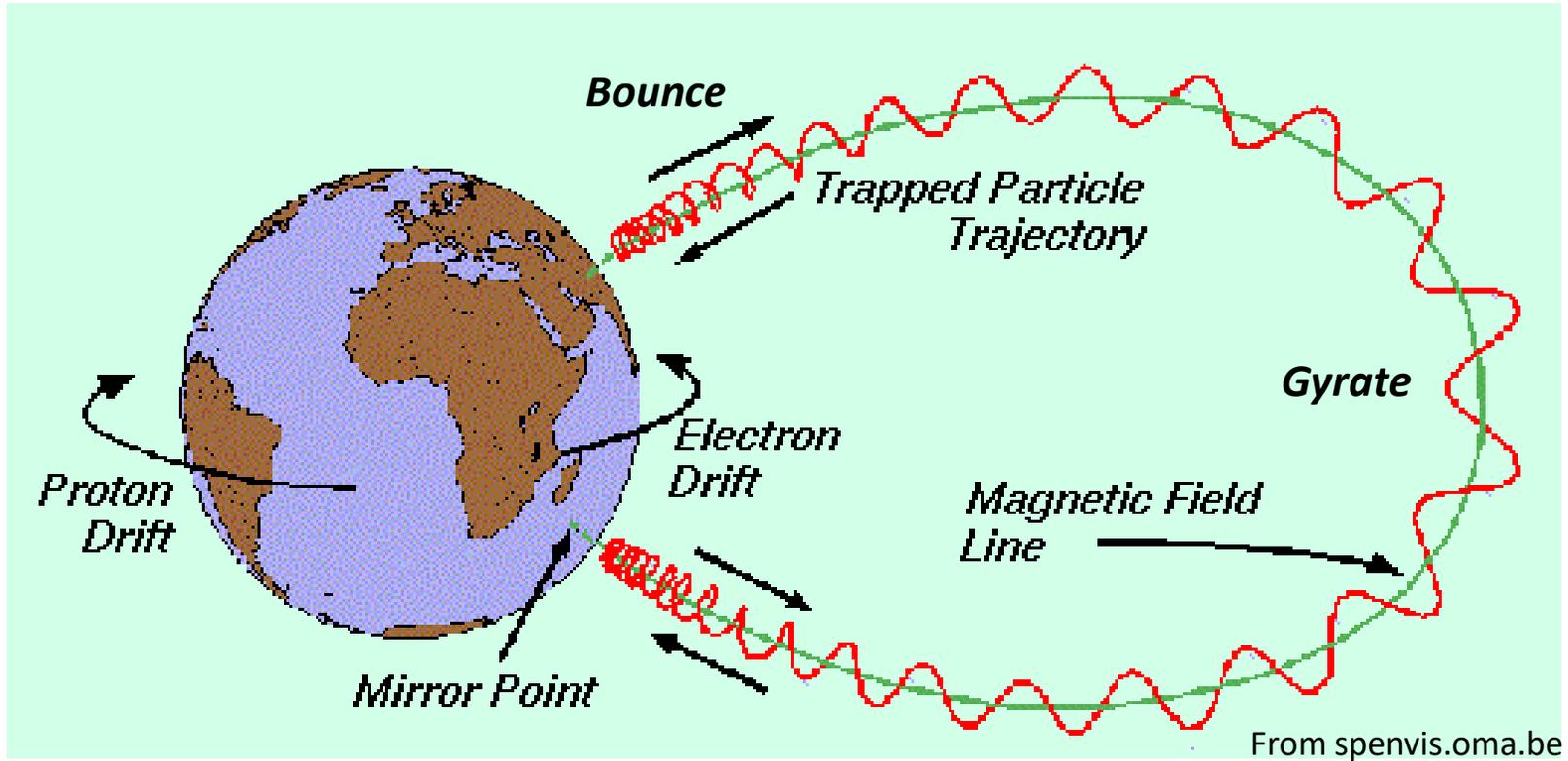
# Earth's Magnetosphere



From Kivelson and Russell [1995]

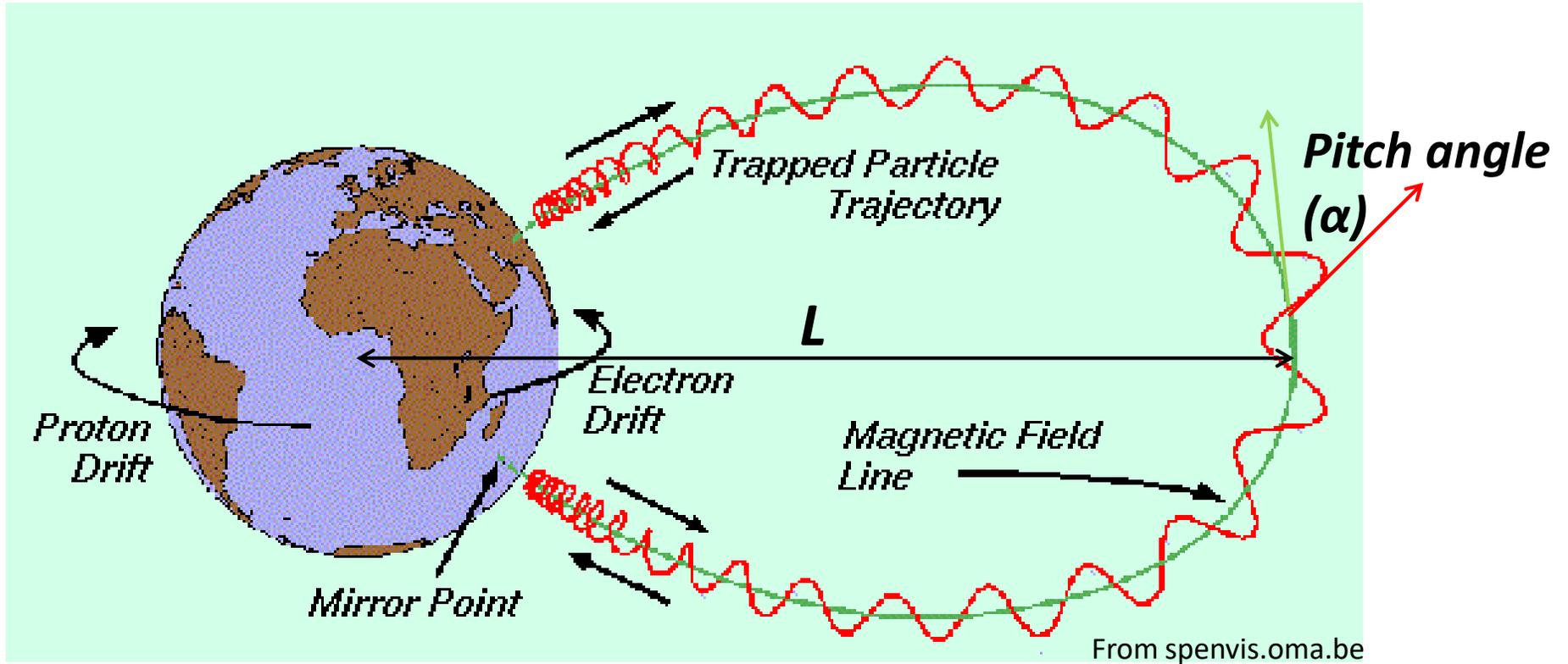
- Earth's magnetosphere is a tear-shaped region carved out of the solar wind by Earth's magnetic field.
- Close to Earth, the Earth's inner magnetosphere is the region where the geomagnetic field resembles the dipole field. It includes the plasmasphere, radiation belts, and ring current.
- The charged particles and current systems existing in the Earth's inner magnetosphere pose potential threats to the spacecraft and human in the space and technical systems on the ground.

# Charged Particle Motions



Time scales (GEO, E=1MeV, $\alpha=60^\circ$ )	Gyro	Bounce	Drift
Electrons	1 ms	0.5 s	10 min
Protons	0.6 s	10 s	7 min

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# Adiabatic Invariants

- First adiabatic invariant (corresponding to gyration):

$$\mu = \frac{P_{\perp}^2}{2m_0B} = \text{const.}$$

- Second adiabatic invariant (corresponding to bounce motion):

$$J = 2\sqrt{2m_0\mu} \int_{S_m}^{S'_m} \sqrt{B_m - B(s)} ds = \text{const.}$$

$$\text{(Or } K = \int_{S_m}^{S'_m} \sqrt{B_m - B(s)} ds = \text{const.)}$$

- Third adiabatic invariant (corresponding to drift motion):

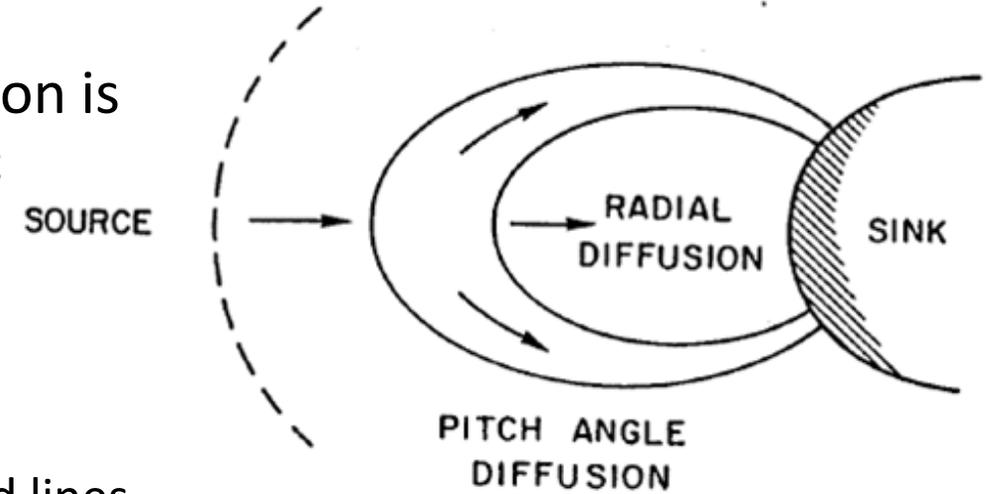
$$\Phi = \int_S \vec{B} \cdot \vec{dS} = \text{const.} ; L^* = \frac{2\pi M}{|\Phi|R_e}$$

- Phase-averaged phase space coordinates:  $(\mu, K, L^*)$

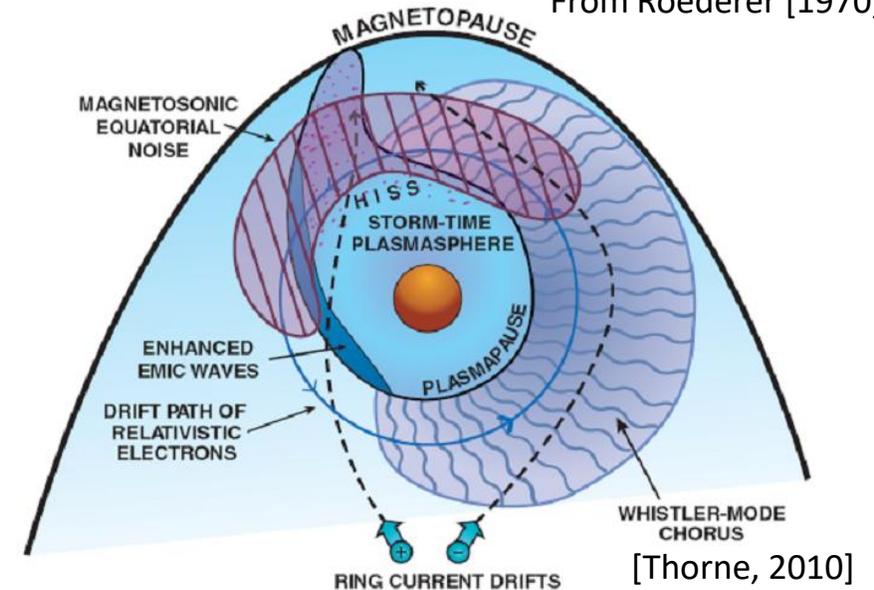
- Phase space density:  $f_p = \frac{j}{p^2}$

# Violations of the Adiabatic Invariants

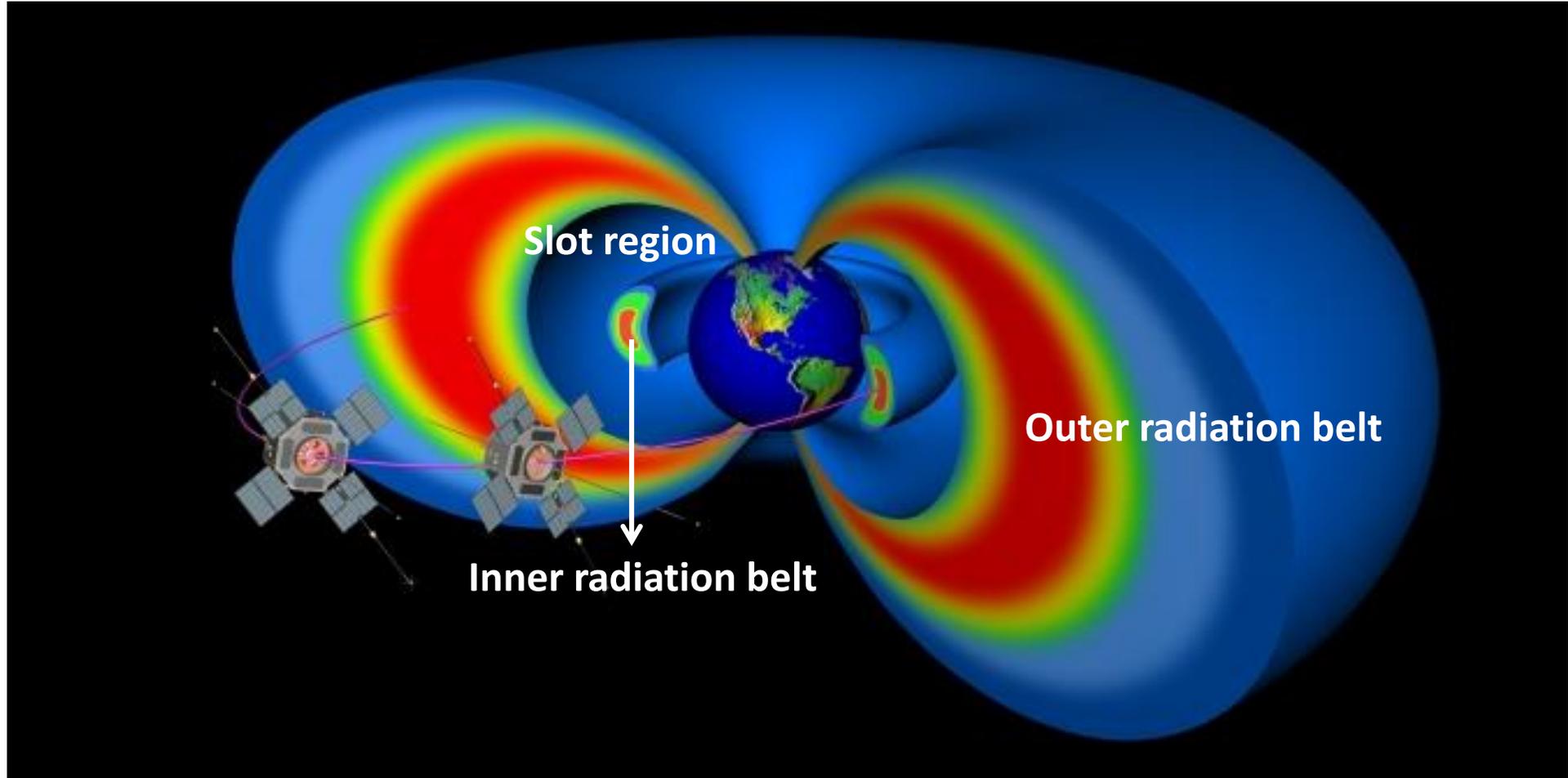
- When the timescale of magnetic/electric field fluctuation is shorter compared to the particle motion, the adiabatic invariant(s) can be violated.
- Violations of the adiabatic invariants -> diffusion
  - Radial diffusion: permits transport of the particles across field lines.
  - Pitch angle diffusion: alters the particle pitch angle.
  - Energy diffusion: changes the particle energy.
- Magnetospheric waves can cause the diffusion processes.
  - ULF waves, chorus waves, hiss waves...



From Roederer [1970]



# The Earth's Radiation Belts

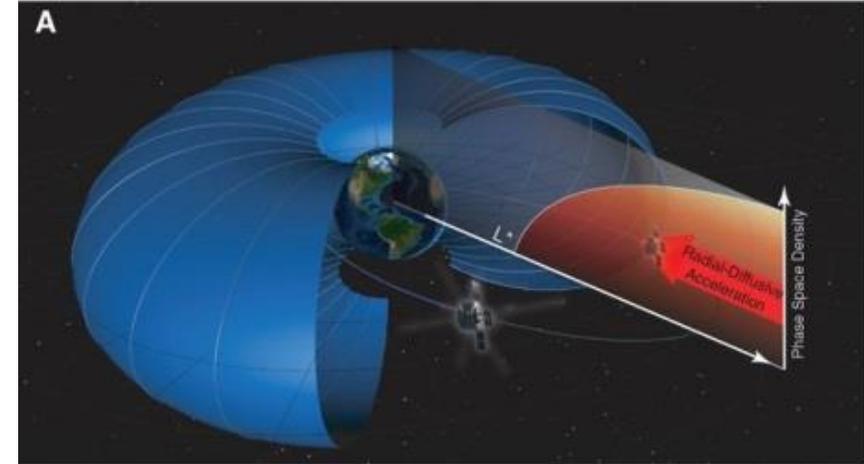


[From nasa.gov]

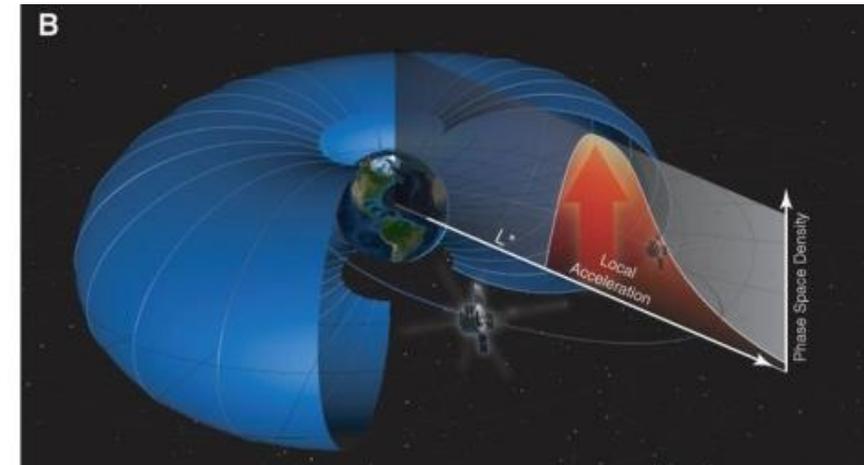
# Source and Loss Processes for Radiation Belt Electrons

- Source processes
  - Inward radial diffusion
  - Local acceleration
- Loss processes
  - Magnetopause shadowing
  - Precipitation into the atmosphere
  - Outward radial diffusion

Inward radial diffusion



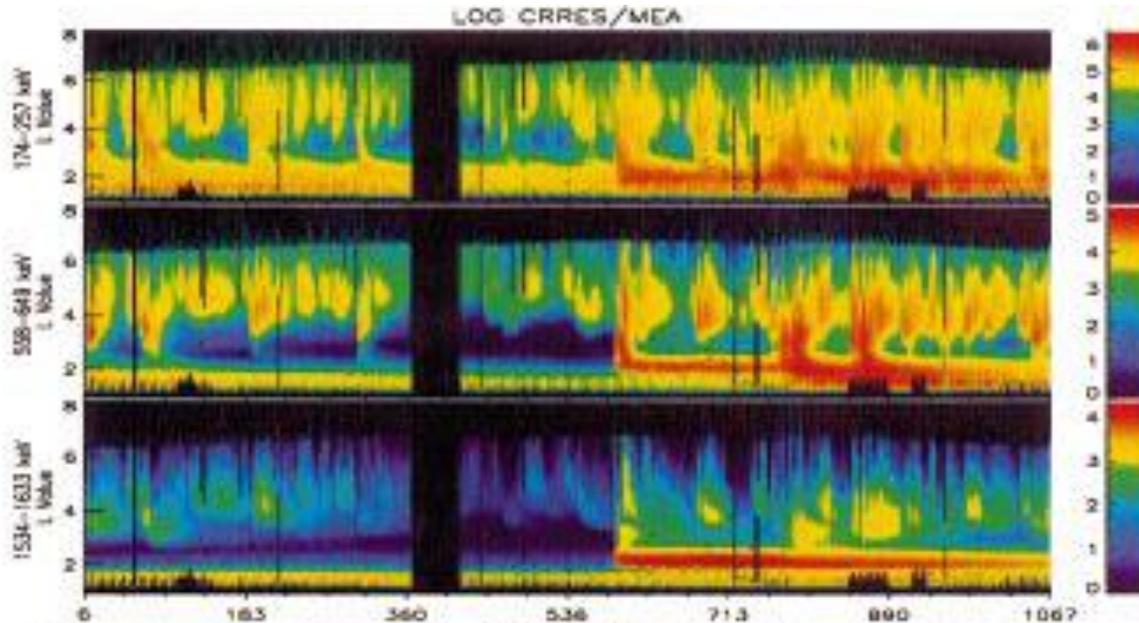
Local acceleration



(Reeves et al., 2013)

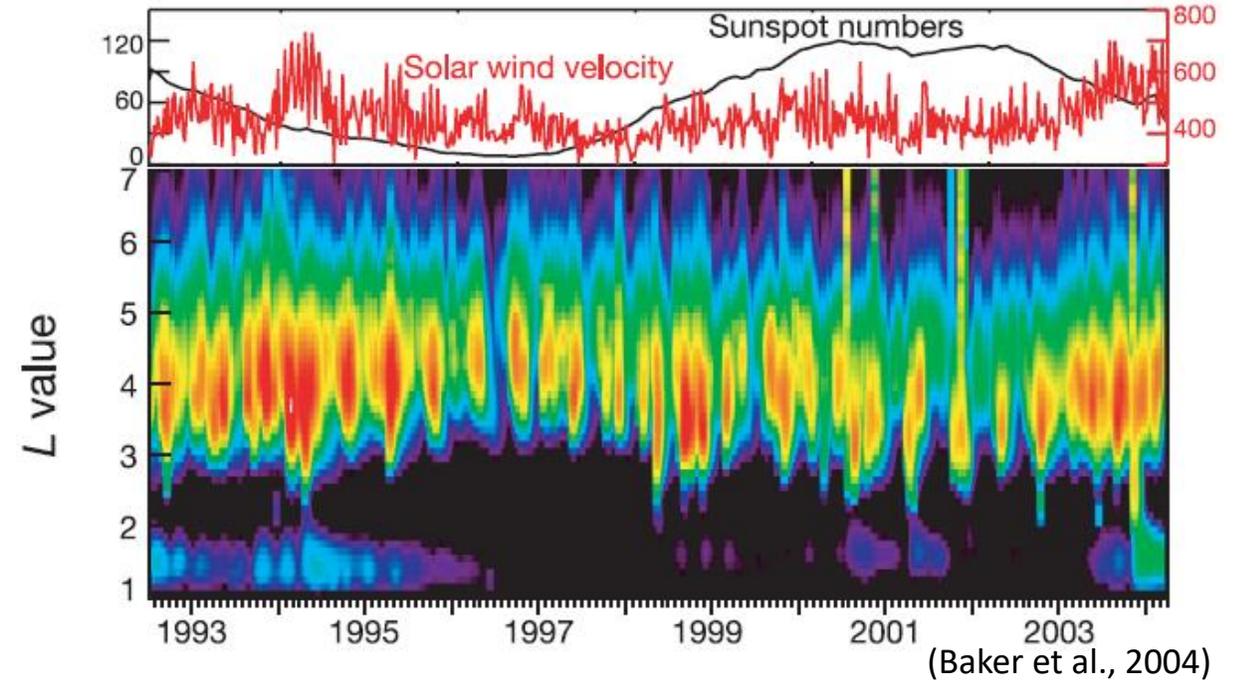
# Long-Term Variations of Radiation Belt Electrons: Measurements Prior to the Van Allen Probes Era

CRRES



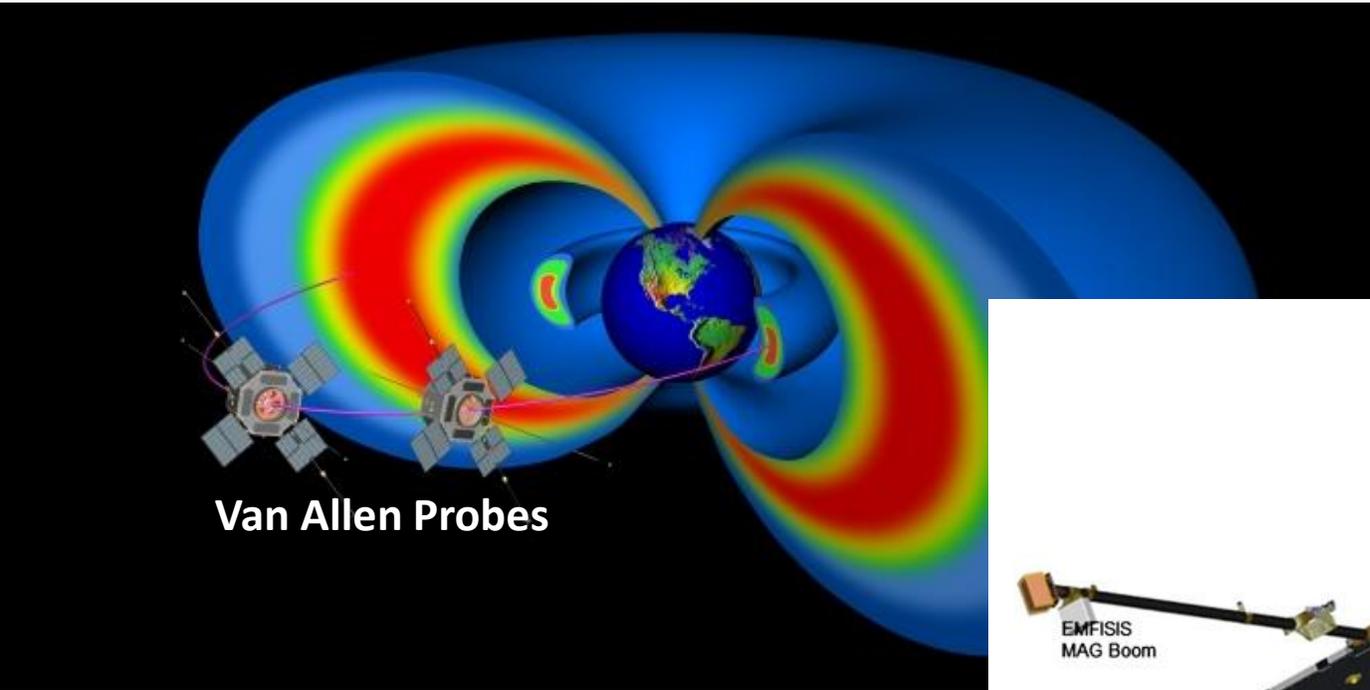
(Li and Temerin, 2001)

SAMPEX



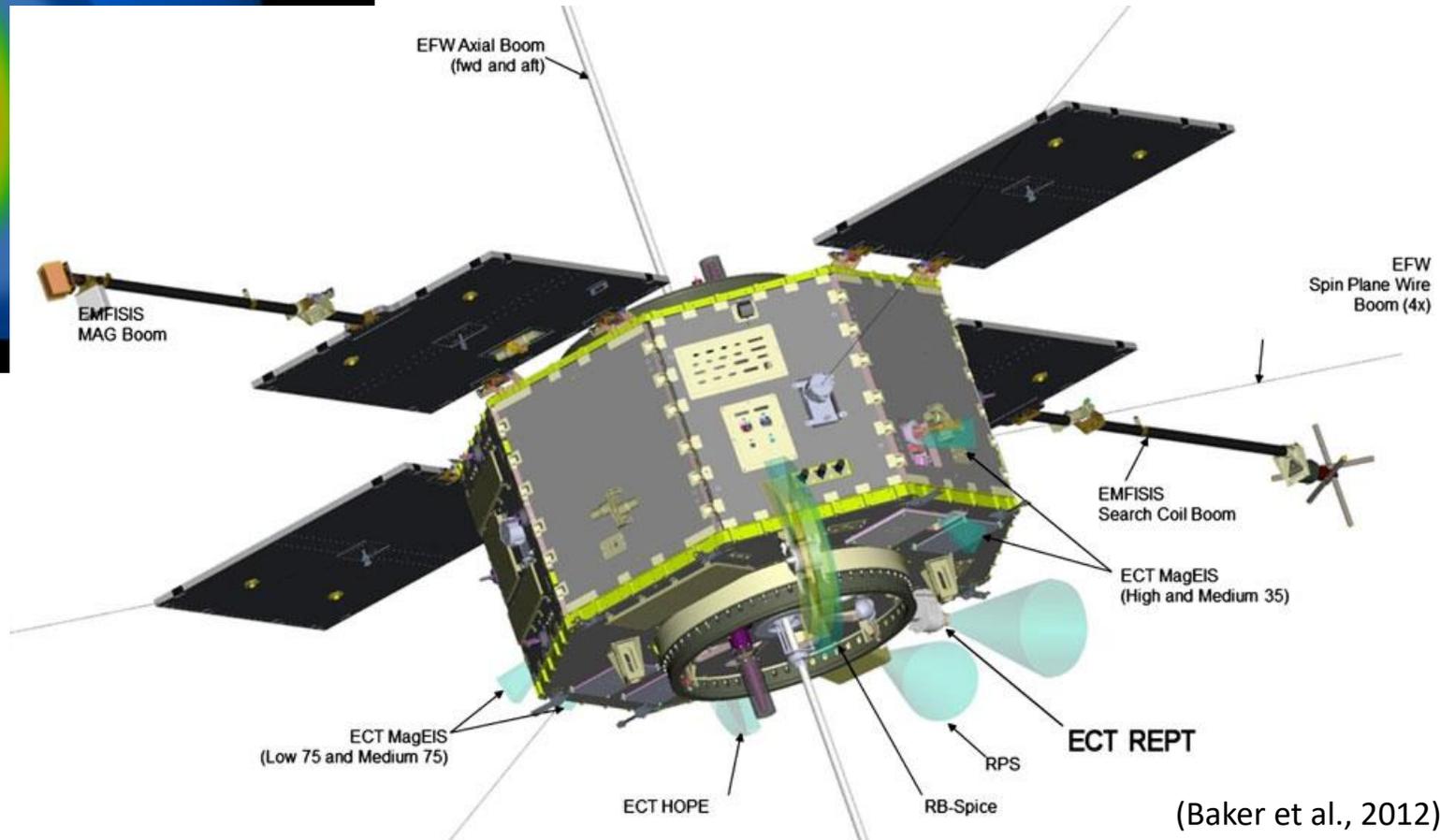
- The outer radiation belt is very dynamic and radiation belt particles are subject to significant influence from the solar wind.

# The Van Allen Probes



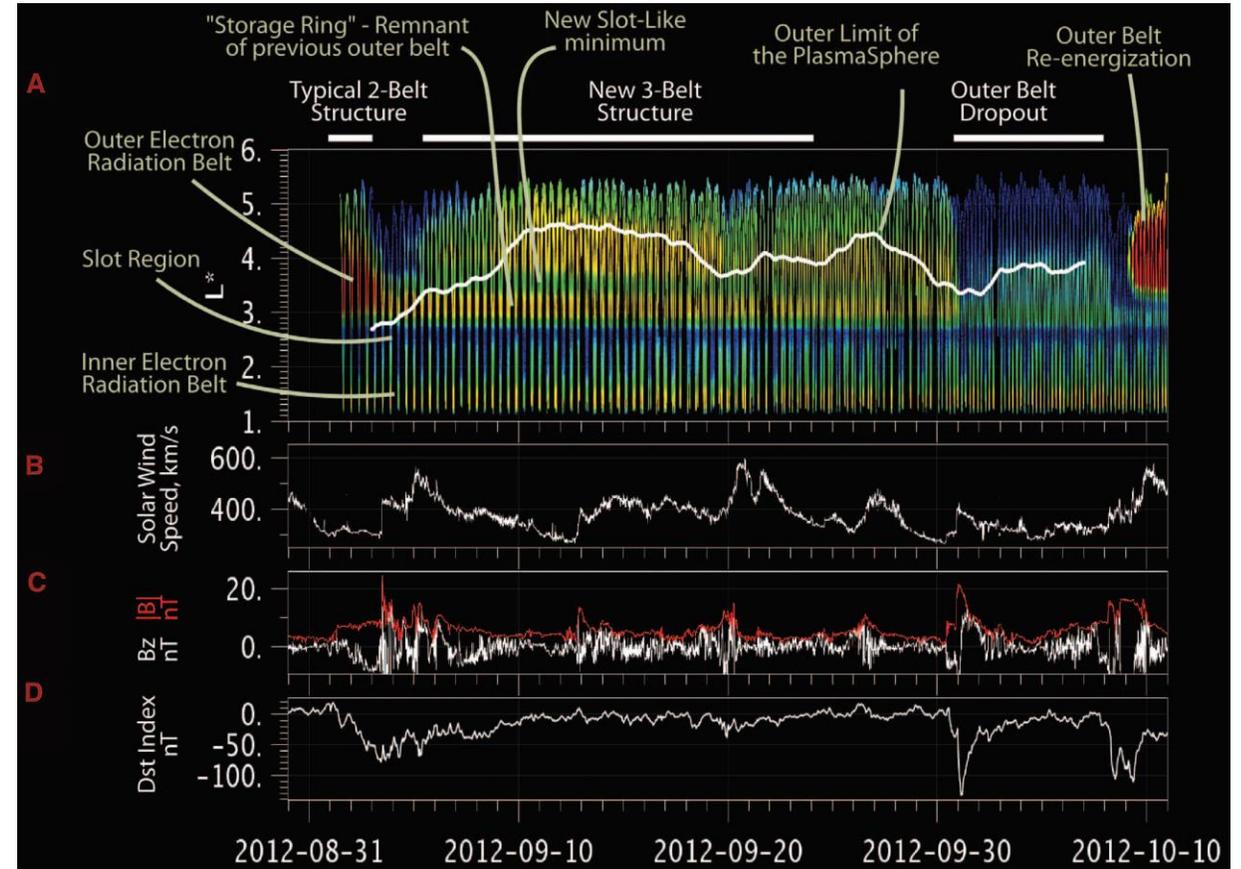
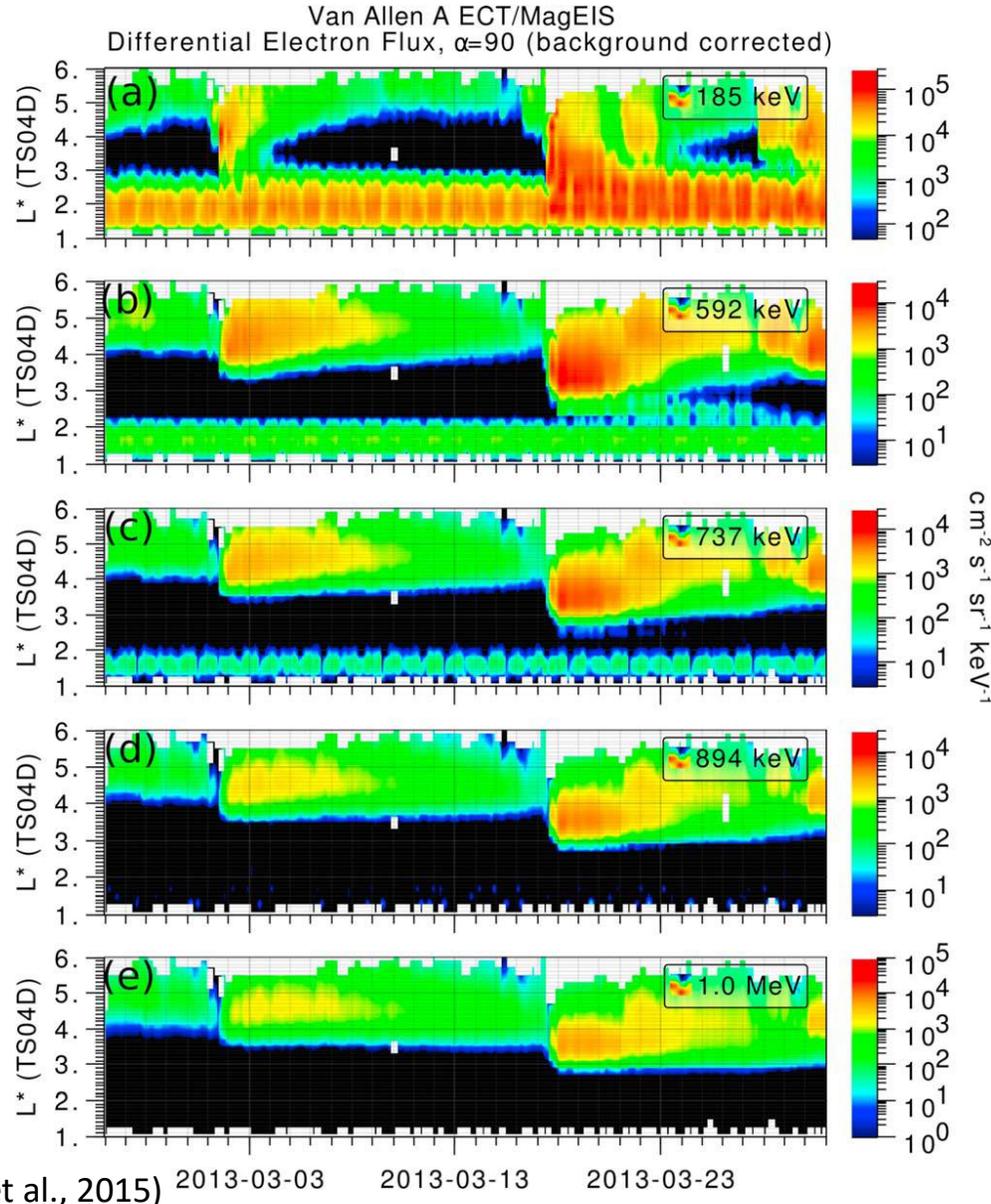
Van Allen Probes

[From nasa.gov]

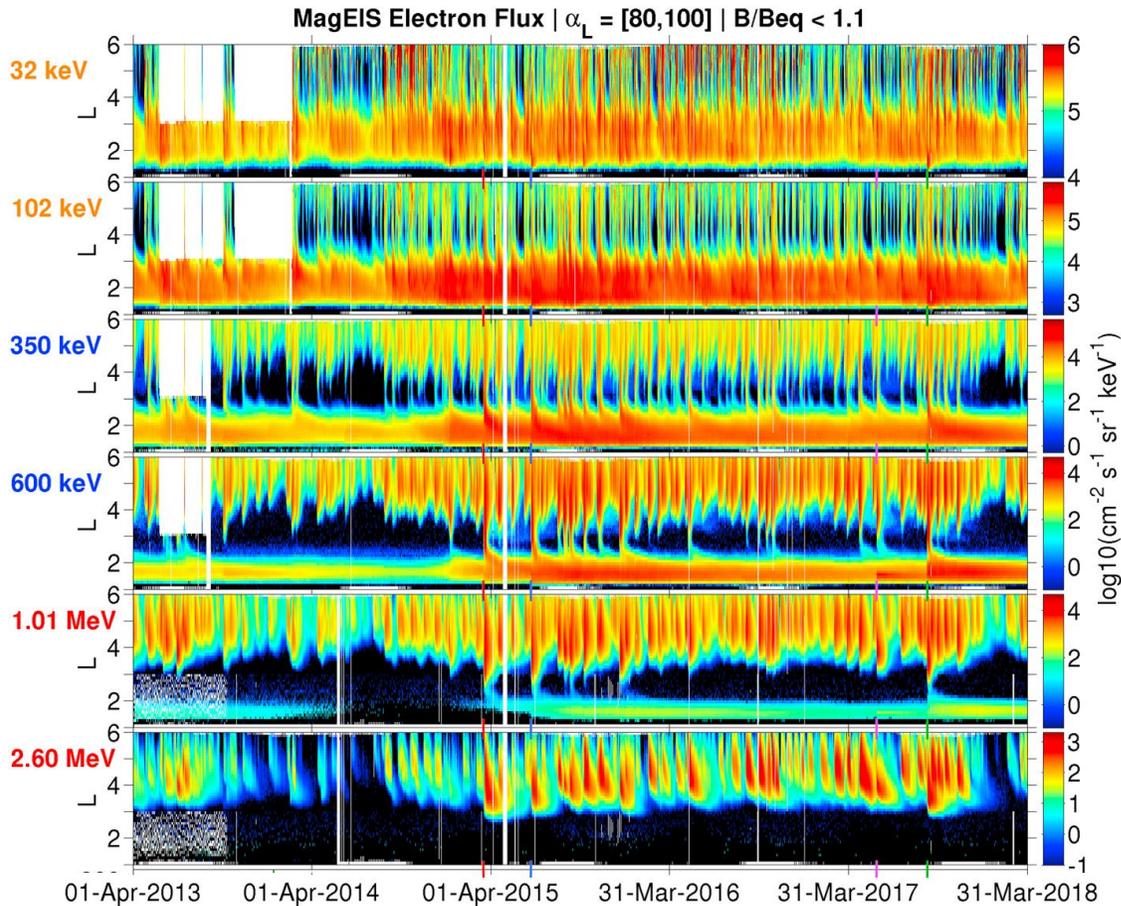


(Baker et al., 2012)

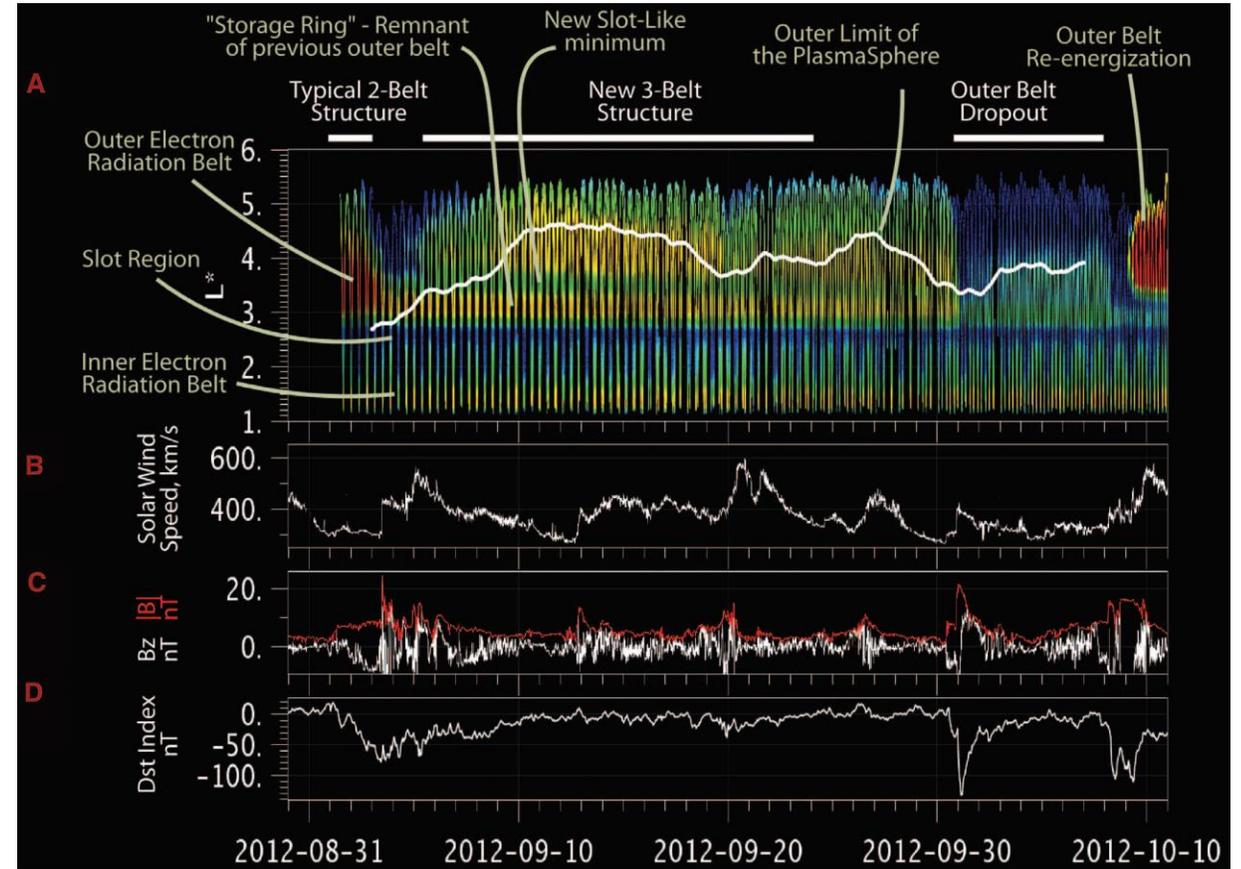
# Radiation Belt Electron Measurements of the Van Allen Probes



- Abundant 100s of keV electrons but limited >MeV electrons exist in the inner radiation belt;
- Three-belt structure of radiation belts is found.

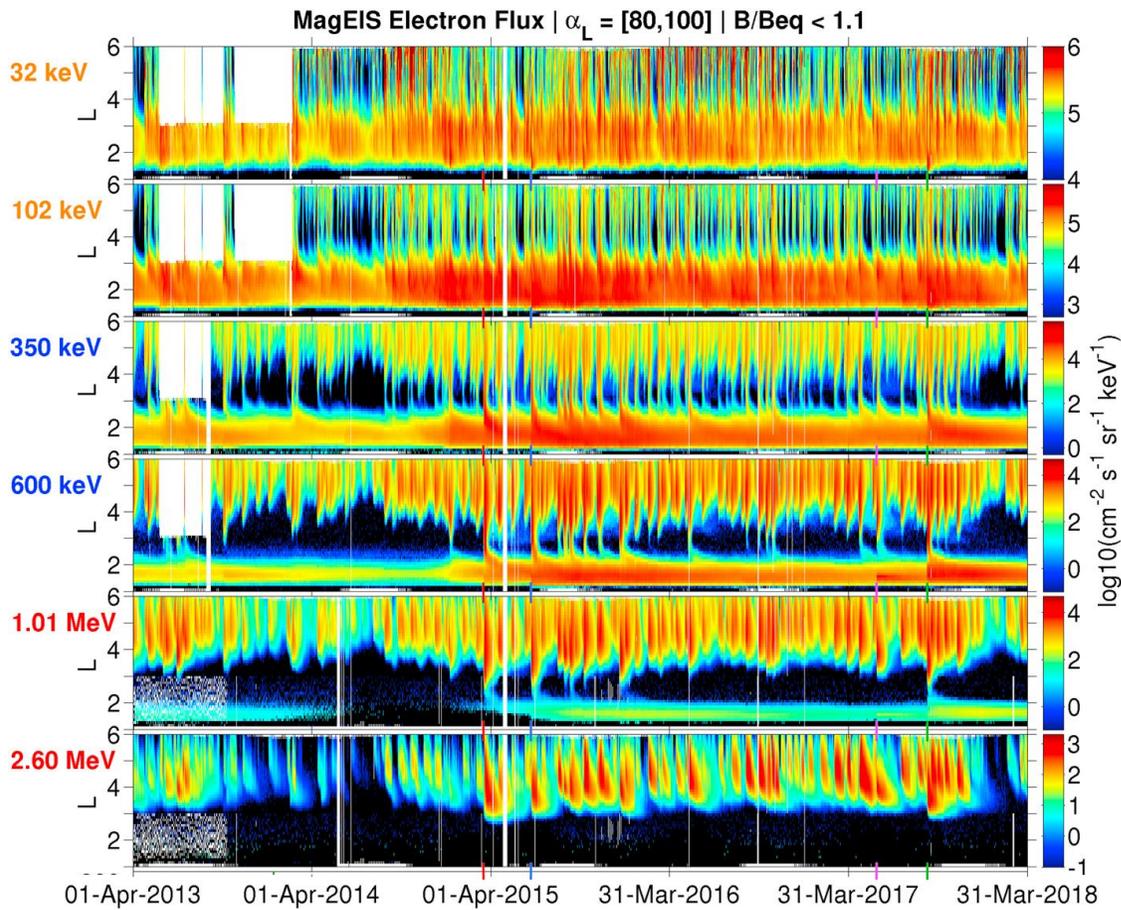


(Claudepierre et al., 2019)

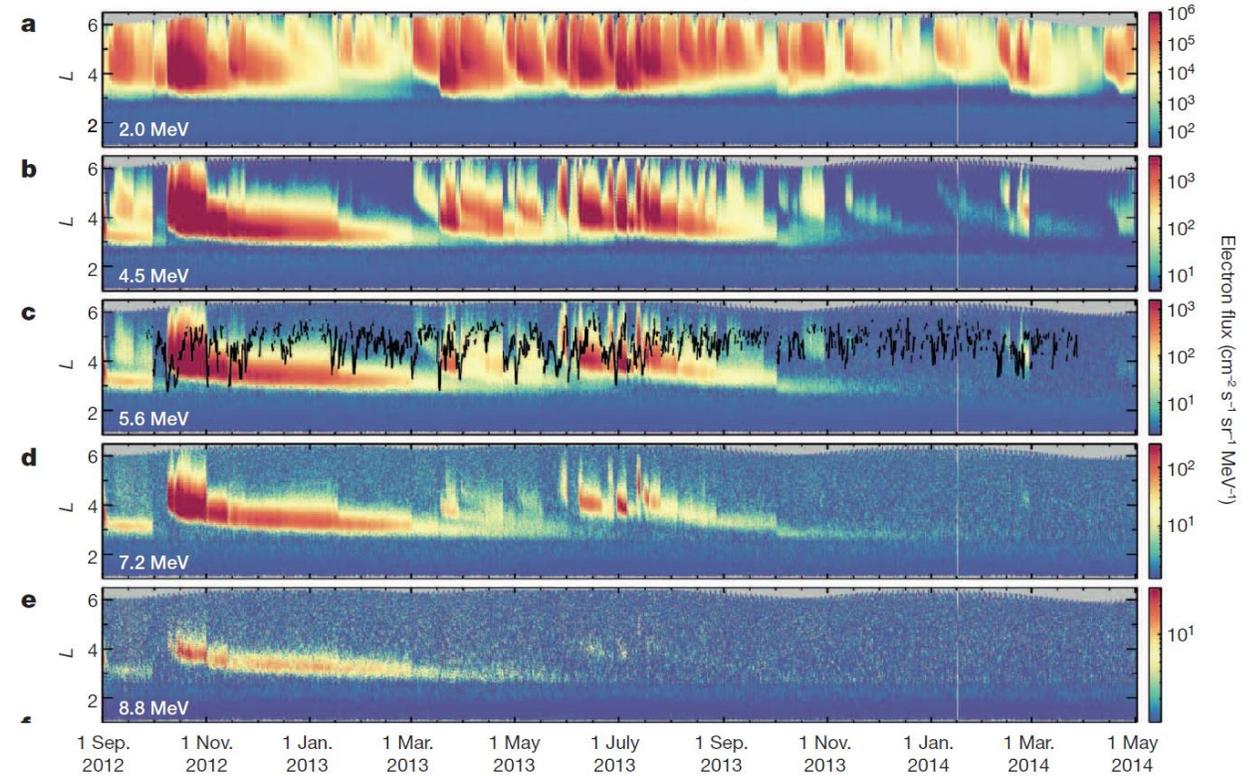


(Baker et al., 2013)

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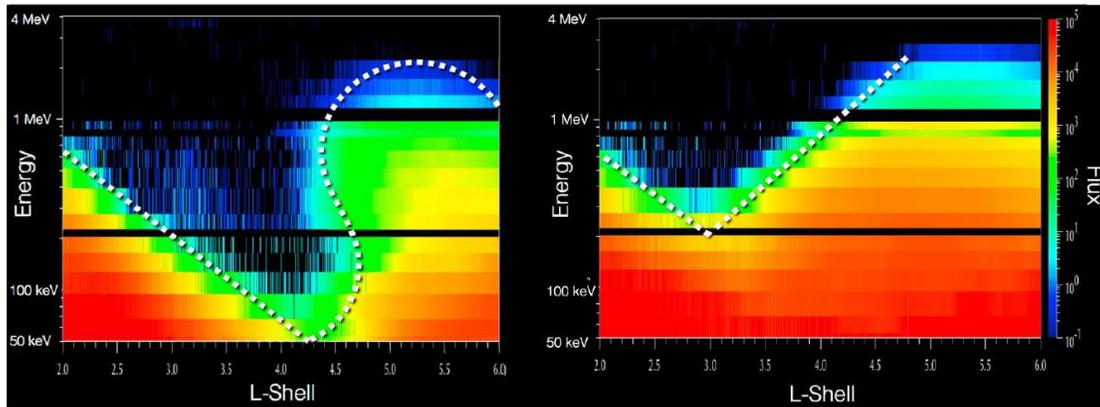
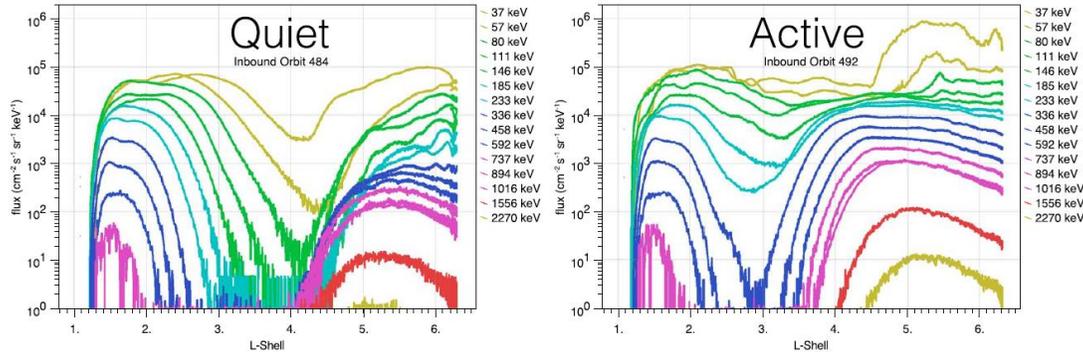
(Claudepierre et al., 2019)



(Baker et al., 2014)

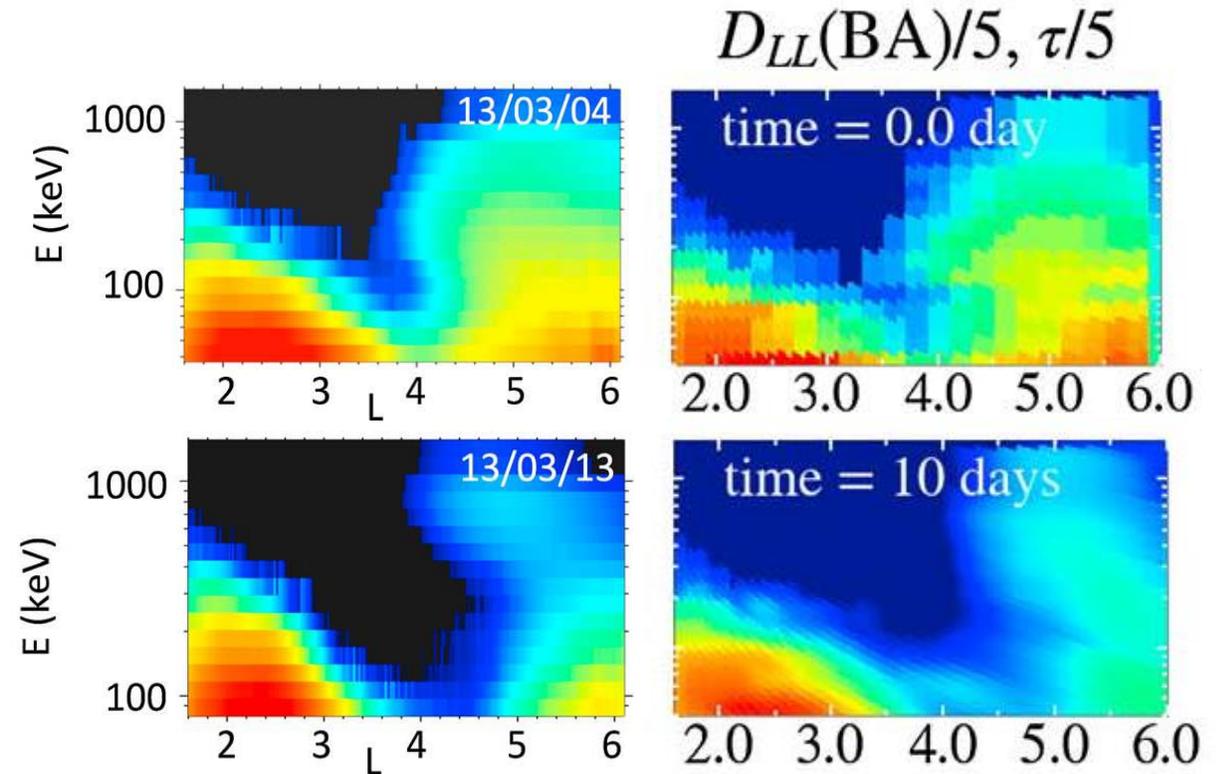
- Impenetrable barrier for multi-MeV electrons exists.

# L- and Energy-Dependent Features of Radiation Belt Electrons



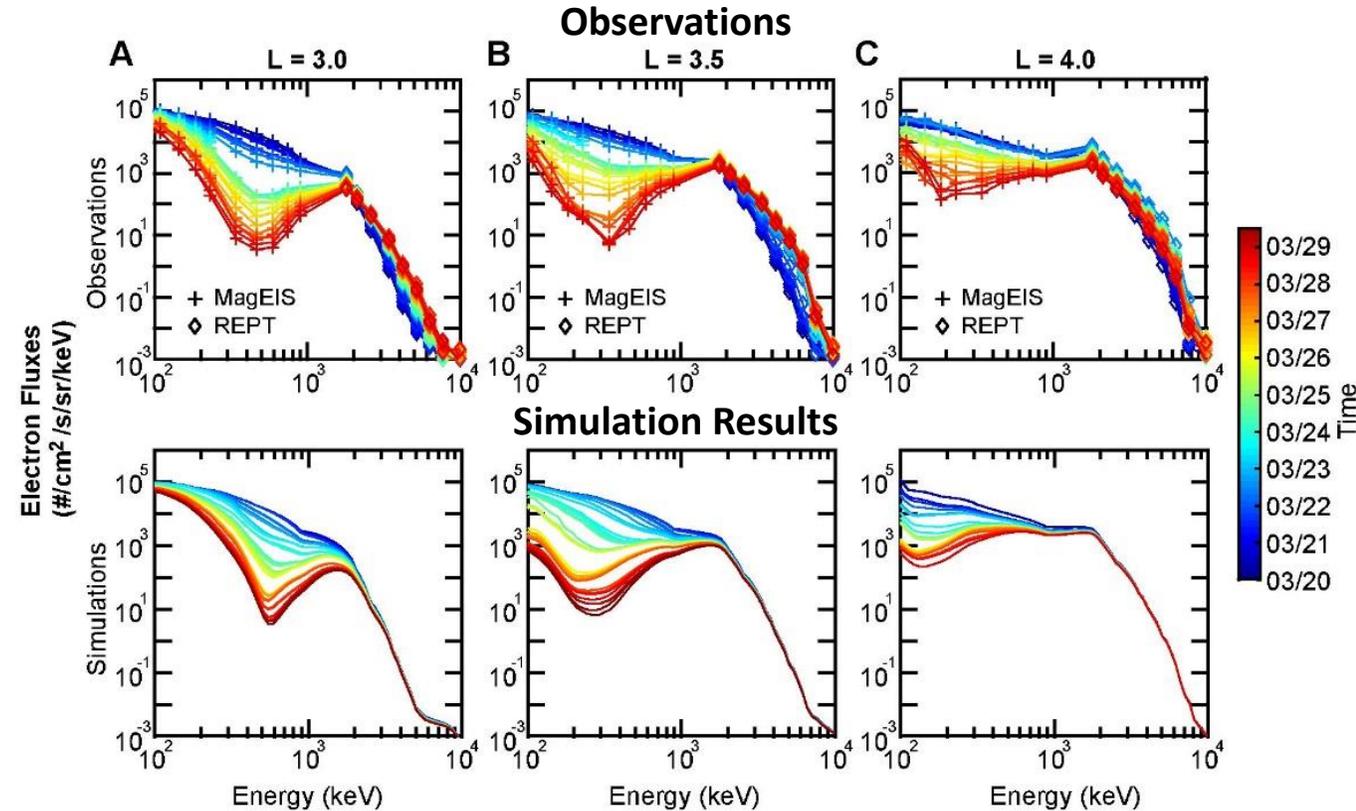
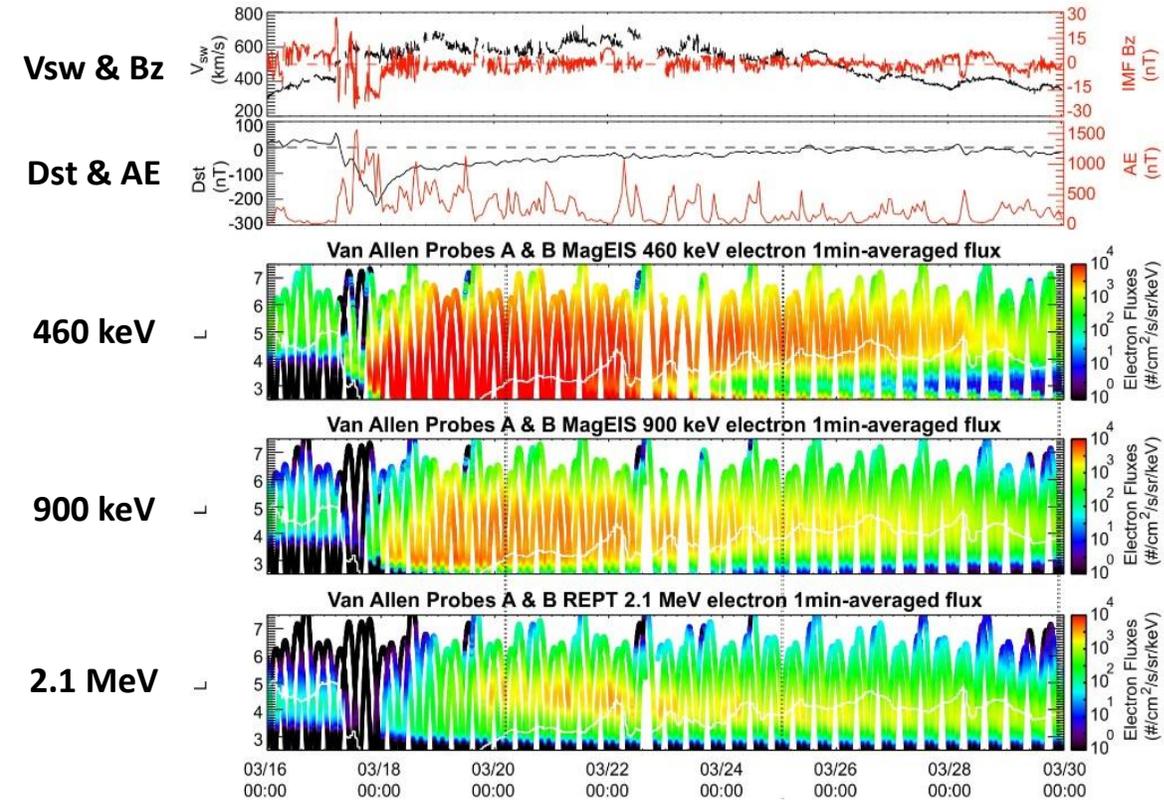
(Reeves et al., 2016)

- S-shaped structure in energy-L distribution is generally present during quiet times and disappears during active times;
- S-shaped structure is formed as a result of plasmaspheric hiss wave scattering.



(Ripoll et al., 2016)

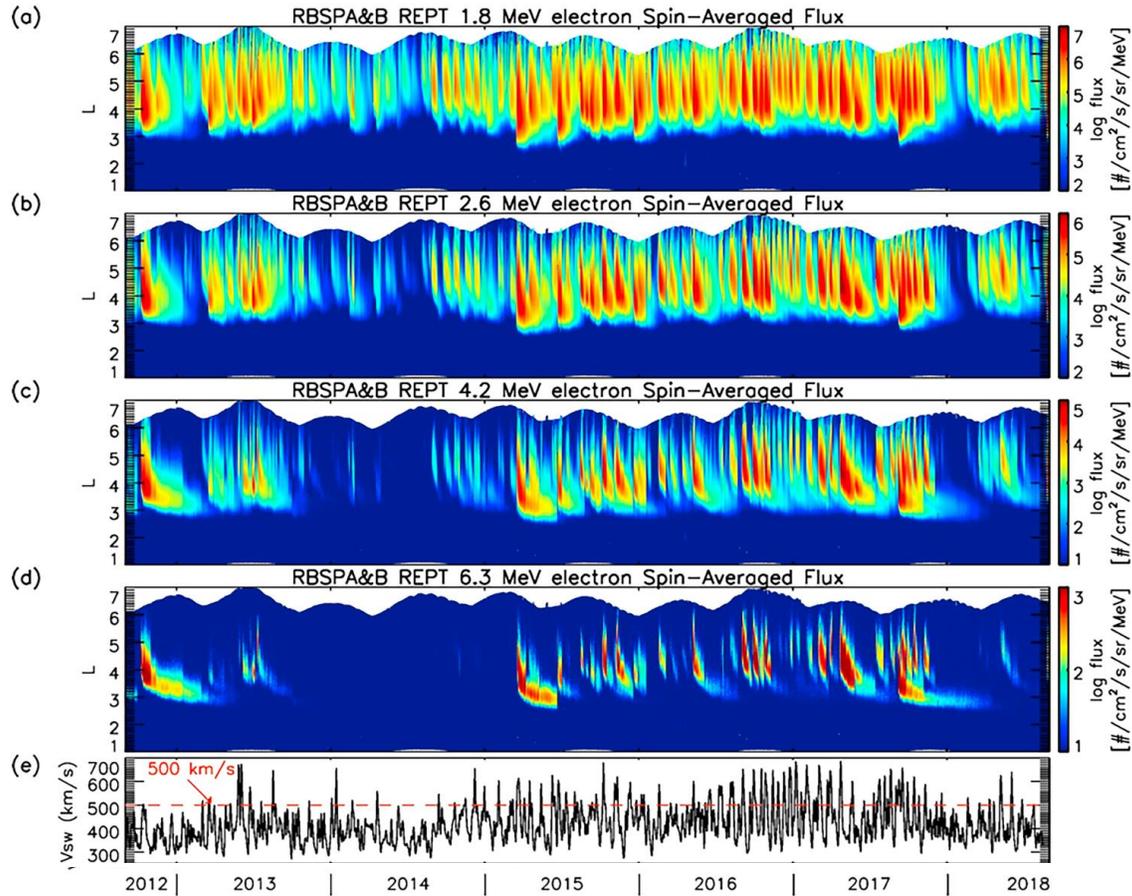
# Bump-On-Tail (BOT) Energy Spectrum of Radiation Belt Electrons



(Zhao et al., 2019)

- Reversed energy spectrum of ~100s of keV – 2 MeV electrons is reported;
- BOT energy spectrum is actually the most prevalent energy spectrum inside the plasmasphere at  $L > \sim 2.6$ ;
- Plasmaspheric hiss wave scattering is responsible for the formation of BOT energy spectrum.

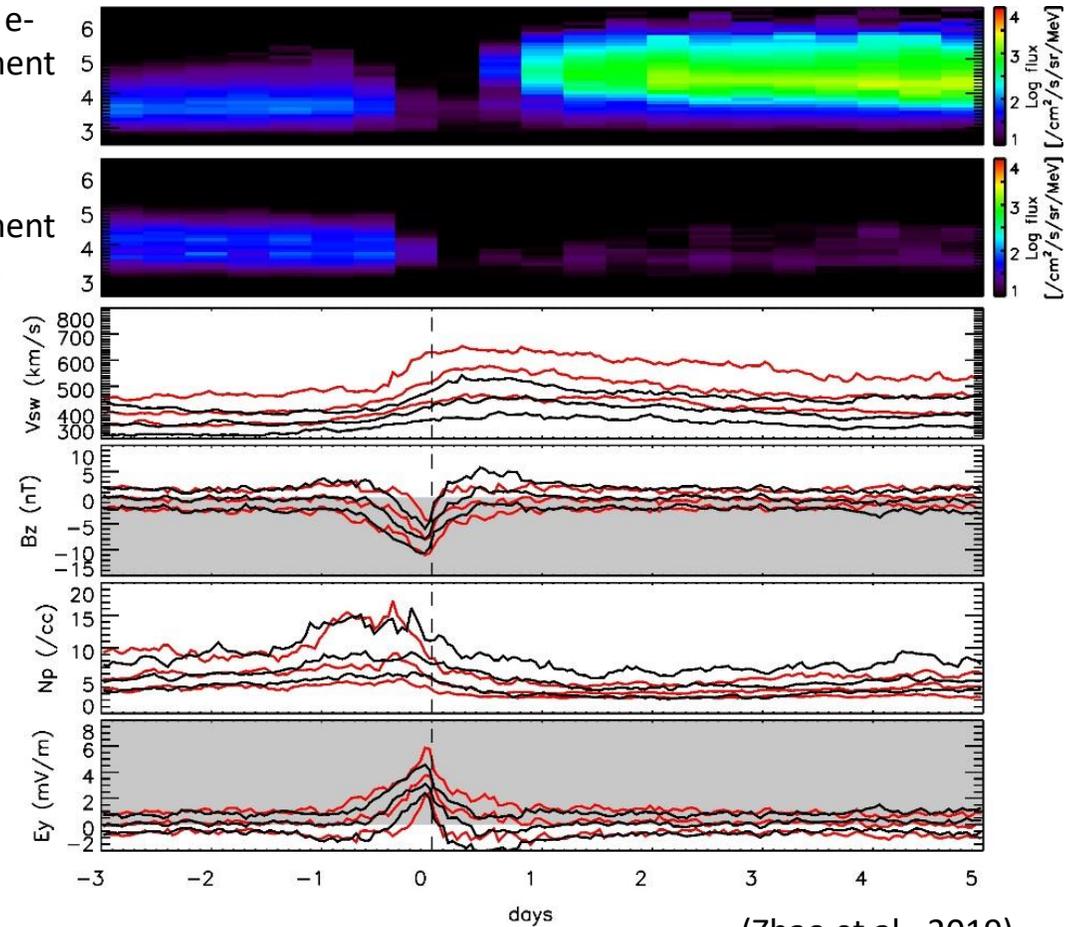
# Effects of Solar Wind on Radiation Belt Electrons



(Baker et al., 2019)

5.2 MeV e-  
enhancement  
events

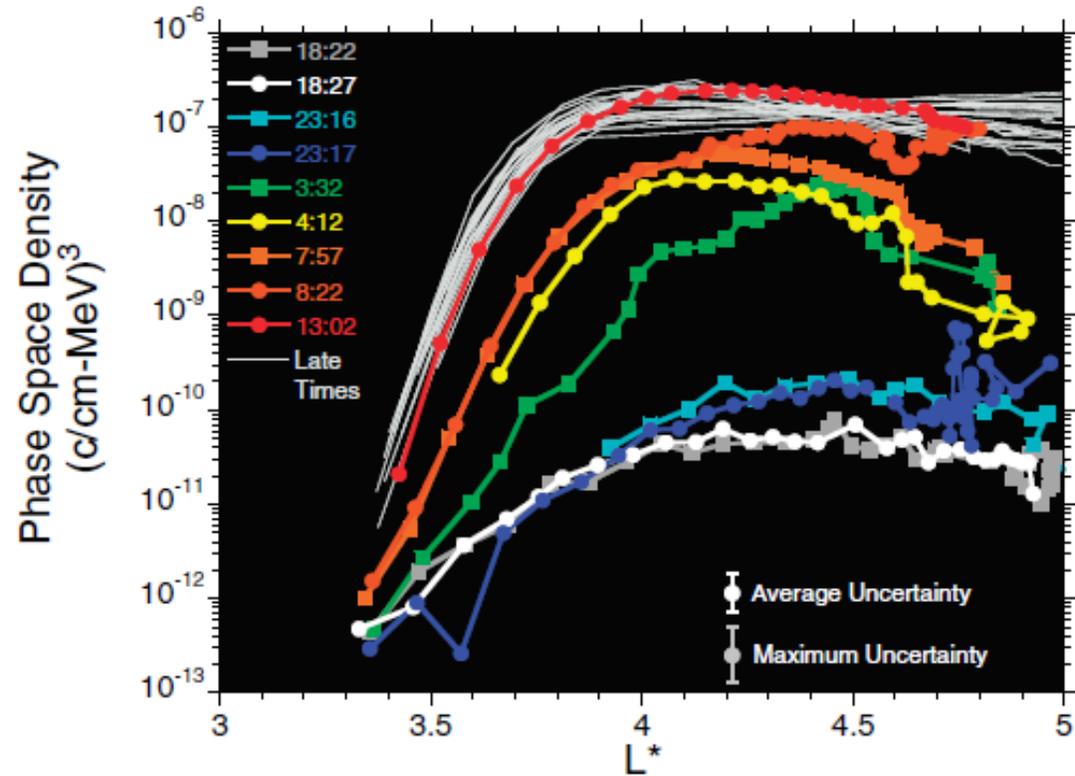
Non-  
enhancement  
events



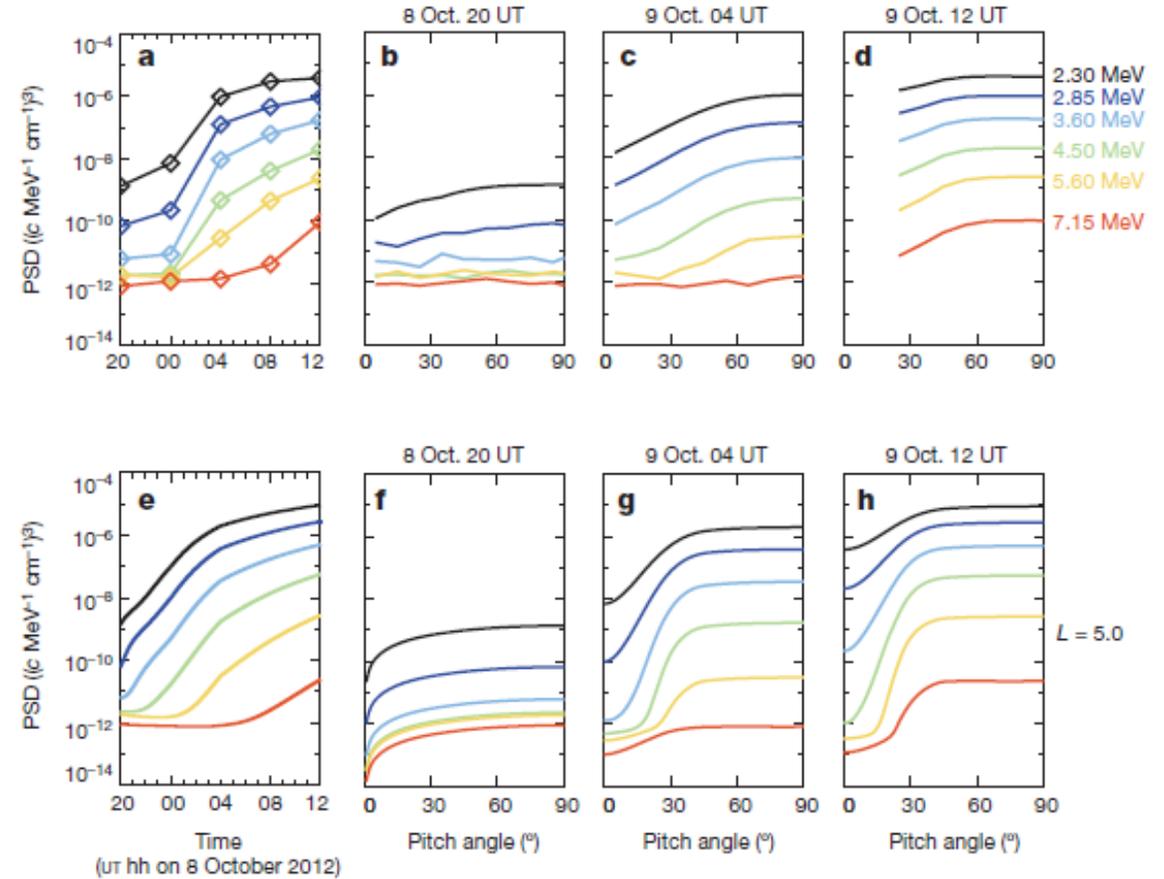
(Zhao et al., 2019)

- Multi-MeV electrons present clear energy-dependent behaviors;
- Solar wind speed is shown to be the most influential solar wind parameter causing multi-MeV electron flux enhancements.

# Acceleration Mechanism of Radiation Belt Electrons



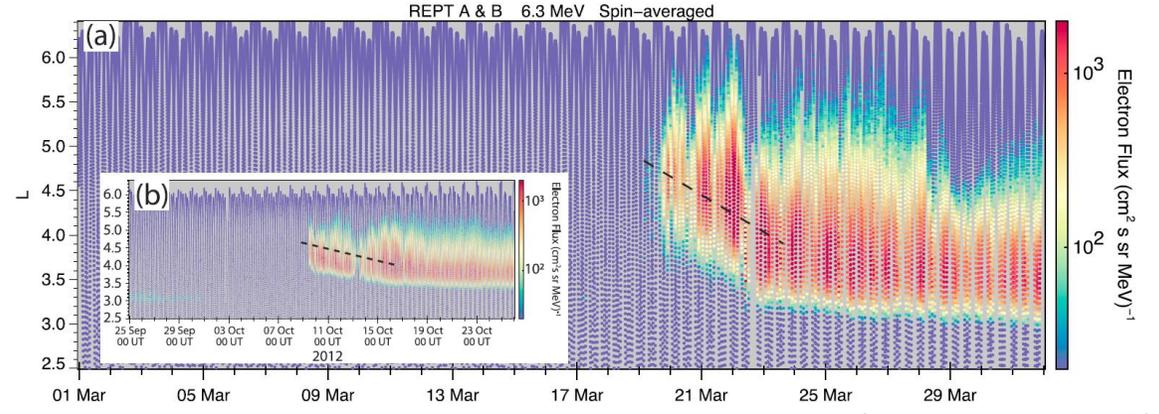
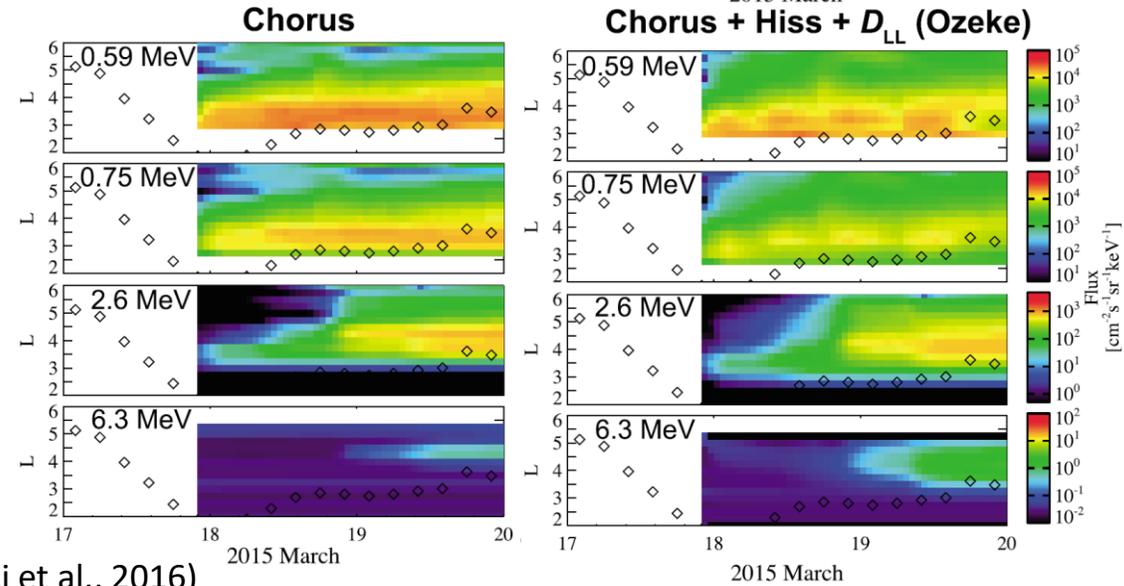
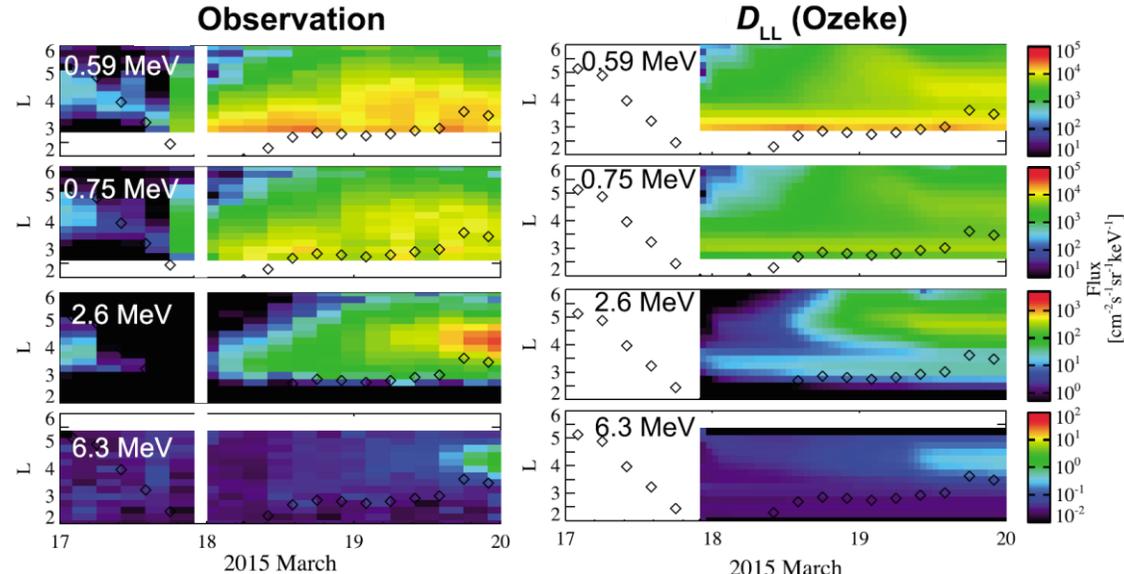
(Reeves et al., 2013)



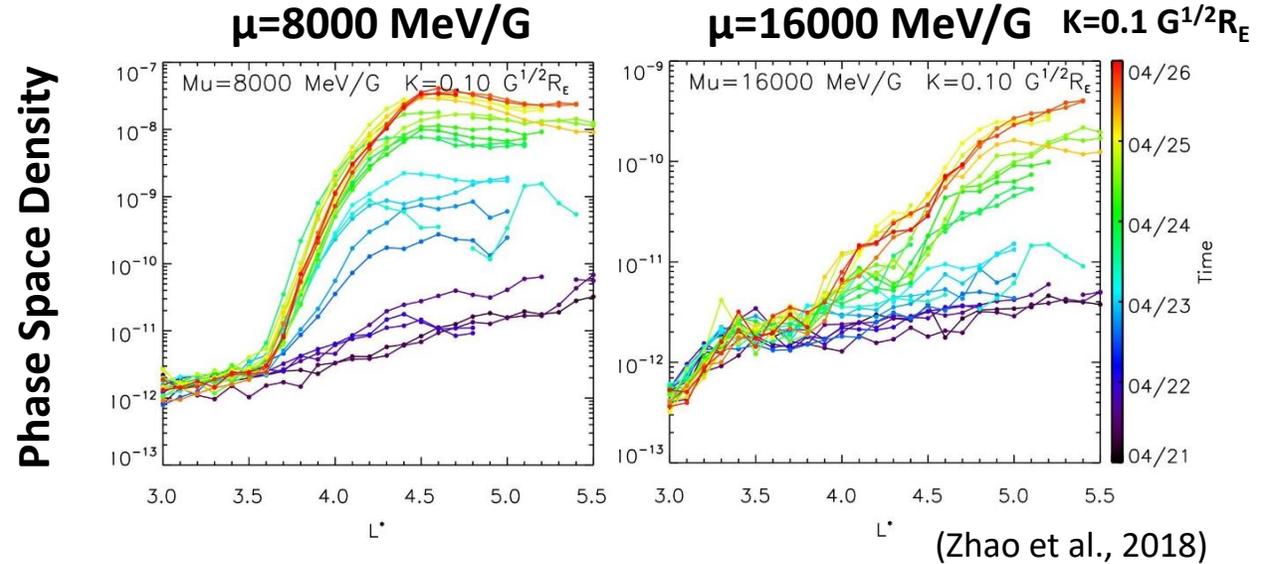
(Thorne et al., 2013)

- Local acceleration caused by whistler mode chorus waves is found to be the main acceleration mechanism for  $\sim 2 - 7$  MeV electrons in the outer radiation belt during the storm of Oct 2012.

# Acceleration Mechanism of Radiation Belt Electrons



(Jaynes et al., 2018)

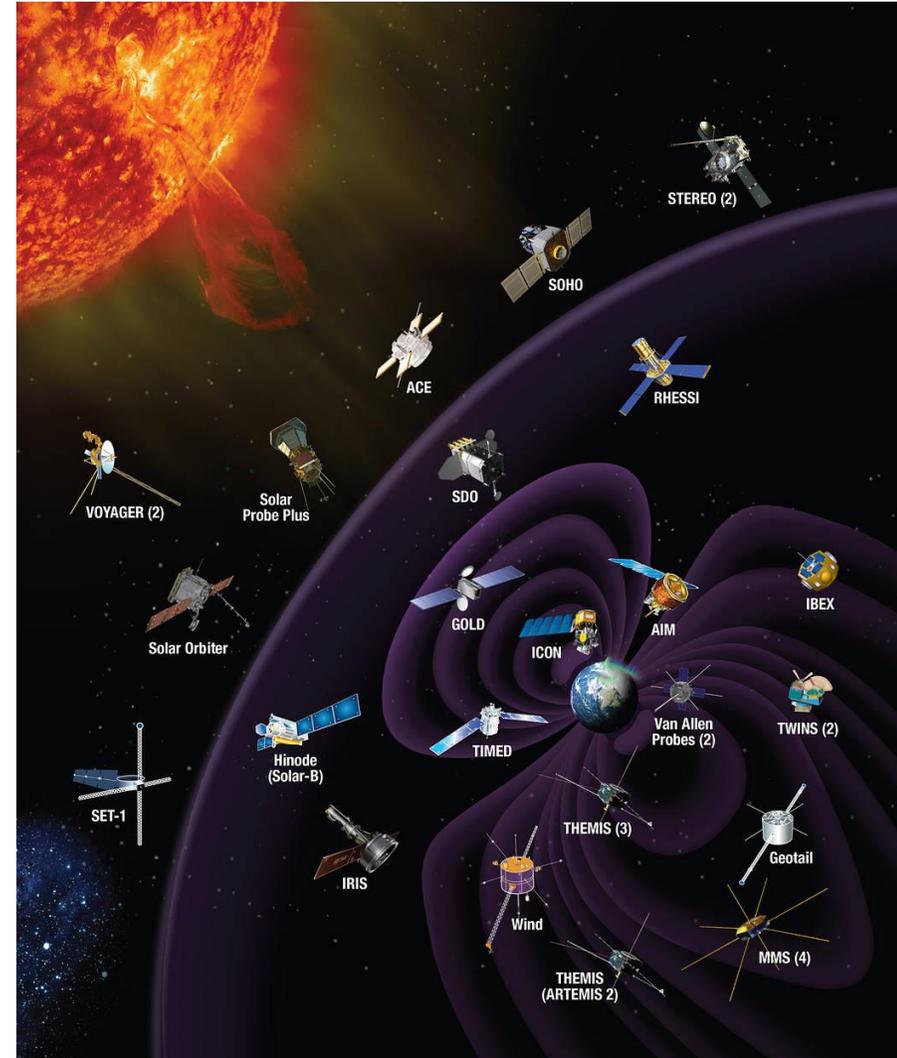


(Zhao et al., 2018)

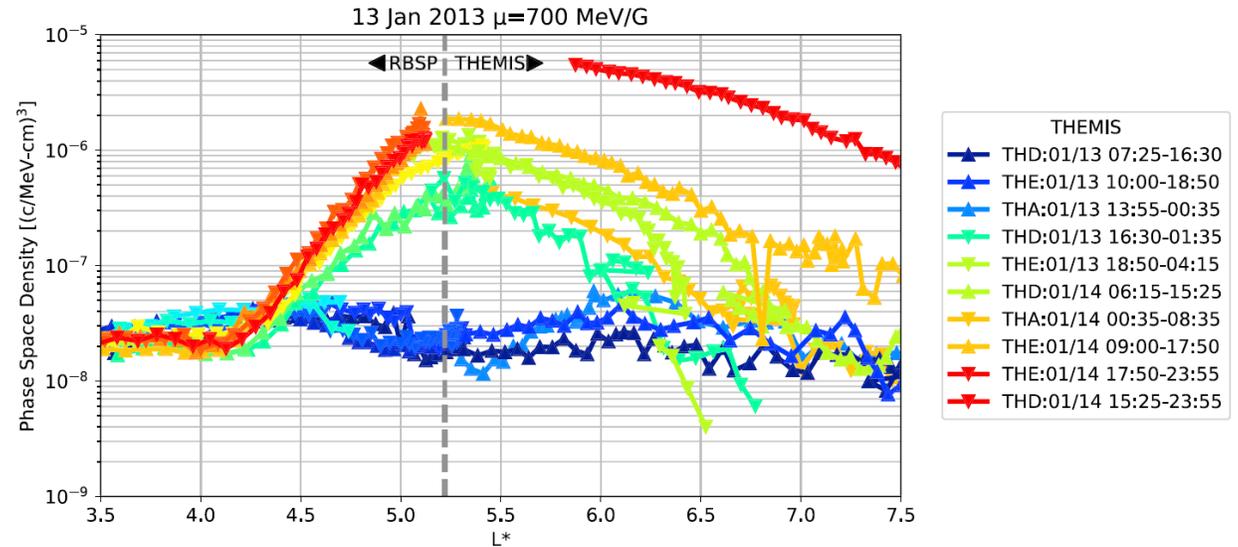
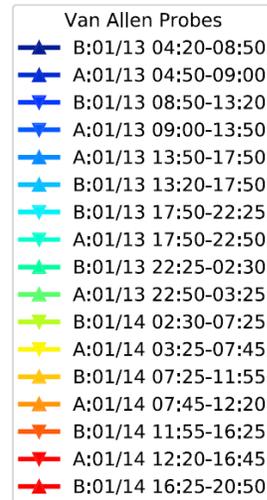
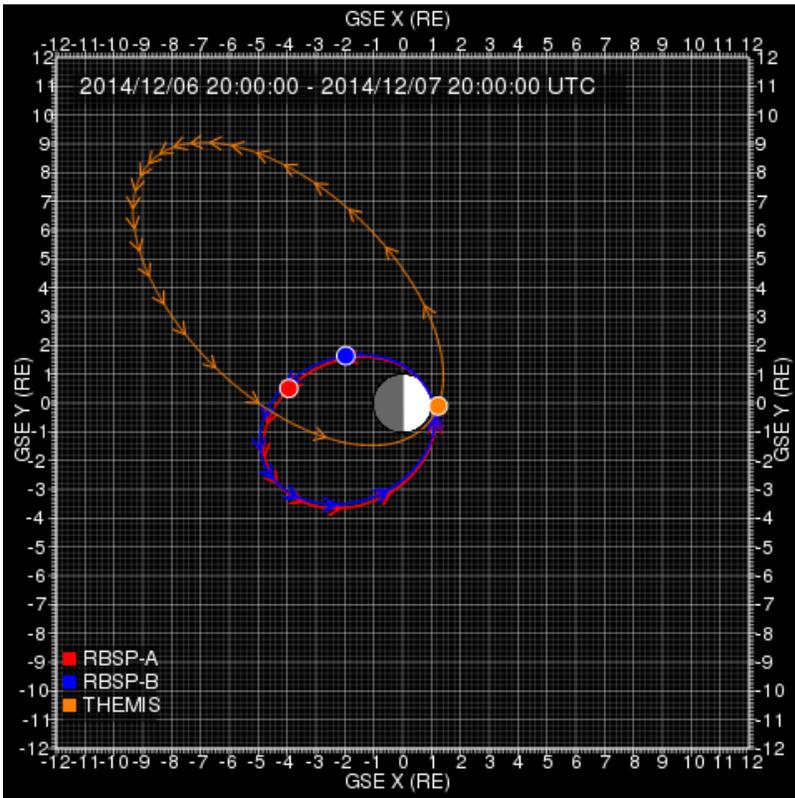
- Inward radial diffusion is also a very important acceleration mechanism for radiation belt electrons.

# System Understanding of Radiation Belt Particle Dynamics through Multi-Spacecraft and Ground-Based Observations and Modeling

- Single-point measurements have limitations in revealing underlying physical mechanisms on the radiation belt particles due to **spatial/temporal ambiguities** and **limited coverage**;
- Through **coordinated measurements** from multi-spacecraft and ground-based observations, combining with theoretical and modeling efforts, advanced understandings of radiation belt particle dynamics on both local and global scales can be gained.



# Multi-Satellite Observations Provide Comprehensive Understanding of Radiation Belt Electron Acceleration Mechanism

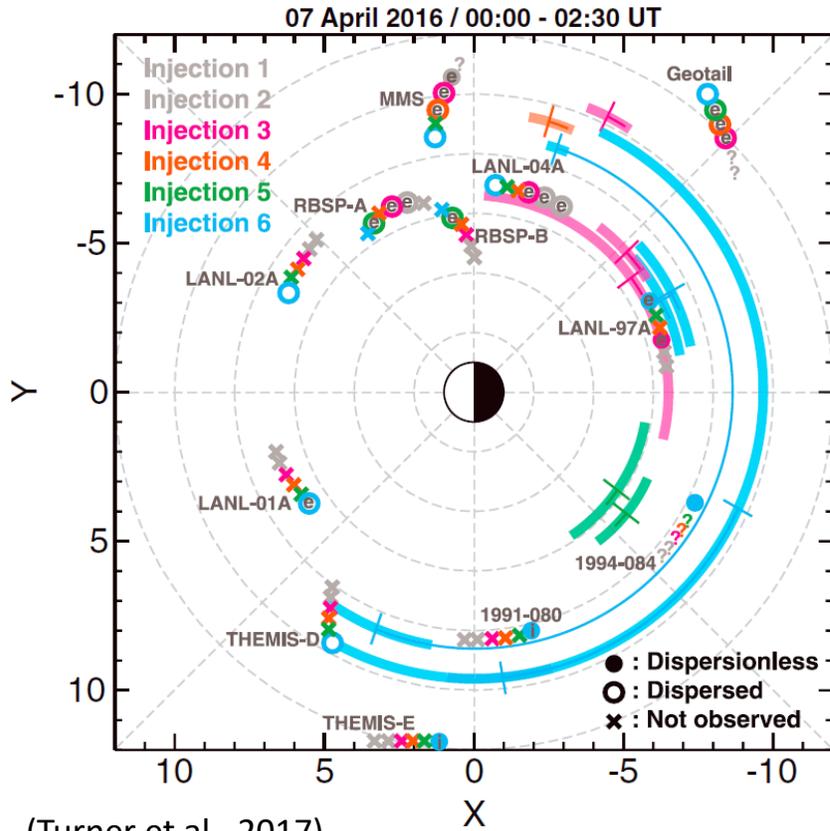


Event type	Van Allen Probes only	THEMIS & Van Allen Probes
Local acceleration dominant	24 (22)	70 (38)
Other	56 (20)	10 (4)
Total	80 (42)	80 (42)

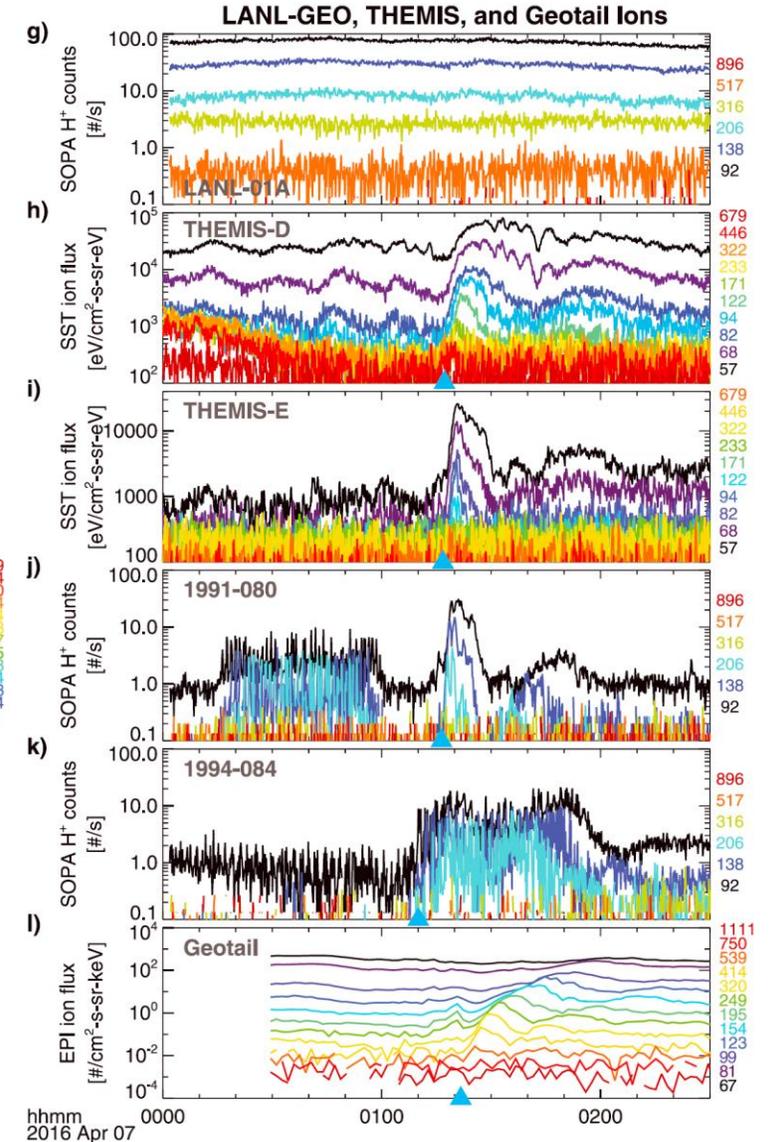
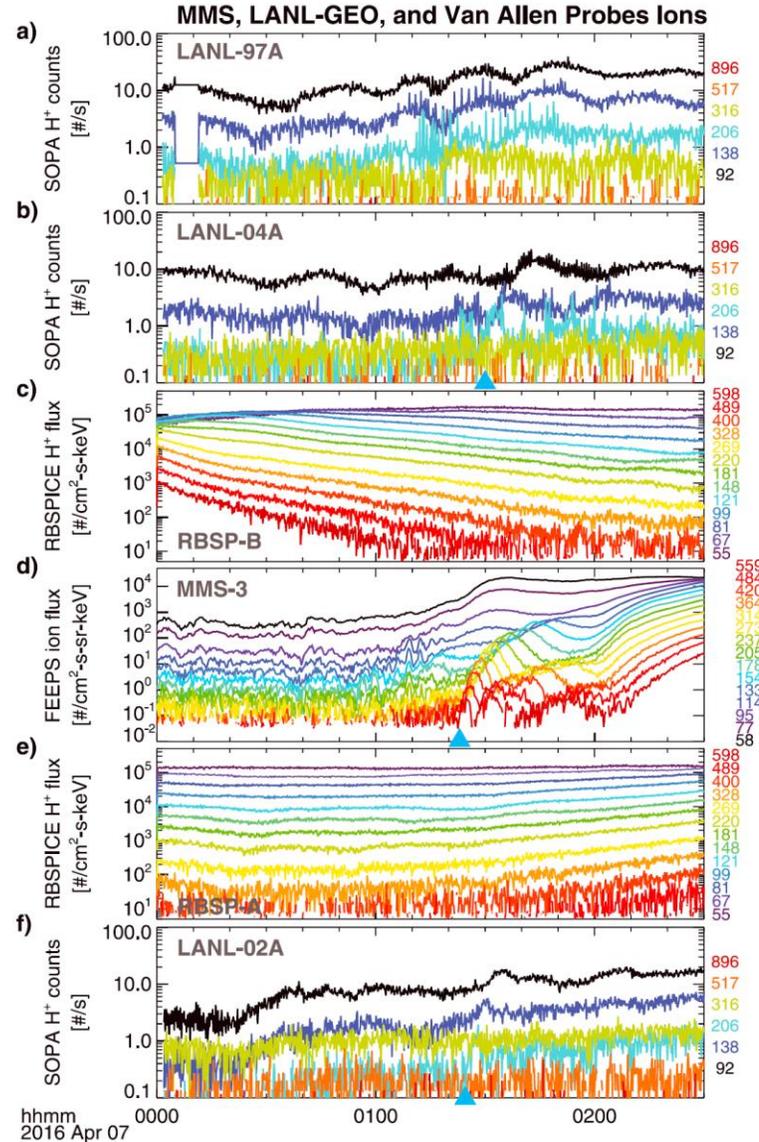
(Boyd et al., 2018)

- Using Van Allen Probes data only ( $L^*$  up to  $\sim 5.5$ ),  $\sim 30\%$  radiation belt electron enhancements show clear local acceleration dominant feature;
- Combining data from Van Allen Probes and THEMIS, most enhancements are found to be dominated by local acceleration.

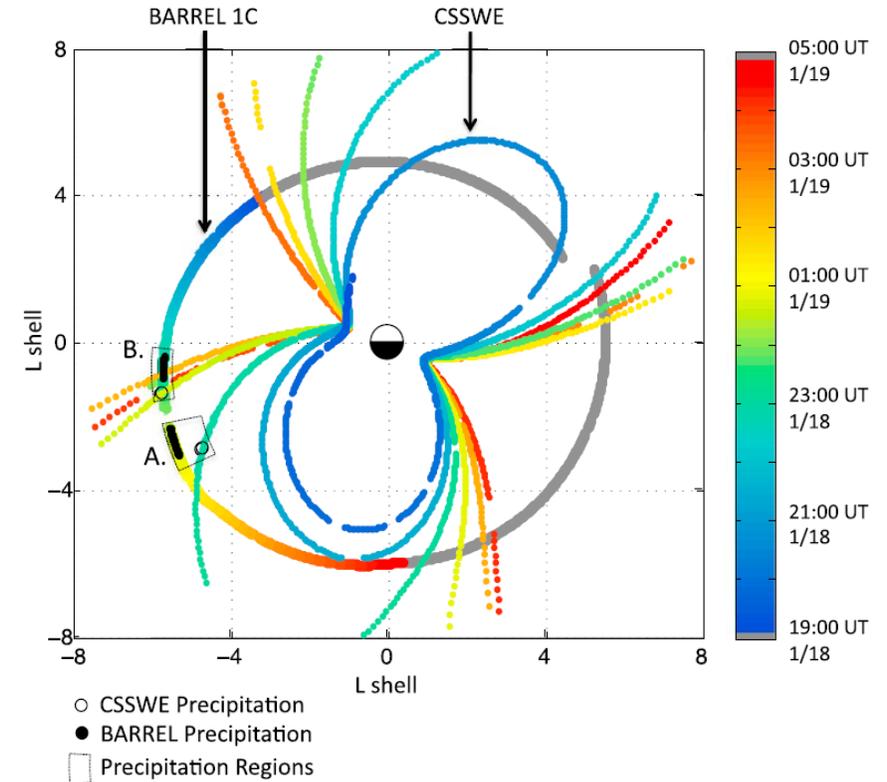
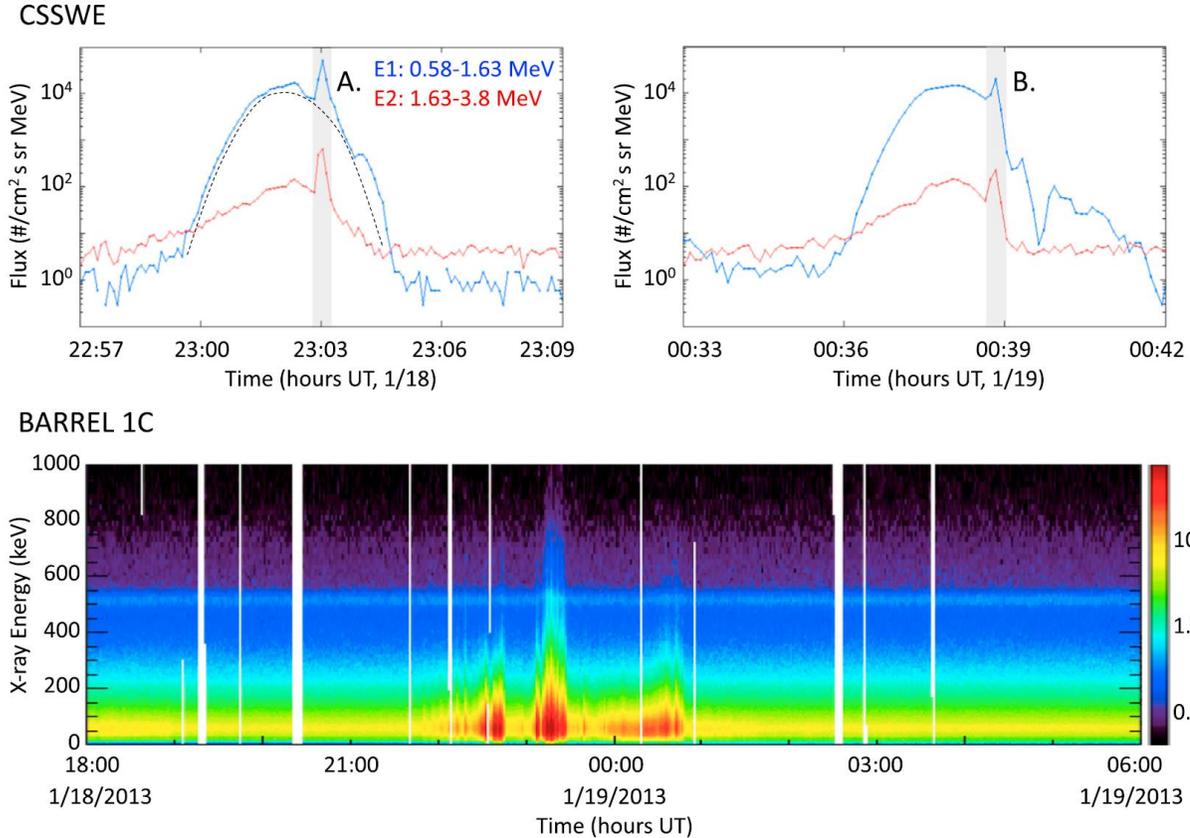
# Multipoint Measurements Provide Global Pictures of Substorm Injections



- Multipoint measurements provide comprehensive information of spatial and temporal evolution of substorm injections.



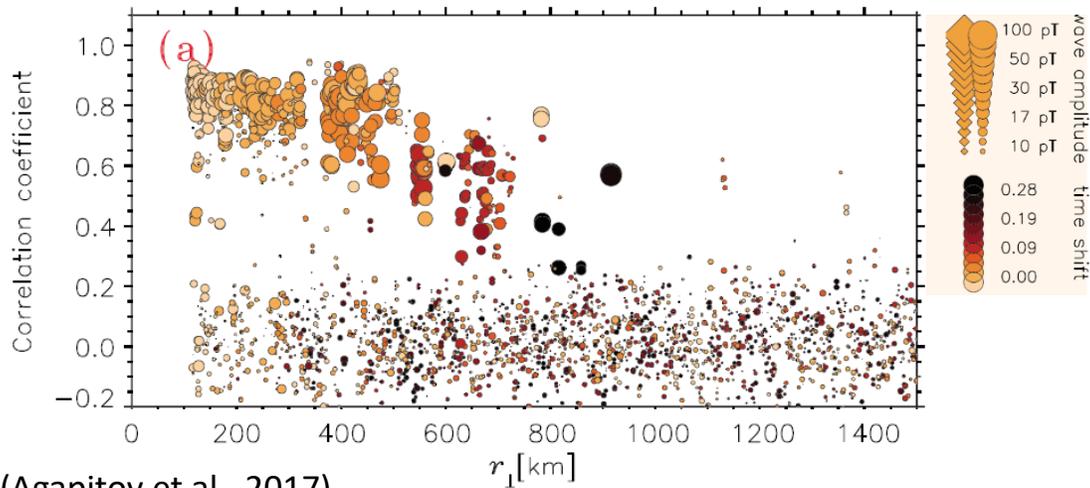
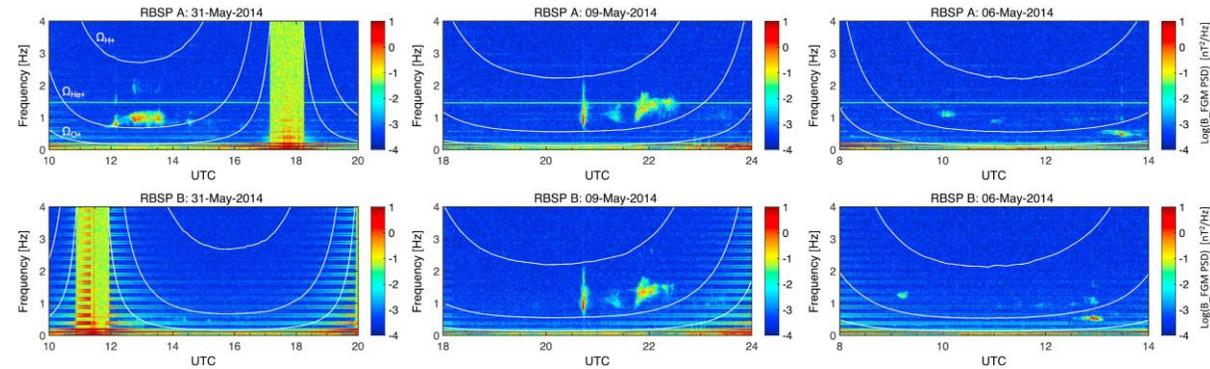
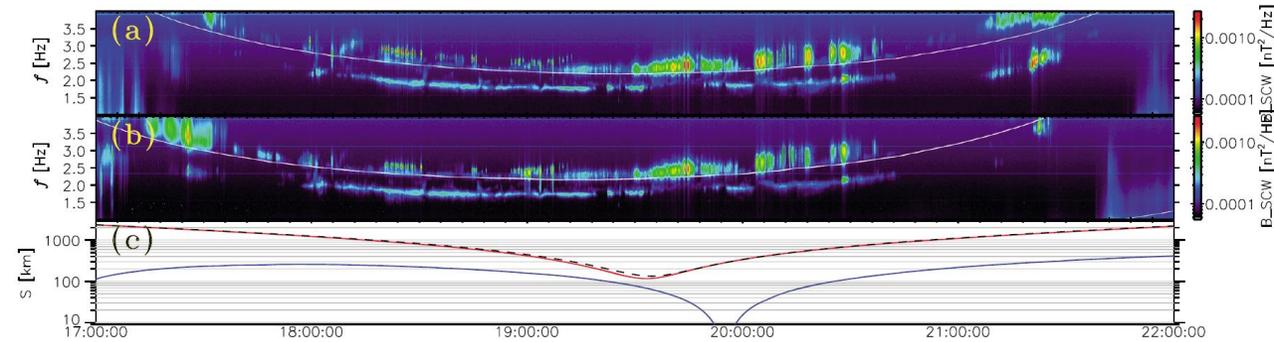
# Multipoint Measurements Provide Important Information of Electron Precipitation



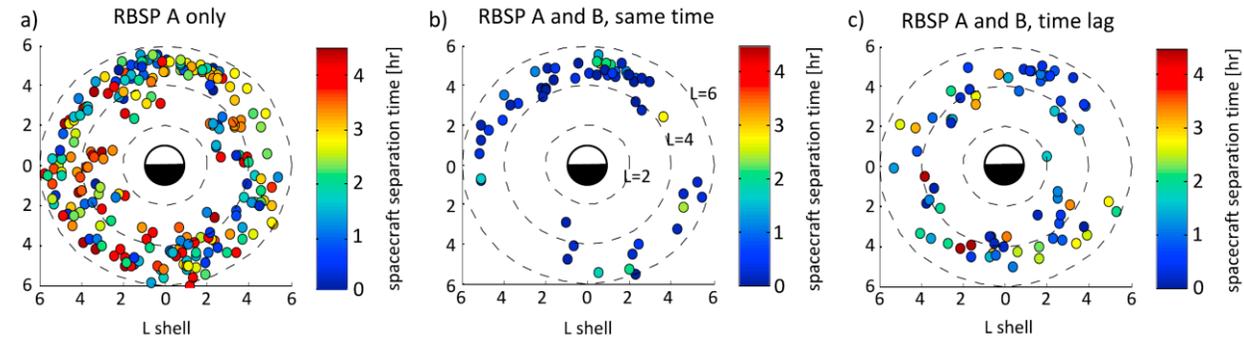
(Blum et al., 2013)

- Both CSSWE and BARREL observed the same precipitation band structures of radiation belt electrons;
- Using data from CSSWE and BARREL, the precipitation regions can be better confined in L and MLT and the electron precipitation can be better quantified.

# Multipoint Measurements Provide Better Understanding of Radiation Belt Wave Properties



(Agapitov et al., 2017)



(Blum et al., 2017)

- Using conjunctive measurements from the two Van Allen Probes, the spatial and temporal scales of chorus waves and EMIC waves are revealed.

# Summary

- Radiation belt particles exhibit significant variations under the effects of various source and loss mechanisms;
- Recent advances have been achieved on understanding the radiation belt particle dynamics thanks to the unprecedented measurements of the Van Allen Probes;
- Coordinated multi-spacecraft and ground-based measurements, combining with theoretical and modeling efforts, will further deepen our understanding of radiation belt particle dynamics.

# Thank you!

# System Understanding of Radiation Belt Particle Dynamics through Multi-Spacecraft and Ground Based Observations and Modeling

- The broad goal of this FG is to deepen understanding of radiation belt particle dynamics through coordinated multi-mission measurements, combining with theoretical and modeling efforts.
- Specific topics include:
  - Improve the understanding of physical mechanisms related to radiation belt electron acceleration and loss on short timescales (minutes to hours);
  - Quantify the radiation belt electron precipitation into the atmosphere and understand the related physical mechanisms;
  - Improve the understanding of the properties and spatiotemporal distribution of waves and their effects on the radiation belt particles;
  - Advance the understanding of inner belt and slot region particle dynamics.

